Hydrodynamics and beyond...

Romuald A. Janik

Jagiellonian University Kraków

M. Heller, RJ, P. Witaszczyk, 1302.0697 [PRL 110, 211602 (2013)] work in progress and results from 1103.3452, 1203.0755

Outline

What is hydrodynamics? (and higher-order hydrodynamics?)

Some questions

Methods: fluid/gravity duality and boost-invariant flow

The Borel plane and gradient expansion

Borel resummed hydrodynamics

Comparisions with numerics

Conclusions

What is hydrodynamics?

- Universal description of the long wavelength degrees of freedom
- Applies equally well at macroscopic and microscopic scales
- Current most relevant example: quark-gluon-plasma produced at RHIC/LHC

Long wavelength description \equiv gradient expansion \equiv expansion in the number of derivatives

What is hydrodynamics?

- Universal description of the long wavelength degrees of freedom
- Applies equally well at macroscopic and microscopic scales
- Current most relevant example: quark-gluon-plasma produced at RHIC/LHC

Long wavelength description \equiv gradient expansion \equiv expansion in the number of derivatives

What is hydrodynamics?

Universal description of the long wavelength degrees of freedom

- Applies equally well at macroscopic and microscopic scales
- Current most relevant example: quark-gluon-plasma produced at RHIC/LHC

Long wavelength description \equiv gradient expansion \equiv expansion in the number of derivatives

What is hydrodynamics?

- Universal description of the long wavelength degrees of freedom
- Applies equally well at macroscopic and microscopic scales
- Current most relevant example: quark-gluon-plasma produced at RHIC/LHC

Long wavelength description \equiv gradient expansion \equiv expansion in the number of derivatives

What is hydrodynamics?

- Universal description of the long wavelength degrees of freedom
- Applies equally well at macroscopic and microscopic scales
- Current most relevant example: quark-gluon-plasma produced at RHIC/LHC

Long wavelength description \equiv gradient expansion \equiv expansion in the number of derivatives

What is hydrodynamics?

- Universal description of the long wavelength degrees of freedom
- Applies equally well at macroscopic and microscopic scales
- Current most relevant example: quark-gluon-plasma produced at RHIC/LHC

Long wavelength description \equiv gradient expansion \equiv expansion in the number of derivatives

What is hydrodynamics?

- Universal description of the long wavelength degrees of freedom
- Applies equally well at macroscopic and microscopic scales
- Current most relevant example: quark-gluon-plasma produced at RHIC/LHC

Long wavelength description \equiv gradient expansion \equiv expansion in the number of derivatives

What is hydrodynamics?

- Universal description of the long wavelength degrees of freedom
- Applies equally well at macroscopic and microscopic scales
- Current most relevant example: quark-gluon-plasma produced at RHIC/LHC

Long wavelength description \equiv gradient expansion \equiv expansion in the number of derivatives

What is hydrodynamics?

- Universal description of the long wavelength degrees of freedom
- Applies equally well at macroscopic and microscopic scales
- Current most relevant example: quark-gluon-plasma produced at RHIC/LHC

Long wavelength description \equiv gradient expansion \equiv expansion in the number of derivatives

Hydrodynamics:

- 1. Concentrates on the dynamics of the energy-momentum tensor $T_{\mu
 u}$
- 2. Amounts to the assumption that $T_{\mu\nu}$ is wholly expressed through the flow velocity u^{μ} , energy density and pressure (E = 3p for conformal fluids)
- 3. Arrange all possible terms by the number of derivatives of u^{μ}
- **5.** Generalized Navier-Stokes equation is just $\partial_{\mu}T^{\mu\nu} = 0$
- $\mathcal{N} = 4$ SYM hydrodynamics:

$$T_{rescaled}^{\mu\nu} = \underbrace{(\pi T)^4 (\eta^{\mu\nu} + 4u^{\mu}u^{\nu})}_{perfect \ fluid} - \underbrace{2(\pi T)^3 \sigma^{\mu\nu}}_{viscosity} + \underbrace{(\pi T^2) \left(\log 2T_{2a}^{\mu\nu} + 2T_{2b}^{\mu\nu} + (2 - \log 2) \left(\frac{1}{3}T_{2c}^{\mu\nu} + T_{2d}^{\mu\nu} + T_{2e}^{\mu\nu}\right)\right)}_{rescaled}$$

second order hydrodynamics

Hydrodynamics:

- 1. Concentrates on the dynamics of the energy-momentum tensor $T_{\mu
 u}$
- 2. Amounts to the assumption that $T_{\mu\nu}$ is wholly expressed through the flow velocity u^{μ} , energy density and pressure (E = 3p for conformal fluids)
- 3. Arrange all possible terms by the number of derivatives of u^{μ}
- **5.** Generalized Navier-Stokes equation is just $\partial_{\mu}T^{\mu\nu} = 0$
- $\mathcal{N} = 4$ SYM hydrodynamics:

$$T_{rescaled}^{\mu\nu} = \underbrace{(\pi T)^4 (\eta^{\mu\nu} + 4u^{\mu}u^{\nu})}_{perfect \ fluid} - \underbrace{2(\pi T)^3 \sigma^{\mu\nu}}_{viscosity} + \underbrace{(\pi T^2) \left(\log 2T_{2a}^{\mu\nu} + 2T_{2b}^{\mu\nu} + (2 - \log 2) \left(\frac{1}{3}T_{2c}^{\mu\nu} + T_{2d}^{\mu\nu} + T_{2e}^{\mu\nu}\right)\right)}_{rescaled}$$

second order hydrodynamics

Hydrodynamics:

- 1. Concentrates on the dynamics of the energy-momentum tensor $T_{\mu
 u}$
- 2. Amounts to the assumption that $T_{\mu\nu}$ is wholly expressed through the flow velocity u^{μ} , energy density and pressure (E = 3p for conformal fluids)
- **3.** Arrange all possible terms by the number of derivatives of u^μ
- Coefficients of these terms ≡ transport coefficients characteristic of the underlying microscopic theory
- **5.** Generalized Navier-Stokes equation is just $\partial_{\mu}T^{\mu\nu} = 0$
- $\mathcal{N} = 4$ SYM hydrodynamics:

$$T_{rescaled}^{\mu\nu} = \underbrace{(\pi T)^4 (\eta^{\mu\nu} + 4u^{\mu}u^{\nu})}_{perfect \ fluid} - \underbrace{2(\pi T)^3 \sigma^{\mu\nu}}_{viscosity} + \underbrace{(\pi T^2) \left(\log 2T_{2a}^{\mu\nu} + 2T_{2b}^{\mu\nu} + (2 - \log 2) \left(\frac{1}{3}T_{2c}^{\mu\nu} + T_{2d}^{\mu\nu} + T_{2e}^{\mu\nu}\right)\right)}_{delta}$$

second order hydrodynamics

Hydrodynamics:

- 1. Concentrates on the dynamics of the energy-momentum tensor $T_{\mu\nu}$
- 2. Amounts to the assumption that $T_{\mu\nu}$ is wholly expressed through the flow velocity u^{μ} , energy density and pressure (E = 3p for conformal fluids)
- 3. Arrange all possible terms by the number of derivatives of u^{μ}
- **5.** Generalized Navier-Stokes equation is just $\partial_{\mu}T^{\mu\nu} = 0$

 $\mathcal{N} = 4$ SYM hydrodynamics:

$$T_{rescaled}^{\mu\nu} = \underbrace{(\pi T)^4 (\eta^{\mu\nu} + 4u^{\mu}u^{\nu})}_{perfect \ fluid} - \underbrace{2(\pi T)^3 \sigma^{\mu\nu}}_{viscosity} + \underbrace{(\pi T^2) \left(\log 2T_{2a}^{\mu\nu} + 2T_{2b}^{\mu\nu} + (2 - \log 2) \left(\frac{1}{3}T_{2c}^{\mu\nu} + T_{2d}^{\mu\nu} + T_{2e}^{\mu\nu}\right)\right)}_{rescaled}$$

second order hydrodynamics

Hydrodynamics:

- 1. Concentrates on the dynamics of the energy-momentum tensor $T_{\mu\nu}$
- 2. Amounts to the assumption that $T_{\mu\nu}$ is wholly expressed through the flow velocity u^{μ} , energy density and pressure (E = 3p for conformal fluids)
- 3. Arrange all possible terms by the number of derivatives of u^{μ}
- 4. Coefficients of these terms \equiv transport coefficients characteristic of the underlying microscopic theory
- **5.** Generalized Navier-Stokes equation is just $\partial_{\mu}T^{\mu\nu} = 0$
- $\mathcal{N} = 4$ SYM hydrodynamics:

$$T_{\text{rescaled}}^{\mu\nu} = \underbrace{(\pi T)^4 (\eta^{\mu\nu} + 4u^{\mu}u^{\nu})}_{\text{perfect fluid}} - \underbrace{2(\pi T)^3 \sigma^{\mu\nu}}_{\text{viscosity}} + \\ + \underbrace{(\pi T^2) \left(\log 2T_{2a}^{\mu\nu} + 2T_{2b}^{\mu\nu} + (2 - \log 2) \left(\frac{1}{3}T_{2c}^{\mu\nu} + T_{2d}^{\mu\nu} + T_{2e}^{\mu\nu}\right)\right)}_{\text{scosity}}$$

second order hydrodynamics

Hydrodynamics:

- 1. Concentrates on the dynamics of the energy-momentum tensor $T_{\mu\nu}$
- 2. Amounts to the assumption that $T_{\mu\nu}$ is wholly expressed through the flow velocity u^{μ} , energy density and pressure (E = 3p for conformal fluids)
- 3. Arrange all possible terms by the number of derivatives of u^{μ}
- 4. Coefficients of these terms \equiv transport coefficients characteristic of the underlying microscopic theory
- **5.** Generalized Navier-Stokes equation is just $\partial_{\mu}T^{\mu\nu} = 0$

$\mathcal{N} = 4$ SYM hydrodynamics:

$$T_{rescaled}^{\mu\nu} = \underbrace{(\pi T)^4 (\eta^{\mu\nu} + 4u^{\mu}u^{\nu})}_{perfect \ fluid} - \underbrace{2(\pi T)^3 \sigma^{\mu\nu}}_{viscosity} + (\pi T^2) \left(\log 2T_{2a}^{\mu\nu} + 2T_{2b}^{\mu\nu} + (2 - \log 2) \left(\frac{1}{3}T_{2c}^{\mu\nu} + T_{2d}^{\mu\nu} + T_{2e}^{\mu\nu}\right)\right)$$

second order hydrodynamics

Hydrodynamics:

- 1. Concentrates on the dynamics of the energy-momentum tensor $T_{\mu\nu}$
- 2. Amounts to the assumption that $T_{\mu\nu}$ is wholly expressed through the flow velocity u^{μ} , energy density and pressure (E = 3p for conformal fluids)
- 3. Arrange all possible terms by the number of derivatives of u^{μ}
- 4. Coefficients of these terms \equiv transport coefficients characteristic of the underlying microscopic theory
- **5.** Generalized Navier-Stokes equation is just $\partial_{\mu}T^{\mu\nu} = 0$
- $\mathcal{N}=4$ SYM hydrodynamics:

$$T_{rescaled}^{\mu\nu} = \underbrace{(\pi T)^4 (\eta^{\mu\nu} + 4u^{\mu}u^{\nu})}_{perfect \ fluid} - \underbrace{2(\pi T)^3 \sigma^{\mu\nu}}_{viscosity} + \underbrace{(\pi T^2) \left(\log 2T_{2a}^{\mu\nu} + 2T_{2b}^{\mu\nu} + (2 - \log 2) \left(\frac{1}{3}T_{2c}^{\mu\nu} + T_{2d}^{\mu\nu} + T_{2e}^{\mu\nu}\right)\right)}_{rescaled \ order \ budgetures integrated and the second an$$

second order hydrodynamics

Bhattacharya, Hubeny, Minwalla, Rangamani

4 / 32

What is the nature of the gradient expansion?

- Suppose we include terms with more and more derivatives (dissipation)
- Is the resulting series asymptotic (zero radius of convergence)?
- What physics is (quantitatively) responsible for the lack of convergence?

Analogy: perturbative expansion and instanton effects...

Question 2

If the hydrodynamic series is only asymptotic, is it Borel summable?

- What are the singularities on the Borel plane and what is their physical interpretation?
- Can we explicitly construct Borel resummed hydrodynamics?

What is the nature of the gradient expansion?

- Suppose we include terms with more and more derivatives (dissipation)
- Is the resulting series asymptotic (zero radius of convergence)?
- What physics is (quantitatively) responsible for the lack of convergence?

Analogy: perturbative expansion and instanton effects...

Question 2

If the hydrodynamic series is only asymptotic, is it Borel summable?

- What are the singularities on the Borel plane and what is their physical interpretation?
- Can we explicitly construct Borel resummed hydrodynamics?

What is the nature of the gradient expansion?

- Suppose we include terms with more and more derivatives (dissipation)
- Is the resulting series asymptotic (zero radius of convergence)?
- What physics is (quantitatively) responsible for the lack of convergence?

