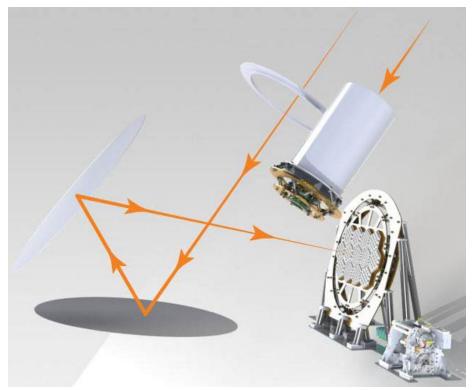
LiteBIRD Focal Plane Overview

Aritoki Suzuki University of California, Berkeley 2015 December 16th

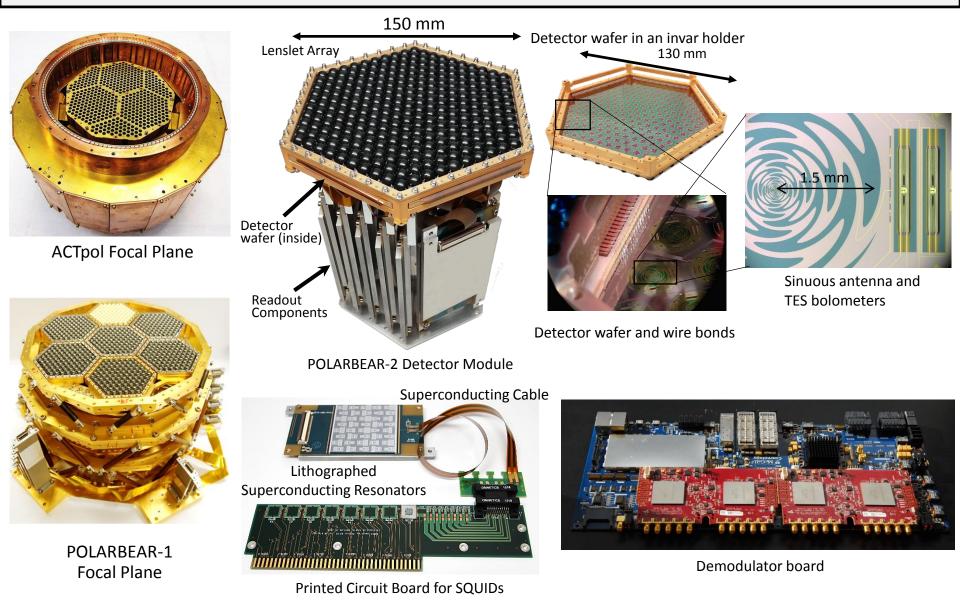
US LiteBIRD Team





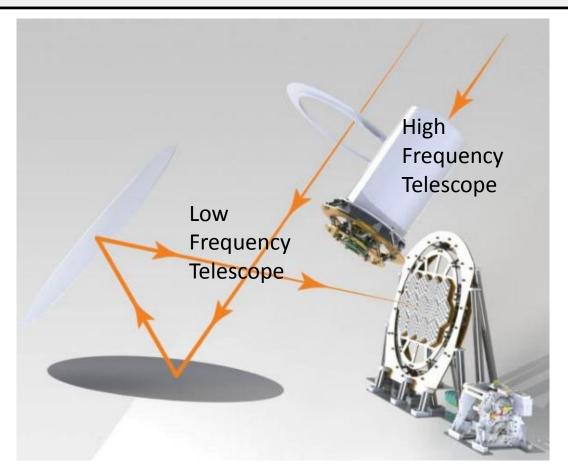
- US LiteBIRD team proposed to NASA, Mission of Opportunity to contribute
 - Analysis (Monday)
 - Sub-Kelvin cryo-coolers (Tuesday)
 - Detectors
 - Readout electronics

Heritage



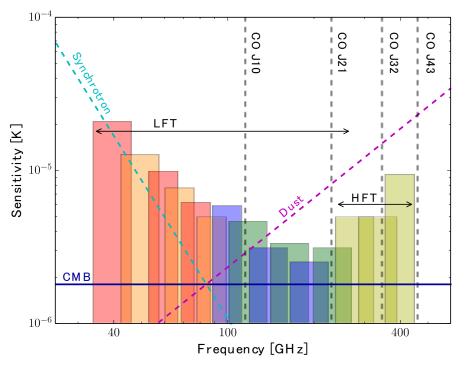
Heritage from ground and balloon based CMB experiments

