

B04

Direct Detection of Cosmic Acceleration by TMT & High Dispersion Spectrograph

次世代超大型光学赤外線望遠鏡TMTと
高分散分光器による宇宙の加速膨張の直接検証

TMT & Overview

Tomo Usuda (NAOJ / TMT-J Director)

B04 Members

◆ PI (研究代表者)

Tomo Usuda (NAOJ TMT-J)

◆ CoI (研究分担者)

Hajime Inaba (AIST) **Optical Laser Frequency Comb**

Karoru Minoshima (UEC) **Optical Laser Frequency Comb**

Takeshi Chiba (Nihon U) Direct detection of acceleration

Toru Misawa (Shinshu U) Lyman- α forests IGM, Obs.

◆ Collaborator (連携研究者)

Feng-Lei Hong (Yokohama Nat. U) **Optical Laser Frequency Comb**

Nobunari Kashikawa (NAOJ) TMT instruments, IGM,

Wako Aoki (NAOJ) TMT instruments, Obs.

Atsushi Onae (AIST) **Optical Fiber Laser Frequency Comb**

Sho Ookubo (AIST) **Optical Fiber Laser Frequency Comb**

◆ Others

Malte Shramm (AIST) **Optical Fiber Laser Frequency Comb**, Obs.,

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Hajime Inaba (AIST) Optical Laser Frequency Comb

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◆ Optical Frequency combs and their applications

Feng-Lei Hong (Yokohama Nat. U) Optical Laser Frequency Comb

Nobunari Kashikawa (NAOJ) TMT instruments, IGM,

Wako Aoki (NAOJ) TMT instruments, Obs.

Atsushi Onae (AIST) Optical Fiber Laser Frequency Comb

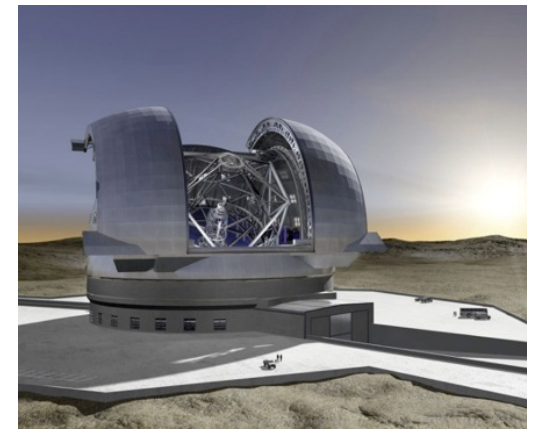
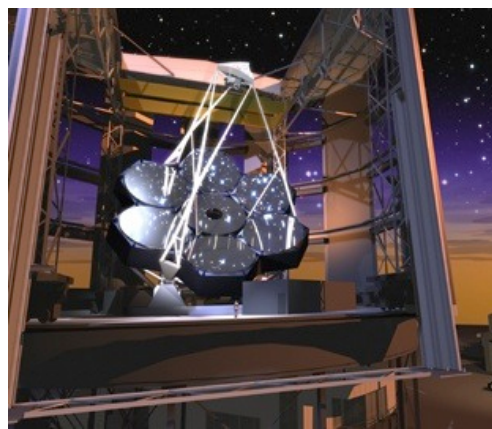
Emerging age of precision spectroscopy in Astronomy using Laser Frequency Combs