Analogy: perturbative expansion and instanton effects...

Question 2

If the hydrodynamic series is only asymptotic, is it Borel summable?

- What are the singularities on the Borel plane and what is their physical interpretation?
- Can we explicitly construct Borel resummed hydrodynamics?

What is the nature of the gradient expansion?

- Suppose we include terms with more and more derivatives (dissipation)
- Is the resulting series asymptotic (zero radius of convergence)?
- What physics is (quantitatively) responsible for the lack of convergence?

Analogy: perturbative expansion and instanton effects...

Question 2

If the hydrodynamic series is only asymptotic, is it Borel summable?

- What are the singularities on the Borel plane and what is their physical interpretation?
- Can we explicitly construct Borel resummed hydrodynamics?

What is the nature of the gradient expansion?

- Suppose we include terms with more and more derivatives (dissipation)
- Is the resulting series asymptotic (zero radius of convergence)?
- What physics is (quantitatively) responsible for the lack of convergence?

Analogy: perturbative expansion and instanton effects...

Question 2

If the hydrodynamic series is only asymptotic, is it Borel summable?

- What are the singularities on the Borel plane and what is their physical interpretation?
- Can we explicitly construct Borel resummed hydrodynamics?

What is the nature of the gradient expansion?

- Suppose we include terms with more and more derivatives (dissipation)
- Is the resulting series asymptotic (zero radius of convergence)?
- What physics is (quantitatively) responsible for the lack of convergence?

Analogy: perturbative expansion and instanton effects...

Question 2

If the hydrodynamic series is only asymptotic, is it Borel summable?

- What are the singularities on the Borel plane and what is their physical interpretation?
- Can we explicitly construct Borel resummed hydrodynamics?

What is the nature of the gradient expansion?

- Suppose we include terms with more and more derivatives (dissipation)
- Is the resulting series asymptotic (zero radius of convergence)?
- What physics is (quantitatively) responsible for the lack of convergence?

Analogy: perturbative expansion and instanton effects...

Question 2

If the hydrodynamic series is only asymptotic, is it Borel summable?

- What are the singularities on the Borel plane and what is their physical interpretation?
- Can we explicitly construct Borel resummed hydrodynamics?

What is the nature of the gradient expansion?

- Suppose we include terms with more and more derivatives (dissipation)
- Is the resulting series asymptotic (zero radius of convergence)?
- What physics is (quantitatively) responsible for the lack of convergence?

Analogy: perturbative expansion and instanton effects...

Question 2

If the hydrodynamic series is only asymptotic, is it Borel summable?

- What are the singularities on the Borel plane and what is their physical interpretation?
- Can we explicitly construct Borel resummed hydrodynamics?

What is the nature of the gradient expansion?

- Suppose we include terms with more and more derivatives (dissipation)
- Is the resulting series asymptotic (zero radius of convergence)?
- What physics is (quantitatively) responsible for the lack of convergence?

Analogy: perturbative expansion and instanton effects...

Question 2

If the hydrodynamic series is only asymptotic, is it Borel summable?

- What are the singularities on the Borel plane and what is their physical interpretation?
- Can we explicitly construct Borel resummed hydrodynamics?

What is the nature of the gradient expansion?

- Suppose we include terms with more and more derivatives (dissipation)
- Is the resulting series asymptotic (zero radius of convergence)?
- What physics is (quantitatively) responsible for the lack of convergence?

Analogy: perturbative expansion and instanton effects...

Question 2

If the hydrodynamic series is only asymptotic, is it Borel summable?

- What are the singularities on the Borel plane and what is their physical interpretation?
- Can we explicitly construct Borel resummed hydrodynamics?

What is the nature of the gradient expansion?

- Suppose we include terms with more and more derivatives (dissipation)
- Is the resulting series asymptotic (zero radius of convergence)?
- What physics is (quantitatively) responsible for the lack of convergence?

Analogy: perturbative expansion and instanton effects...

Question 2

If the hydrodynamic series is only asymptotic, is it Borel summable?

- What are the singularities on the Borel plane and what is their physical interpretation?
- Can we explicitly construct Borel resummed hydrodynamics?

Is there any practical motivation for looking at high order hydrodynamics?

- Lublinsky, Shuryak proposed a resummation based on linearized hydrodynamic modes from AdS/CFT with phenomenological motivation
- In our previous work [Heller, RJ, Witaszczyk] we considered the evolution of a spacetime dual to a plasma system evolving from some nonequilibrium initial conditions and its transition to hydrodynamics

Large and intermediate initial data

'Small' initial data

Is there any practical motivation for looking at high order hydrodynamics?

