Concluding Remark

Eiichiro Komatsu (Max-Planck-Institut für Astrophysik) "Cosmic Acceleration" Kick-off Meeting September 21, 2014

Great start

- ~100 participants in the middle of holidays, in such a short notice
- Remarkable enthusiasm in the community!

Just in case you have not realised already...

- Observational cosmology in Japan is <u>exploding</u>
- This is unprecedented: we have never seen the rapid development of this magnitude before
 - I left this country, because there was no observational cosmology in Japan back in 1999
- HUGE change now!

Amazing Experiments

- **B01**: CMB [*Simons Array* and *LiteBIRD*]
- **B02**: Imaging LSS [*HSC*]
- **B03**: Spectroscopic LSS [*PFS* (and *eBOSS*)]
- **B04**: Direct measurement of acceleration [*TMT*]
- All of these are front-runners in the world
 - It's like ... wow. They are potential game-changers

Amazing Experiments

Do participate!

- BO You may contribute to changing the history of cosmology and astrophysics
- B0

B0

B0

- Do not miss this opportunity!
- Apply for positions; apply for grants
- It's like ... wow. They are potential game-changers

Amazing Theory Groups

- **A01**: Inflation
- A02: Dark Matter
- A03: Dark Energy
- **C01**: Fundamental Physics
- They have been front-runners already. Just let them do whatever they want

Amazing Theory In my humble opinion:

- A0 Theorists should do whatever they want,
 A0 so they do not have to aim at the following specifically, but:
- A0
- C0

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Perhaps one of the useful outcomes from theoretical studies would be to propose new observables/ measurements





Personally:

Tell us what to measure! • A01: Inflat **One useful question** • A02: Dark 00 to ask: oute Has [your] theoretical A03: Dark study changed the Per les way we take/analyse **C01**: Fund Bu JID the data? be the other way around They have ust let them do whatever they want

3 pillars of science (theory)

| | | [A01] Inflation Sasaki (Kyoto) | [A02] fluent. & struct. Takahashi (Tohoku) | [A03] Dark Energy Sugiyama (Nagoya) |
|---|---|--|---|--|
| | [B01] CMB polariz. Hazumi (KEK) | ζ, r, n_s direct evidence | CMB lensing isocurv. $m_{ u}, N_{ u}$ | cosmo. params CMB lensing |
| | [B02] Subaru galaxy imaging Miyazaki(NAOJ) | Lensing $\rightarrow b(k)$ $\rightarrow P_{\text{primod}}(k)$ | weak lensing $m_{ u}$ non-std. DM | weak lensing SNe, γ |
| _ | [B03] galaxy spectroscopy Takada(KIPMU) | primord. NG $\Omega_K, n_s, lpha_s$ | isocurv. DM in dSph gals. $P(k), m_ u$ | BAO, RSD $\Omega_{ m de}(z), \gamma$ |
| | [B04] TMT Usuda (NAOJ) | QED coupling (α) space time var. | Lyman-α forests IGM | direct detection of acceleration |

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| | [B02] Subaru galaxy imaging Miyazaki(NAOJ) | Lensi $\rightarrow P_{\mathfrak{p}}$ | Synergy between the groups is the | | weak lensing SNe, γ |
| | [B03] galaxy spectroscopy Takada(KIPMU) | prim Ω_K | most important success criterion! | | BAO, RSD $\Omega_{ m de}(z), \gamma$ |
| | [B04] TMT Usuda (NAOJ) | QED coupling (α) space time var. | | Lyman-α forests IGM | direct detection of acceleration |

C01: ultimate, theory Ooguri(Caltech)

//X00: organization //Murayama (IPMU)

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lensing

SNe!

BAQRSD

 $\Omega_{de}(z), \gamma$

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IGM

D01: ultimate analysis Komatsu(MPA)

Your home work: Think about grant proposals!

- This project comes with opportunities for 14 participating proposals in JFY2015!
- E01: Theoretical and/or numerical study related to the evolution and structure of the Universe ~9 <¥1M
- E02: Experimental and observational study related to the evolution and structure of the Universe ~2 <¥2M
- E03: Study that can bridge between theoretical, numerical and experimental/observational studies related to the evolution and structure of the Universe ~2 < ¥4M