Malte Shramm (AIST) Optical Fiber Laser Frequency Comb, Obs.,

3 ELT projects in the world

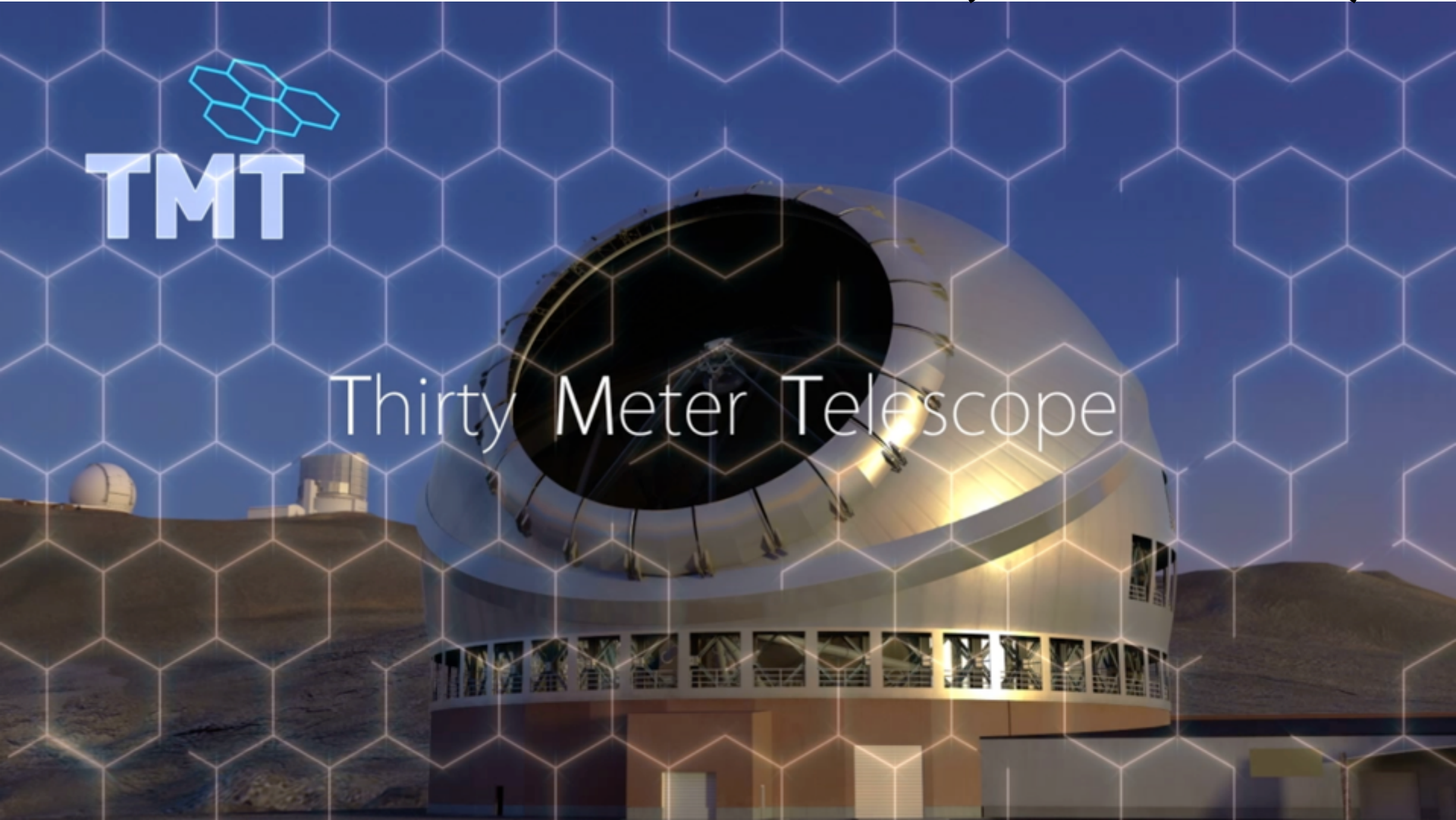
Project	TMT (Thirty Meter Telescope)	GMT (Giant Magellan Telescope)	E-ELT (European-Extremely Large Telescope)
M1	30m	22m (8.4m x 7)	39m
Site	Hawaii, Mauna Kea (4,012m), North	Chile, Las Campanas (2,550m), South	Chile, Cerro Armazones (3,060m)
Budget	1,580M USD	1,100M USD (estimate)	2,000M USD (estimate)
Construction	2014~2024	2013~2021 (with 4 of 7 mirrors)	2014~2025
Member	Japan, USA (Caltech, UC, (AURA)), Canada, China, India	USA (Carnegie, Arizona, etc.), Australia, S.Korea, SaoPaolo)	ESO (including Brasil)



TMT on Mauna Kea (CG: 140 sec)