- Lublinsky, Shuryak proposed a resummation based on linearized hydrodynamic modes from AdS/CFT with phenomenological motivation
- In our previous work [Heller, RJ, Witaszczyk] we considered the evolution of a spacetime dual to a plasma system evolving from some nonequilibrium initial conditions and its transition to hydrodynamics

Large and intermediate initial data

'Small' initial data

Is there any practical motivation for looking at high order hydrodynamics?

- Lublinsky, Shuryak proposed a resummation based on linearized hydrodynamic modes from AdS/CFT with phenomenological motivation
- In our previous work [Heller, RJ, Witaszczyk] we considered the evolution of a spacetime dual to a plasma system evolving from some nonequilibrium initial conditions and its transition to hydrodynamics

Large and intermediate initial data

'Small' initial data

Is there any practical motivation for looking at high order hydrodynamics?

- Lublinsky, Shuryak proposed a resummation based on linearized hydrodynamic modes from AdS/CFT with phenomenological motivation
- In our previous work [Heller, RJ, Witaszczyk] we considered the evolution of a spacetime dual to a plasma system evolving from some nonequilibrium initial conditions and its transition to hydrodynamics

Large and intermediate initial data

'Small' initial data

Is there any practical motivation for looking at high order hydrodynamics?

- Lublinsky, Shuryak proposed a resummation based on linearized hydrodynamic modes from AdS/CFT with phenomenological motivation
- In our previous work [Heller, RJ, Witaszczyk] we considered the evolution of a spacetime dual to a plasma system evolving from some nonequilibrium initial conditions and its transition to hydrodynamics

Large and intermediate initial data

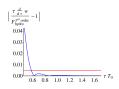
'Small' initial data

Is there any practical motivation for looking at high order hydrodynamics?

- Lublinsky, Shuryak proposed a resummation based on linearized hydrodynamic modes from AdS/CFT with phenomenological motivation
- In our previous work [Heller, RJ, Witaszczyk] we considered the evolution of a spacetime dual to a plasma system evolving from some nonequilibrium initial conditions and its transition to hydrodynamics



'Small' initial data

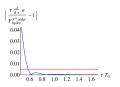


Is there any practical motivation for looking at high order hydrodynamics?

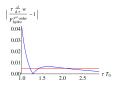
- Lublinsky, Shuryak proposed a resummation based on linearized hydrodynamic modes from AdS/CFT with phenomenological motivation
- In our previous work [Heller, RJ, Witaszczyk] we considered the evolution of a spacetime dual to a plasma system evolving from some nonequilibrium initial conditions and its transition to hydrodynamics

Large and intermediate

initial data



'Small' initial data



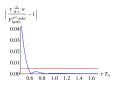
Question 3

Is there any practical motivation for looking at high order hydrodynamics?

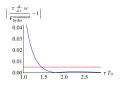
- Lublinsky, Shuryak proposed a resummation based on linearized hydrodynamic modes from AdS/CFT with phenomenological motivation
- In our previous work [Heller, RJ, Witaszczyk] we considered the evolution of a spacetime dual to a plasma system evolving from some nonequilibrium initial conditions and its transition to hydrodynamics

Large and intermediate

initial data



'Small' initial data



What are the deviations from all-order hydrodynamics? (possible phenomenological models?)

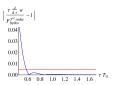
Question 3

Is there any practical motivation for looking at high order hydrodynamics?

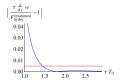
- Lublinsky, Shuryak proposed a resummation based on linearized hydrodynamic modes from AdS/CFT with phenomenological motivation
- In our previous work [Heller, RJ, Witaszczyk] we considered the evolution of a spacetime dual to a plasma system evolving from some nonequilibrium initial conditions and its transition to hydrodynamics

Large and intermediate

initial data



'Small' initial data



What are the deviations from all-order hydrodynamics? (possible phenomenological models?)