Low/High Frequency Telescope



- Low Frequency Telescope \rightarrow 40 GHz ~ 235 GHz
- High Frequency Telescope \rightarrow 280 GHz ~ 402 GHz

Frequency Coverage



• 15 bands for foreground monitoring

• Multiple components dust, Synchrotron with $\Delta\beta$

• Low Frequency Telescope (LFT)

- Low Frequency Detector (LF) 40~88.5 GHz
- Mid Frequency Detector (MF) 100~234.9 GHz
- High Frequency Telescope (HFT)
 - High Frequency Detector (HF) 280~402.1 GHz

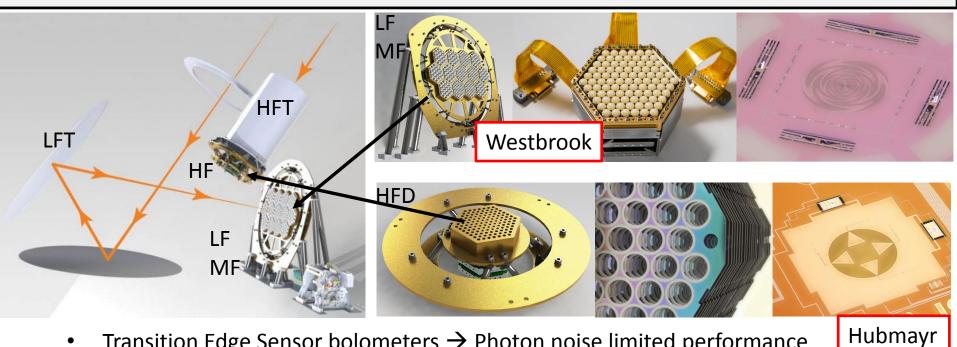
2 triplexer pixels

2 triplexer pixels

3 single band pixels

12/16/2015

Feeds and Detectors



- Transition Edge Sensor bolometers \rightarrow Photon noise limited performance
- Heritage from CMB polarization experiments
- **LF and MF**: Lenslet coupled broadband antenna array
 - POLARBEAR-2, Simons Array, SPT-3G
 - Talk by Ben Westbrook
- **HF**: Horn + OMT array
 - SPT-Pol, ACT-Pol
 - Talk by Hannes Hubmayr

Fabrication Facility and Team

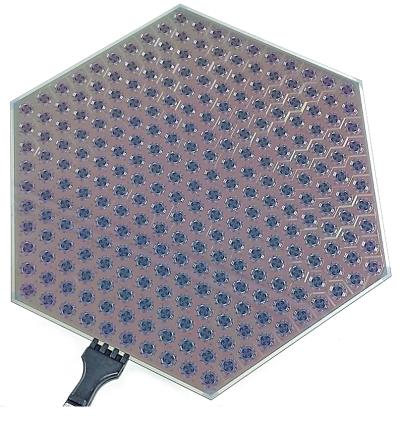


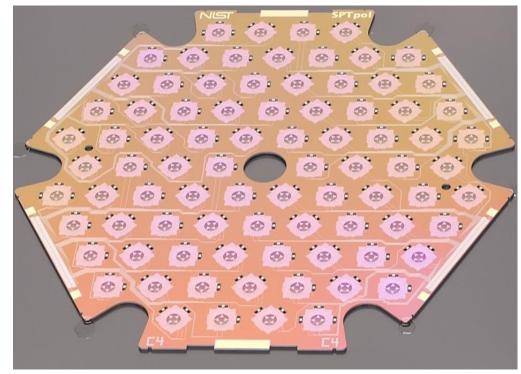
UC Berkeley Marvell Nanofabrication Facility

NIST Boulder Micro Fabrication Facility

- UC Berkeley Marvell Nanofabrication Facility
 - 15,000 square feet class 100 clean room
 - Two dedicated fabrication engineers and one dedicated tool engineer at UC Berkeley
 - Stanford fabrication engineer to fabricate at UC Berkeley
- NIST Boulder Micro Fabrication Facility
 - 18,000 square feet class 100 clean room
 - Two dedicated fabrication engineers for this project