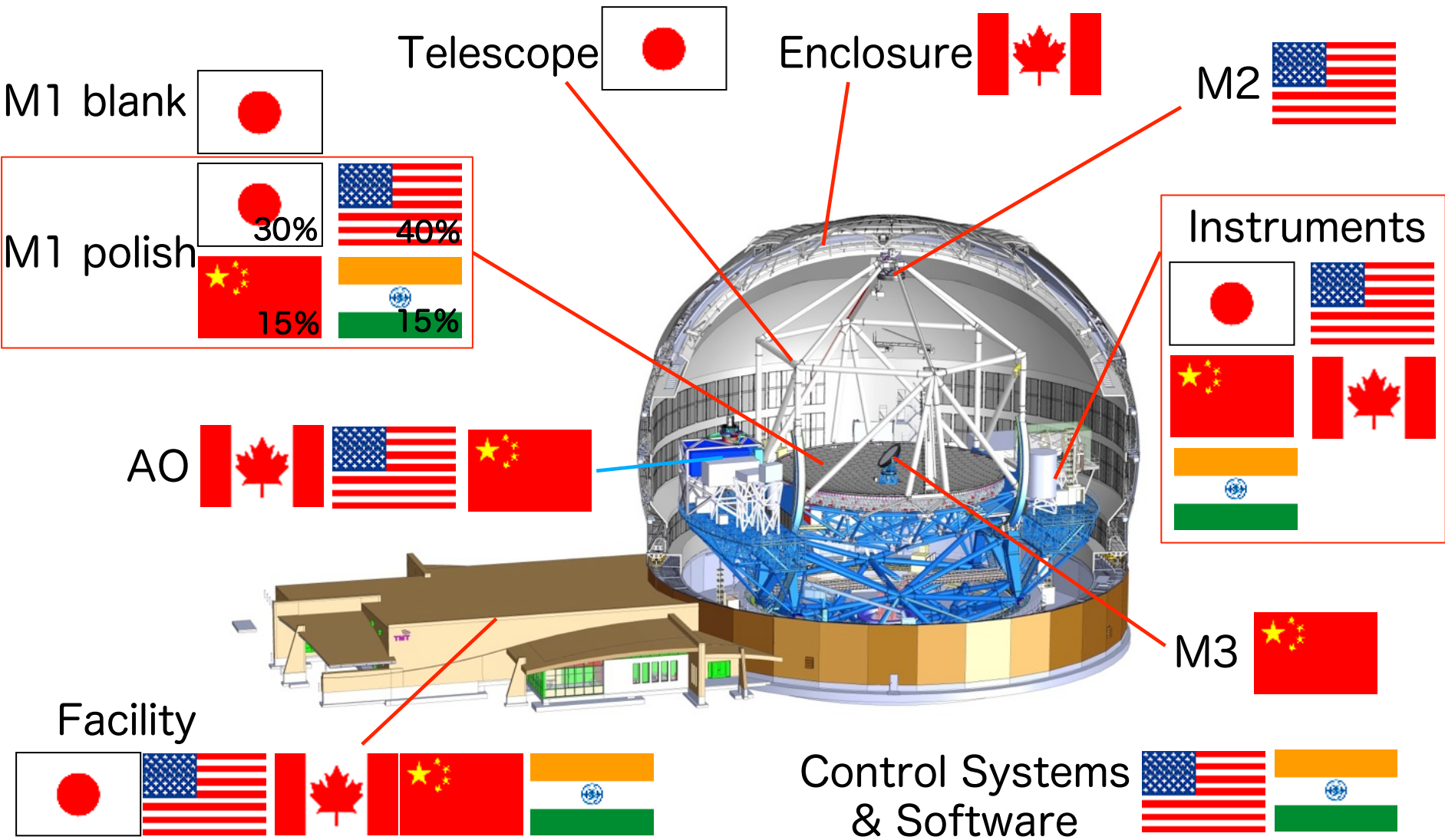


Thirty Meter Telescope



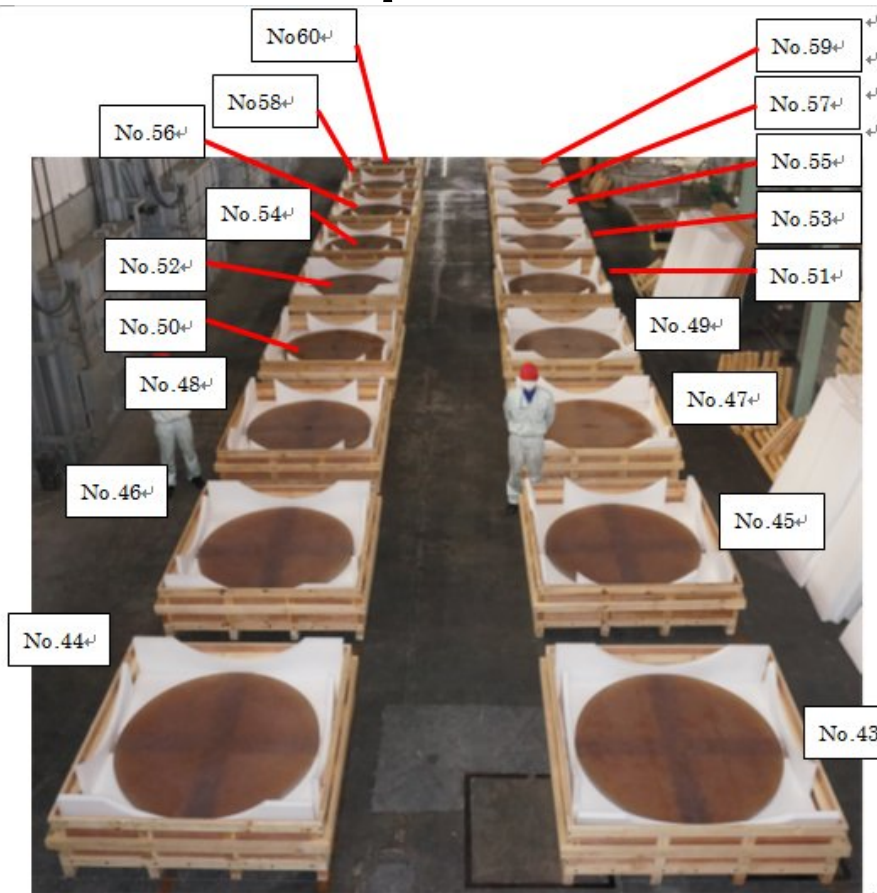
You can download the movie
<http://tmt.nao.ac.jp/gallery/index.html>

Members' contributions to TMT



Providing M1 segment mirrors

- ◆ Prototype of M1 was made in 2013
- ◆ **99 plano-plano segment blanks by March 2015**
- ◆ Mass production is on going



TMT Science Instruments

Opt

WFOS
(Multi-Obj Spec)

no AO/GLAO

HROS
(High Spec Res)

no AO/GLAO

2nd gen.
2026~

NIR

IRIS
(IFU)

NFIRAOS

IRMS
(Multi-Obj Spec)

NFIRAOS

NIRES-Blue
(High Res J~K)

NFIRAOS

IRMOS
(Wide-field MO IFU)

MOAO

Early light
2024~

NIRES-Red
(High Res J~K)

MIR-AO

PFI
(Exoplanet finder)

ExAO

MIR

MICHI
(MIR Cam/Spec)

MIR-AO

3 pillars of science (theory)

4 approaches (expt, obs)

	[A01] Inflation Sasaki (Kyoto)	[A02] fluent. & struct. Takahashi (Tohoku)	[A03] Dark Energy Sugiyama (Nagoya)
[B01] CMB polariz. Hazumi (KEK)	ζ, r, n_s direct evidence	CMB lensing isocurv. m_ν, N_ν	cosmo. params CMB lensing
[B02] Subaru galaxy imaging Miyazaki(NAOJ)	Lensing $\rightarrow b(k)$ $\rightarrow P_{\text{primod}}(k)$	weak lensing m_ν non-std. DM	weak lensing SNe, γ
[B03] galaxy spectroscopy Takada(KIPMU)	primord. NG Ω_K, n_s, α_s	isocurv. DM in dSph gals. $P(k), m_\nu$	BAO, RSD $\Omega_{\text{de}}(z), \gamma$
[B04] TMT Usuda (NAOJ)	(2) QED coupling (α) space time variation	(3) Lyman- α forests IGM	(1) Direct detection of acceleration

important observables at each intersection

B04 Objectives

◆ 3 Science topics

(1) **Direct Detection of Cosmic Acceleration** 「宇宙の加速膨張の直接測定」 **Chiba**

(2) **Constraining the spatial & temporal variation of the fine structure constant** 「物理定数（微細構造定数、陽子・電子質量比）の時間変化に対する制限」 **Chiba**
Misawa

(3) **IGM tomography by QSOs & galaxies**
Absorption Lines 「銀河間物質の3次元構造の解明」
Misawa, Kashikawa

◆ R&D Plans for TMT High Dispersion Spectrograph

◆ Develop **Optical Laser Frequency Comb** for very accurate / stable λ Calibrations **Inaba, Minoshima, Hong, Onae, Ookubo, Shramm**

◆ Install the comb on **Subaru HDS for on-site tests** to know technical difficulties and establish the observation strategies for TMT **Aoki, Misawa, Kashikawa**

(1) Direct Detection of Cosmic Acceleration

Future Universe?

What is Dark Energy?

Universe after 10 yrs

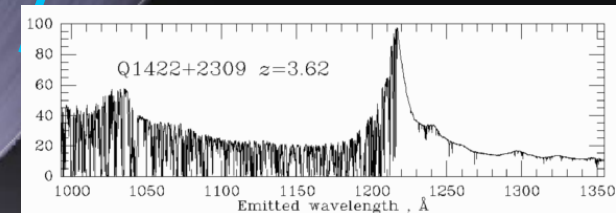
Current Universe

First direct detection of
time variance toward same
objects

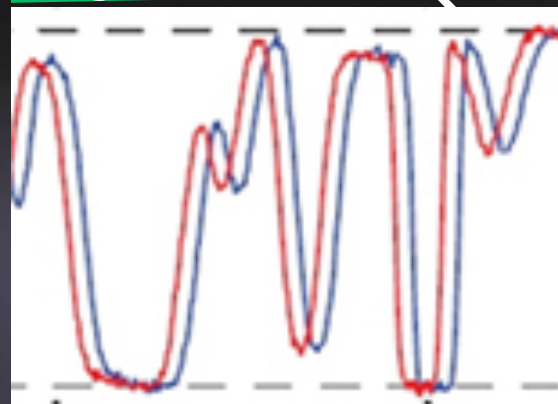
Expanding Universe

Big Bang

$$\dot{z} = (1+z)H_0 - H(z)$$



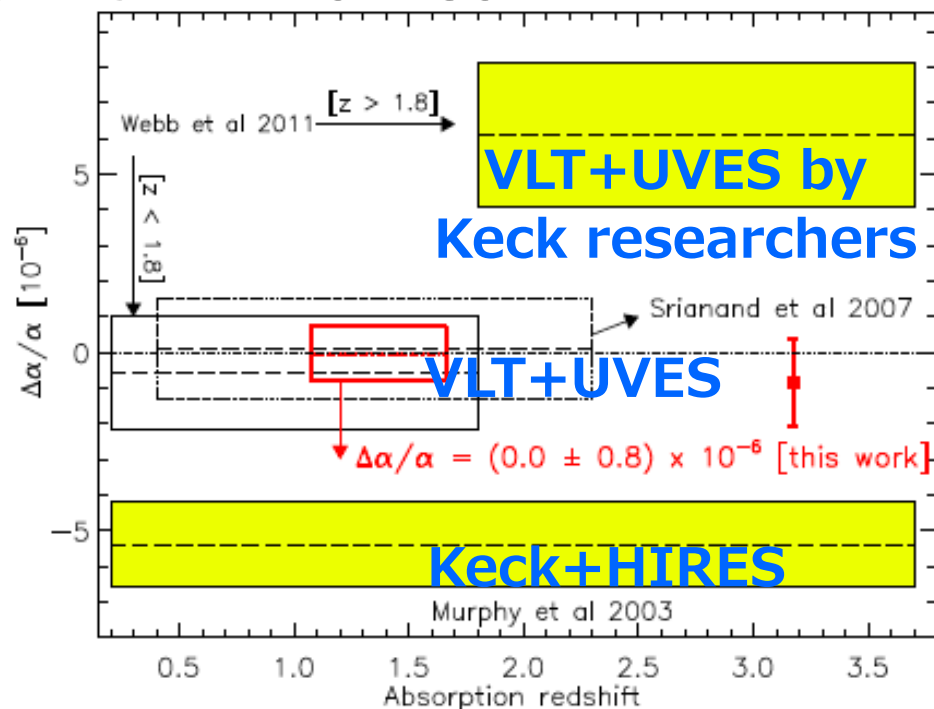
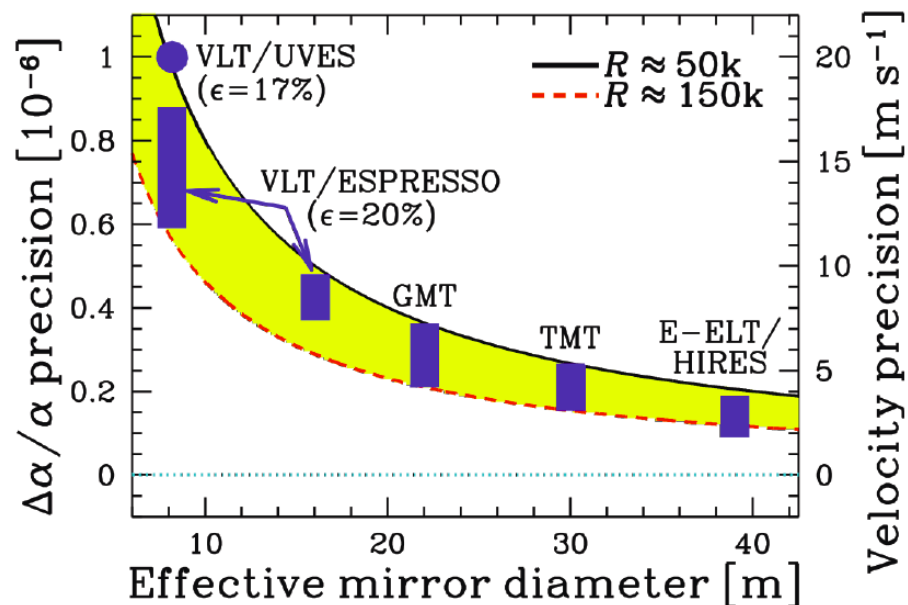
Very accurate
measurements (6cm/s)
of Lyman α absorption
lines in ~ 10 years



- ◆ High S/N and R obs. for ~ 20 Ly α forests over >10 years
- ◆ Direct and model-independent measure of the expansion history