Fabrication Experience





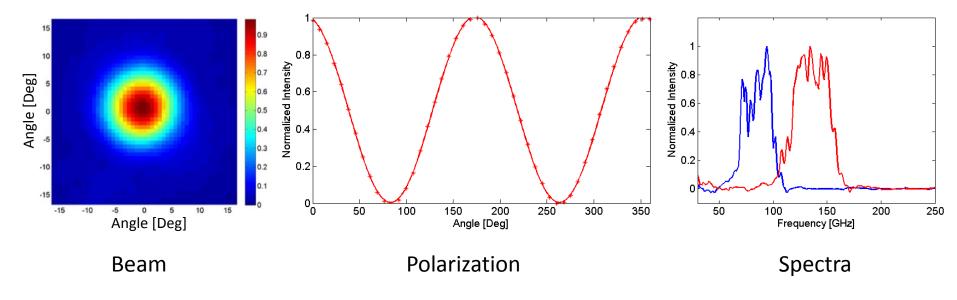
POLARBEAR-2 Detector Array fabricated at UC Berkeley SPTpol Detector Array fabricated at NIST

History of supplying detector arrays to many CMB experiments

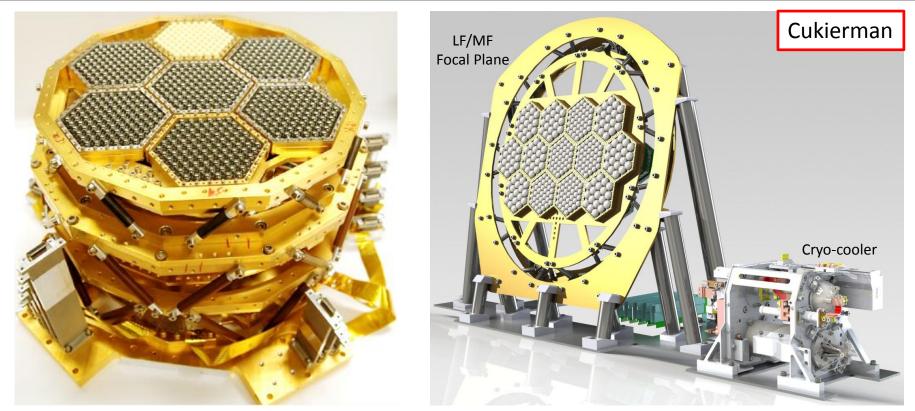
Detector Performance

 NIST fabricated TES bolometer with: Noise Equivalent Power (NEP)=4×10–18W/ V Hz at signal frequencies below 1 Hz with thermal conductance of g = 10 pW/K Meets LiteBIRD spec

Optical performance of LiteBIRD spec sinuous antenna pixel from UC Berkeley



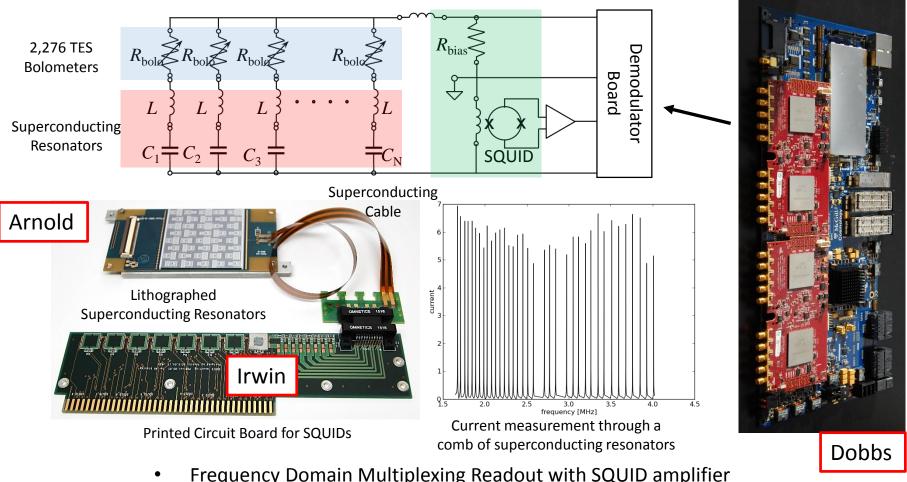
Focal Plane Design



POLARBEAR-1 Focal Plane Tower

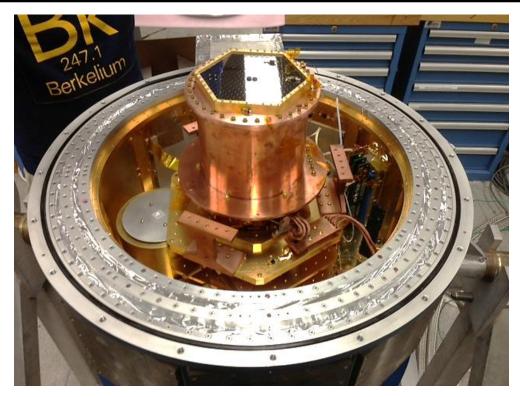
- 100 mK base temperature
- Cryo-cooler with multiple temperature stages
- Design to balance thermal load and structural strength
- Current design is based on deployed CMB experiments
- Talk by Ari Cukierman

Readout Electronics



- Frequency Domain Multiplexing Readout with SQUID amplifier
- Ground and balloon CMB experiments heritage
- Talk by Kam Arnold Cold electronics
- Talk by Matt Dobbs Warm electronics
- Talk by Kent Irwin SQUID in space

Test Facility











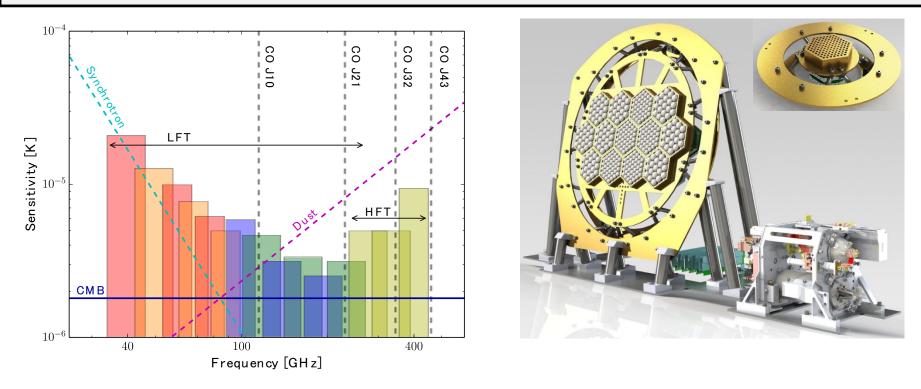


- Dedicated test cryostats at
 - Berkeley, Stanford, NIST, U of Colorado, UW-Madision

Subset of Ongoing Studies

- Robustness of frequency coverage
 → Foreground joint study group
- Radiation effects on focal plane components
 → Radiated sample at HIMAC
- Thermal and structural test
 → Vibration test at Space Science Lab, Berkeley

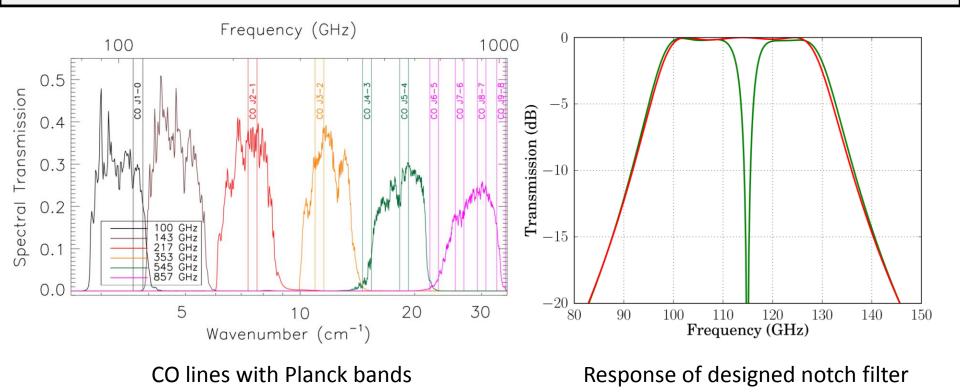
Conclusion



- LiteBIRD focal plane components have heritage from ground and balloon CMB polarization experiments
- 15 frequency bands for foreground monitoring
- Experienced experimental CMB team

Back Up

CO – Notch Filter



• Notch filter to take CO line out