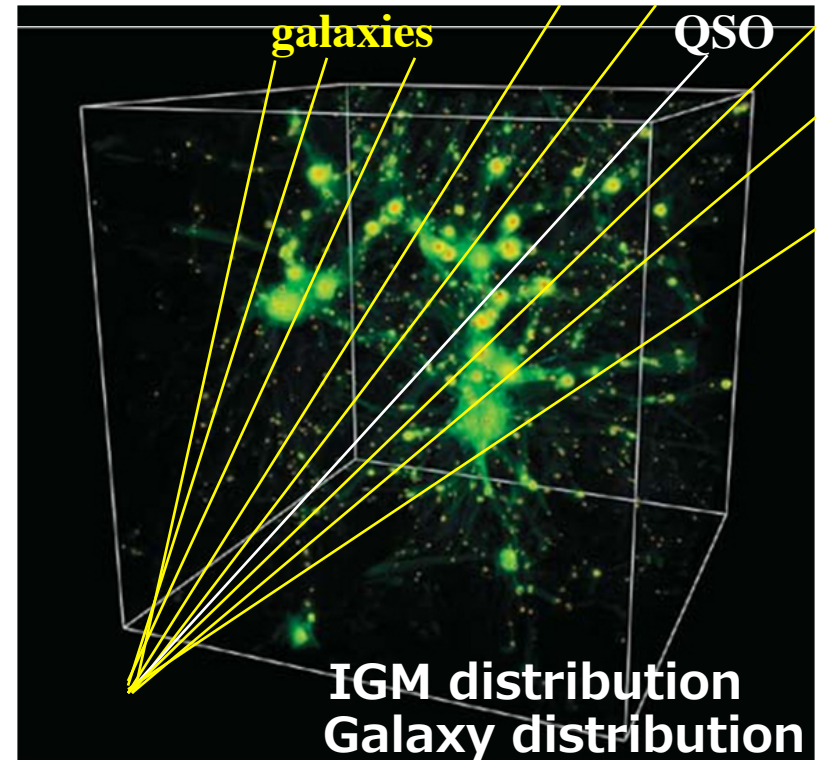
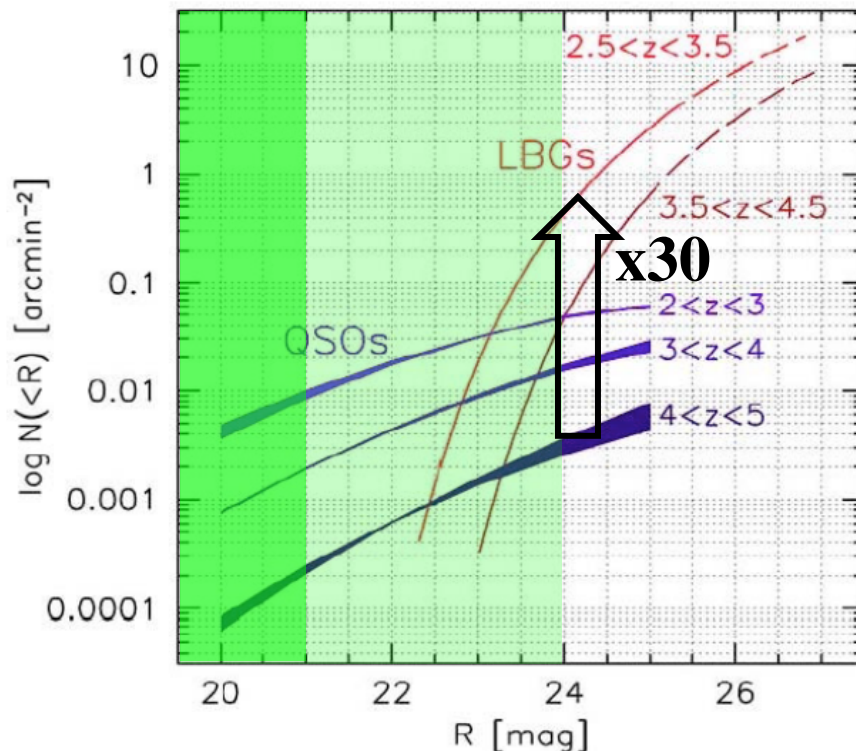
(2) Constraining the spatial & temporal variation of the fine structure constant

- ◆ Fine structure constant $\alpha \equiv 2\pi e^2 / hc$
- ◆ FeII & ZnII lines in QSO absorptions are used for the constraint (λ difference $\propto \alpha^2$)
- ◆ Constrain in laboratory (yr^{-1}) $\Delta\alpha / \alpha = (-1.6 \pm 2.3) \times 10^{-17}$
- ◆ Unsolved by observations with VLT & Keck

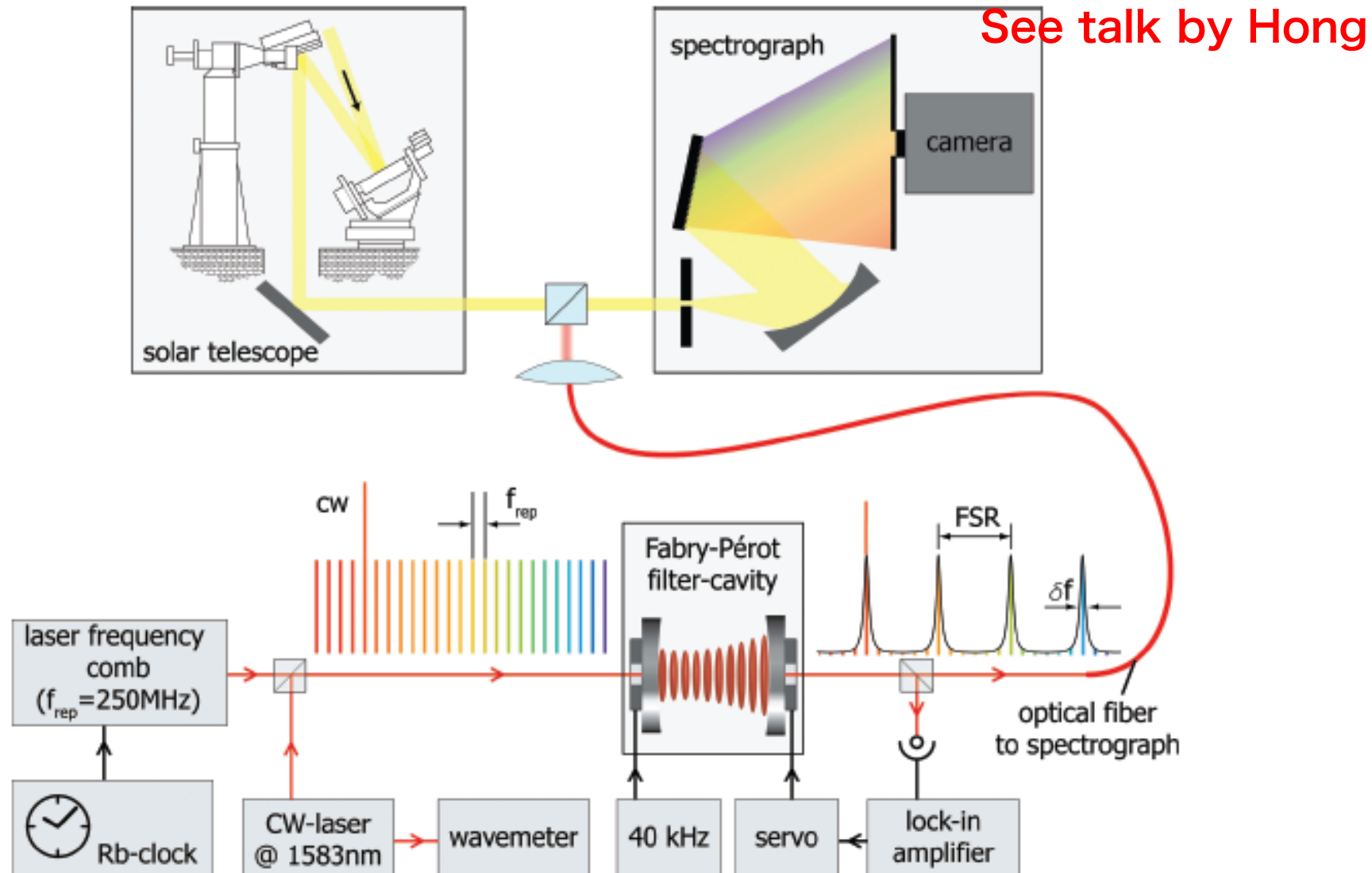


(3) IGM tomography

- ◆ Limiting mag ~ 24 with $R=10,000$
- ◆ Number density of galaxies is $>$ QSO ($2/\text{arcmin}^2$)
- ◆ Enable to study physical and chemical properties (e.g, spatial distribution, kinematic pattern, chemical composition, ionization condition, volume density, and gas temperature) in $d < 300$ kpc



Optical Laser Frequency Comb for very accurate λ calibrations



Optical Laser Frequency Comb for very accurate λ calibrations



Wavelength (900~2,200 nm) See talk by Hong

Summary

- ◆ TMT on-site construction has started in 2014. Design and fabrication works are on going in the partner countries.
- ◆ 3 Science Topics:
 - (1) Direct Detection of Cosmic Acceleration
 - (2) Constraining the spatial & temporal variation of the fine structure constant
 - (3) IGM tomography by QSOs & galaxies
Absorption Lines
- ◆ Technical R&D Plans
 - ◆ Develop Optical Laser Frequency Comb for very accurate / stable λ Calibrations
 - ◆ Install the comb on Subaru HDS for on-site tests to know technical difficulties and establish the observation strategies for TMT