Final motivation:

 $\mathcal{N}=4$ SYM provides for us (through the AdS/CFT correspondence) the **only** physical system for which one can systematically compute high orders of the gradient expansion and examine the above questions

Final motivation:

 $\mathcal{N}=4$ SYM provides for us (through the AdS/CFT correspondence) the **only** physical system for which one can systematically compute high orders of the gradient expansion and examine the above questions

- Approach pioneered by Bhattacharya, Hubeny, Minwalla, Rangamani
- One starts with a boosted planar black hole representing a plasma system moving with uniform velocity u^µ and with temperature T
- ▶ One promotes u^µ and T to slowly varying functions one has to correct the metric iteratively in an expansion in gradients
- At each order one looks for a (regular) solution of

- Rather complicated to perform the expansion analytically:
 in general carried out to 2nd order (2nd order viscous hydrodynamics)
 - in boost-invariant case up to 3rd order

Approach pioneered by Bhattacharya, Hubeny, Minwalla, Rangamani

- One starts with a boosted planar black hole representing a plasma system moving with uniform velocity u^µ and with temperature T
- ▶ One promotes u^µ and T to slowly varying functions one has to correct the metric iteratively in an expansion in gradients
- At each order one looks for a (regular) solution of

- Rather complicated to perform the expansion analytically:
 in general carried out to 2nd order (2nd order viscous hydrodynamics)
 - in boost-invariant case up to 3rd order

- Approach pioneered by Bhattacharya, Hubeny, Minwalla, Rangamani
- One starts with a boosted planar black hole representing a plasma system moving with uniform velocity u^µ and with temperature T
- One promotes u^µ and T to slowly varying functions one has to correct the metric iteratively in an expansion in gradients
- At each order one looks for a (regular) solution of

- Rather complicated to perform the expansion analytically:
 in general carried out to 2nd order (2nd order viscous hydrodynamics)
 - in boost-invariant case up to 3rd order

- Approach pioneered by Bhattacharya, Hubeny, Minwalla, Rangamani
- One starts with a boosted planar black hole representing a plasma system moving with uniform velocity u^µ and with temperature T
- One promotes u^µ and T to slowly varying functions one has to correct the metric iteratively in an expansion in gradients
- At each order one looks for a (regular) solution of

- Rather complicated to perform the expansion analytically:
 in general carried out to 2nd order (2nd order viscous hydrodynamics)
 - in boost-invariant case up to 3rd order

- Approach pioneered by Bhattacharya, Hubeny, Minwalla, Rangamani
- One starts with a boosted planar black hole representing a plasma system moving with uniform velocity u^µ and with temperature T
- ► One promotes u^µ and T to slowly varying functions one has to correct the metric iteratively in an expansion in gradients
- At each order one looks for a (regular) solution of

- Rather complicated to perform the expansion analytically:
 in general carried out to 2nd order (2nd order viscous hydrodynamics)
 - in boost-invariant case up to 3rd order

- Approach pioneered by Bhattacharya, Hubeny, Minwalla, Rangamani
- One starts with a boosted planar black hole representing a plasma system moving with uniform velocity u^µ and with temperature T
- ► One promotes u^µ and T to slowly varying functions one has to correct the metric iteratively in an expansion in gradients
- At each order one looks for a (regular) solution of

- Rather complicated to perform the expansion analytically:
 in general carried out to 2nd order (2nd order viscous hydrodynamics)
 - in boost-invariant case up to 3rd order

- Approach pioneered by Bhattacharya, Hubeny, Minwalla, Rangamani
- One starts with a boosted planar black hole representing a plasma system moving with uniform velocity u^µ and with temperature T
- One promotes u^µ and T to slowly varying functions one has to correct the metric iteratively in an expansion in gradients
- At each order one looks for a (regular) solution of

- Rather complicated to perform the expansion analytically:
 in general carried out to 2nd order (2nd order viscous hydrodynamics)
 - in boost-invariant case up to 3rd order

- Approach pioneered by Bhattacharya, Hubeny, Minwalla, Rangamani
- One starts with a boosted planar black hole representing a plasma system moving with uniform velocity u^µ and with temperature T
- ► One promotes u^µ and T to slowly varying functions one has to correct the metric iteratively in an expansion in gradients
- At each order one looks for a (regular) solution of

- Rather complicated to perform the expansion analytically:
 in general carried out to 2nd order (2nd order viscous hydrodynamics)
 - in boost-invariant case up to 3rd order

- At each order we have a set of coupled *linear* ODE's
- Very simple to solve numerically (with very high precision!) Much simpler than normal numerical relativity which involves solving nonlinear PDE's!!

At each order we have a set of coupled linear ODE's

Very simple to solve numerically (with very high precision!) Much simpler than normal numerical relativity which involves solving nonlinear PDE's!!

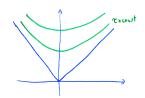
- At each order we have a set of coupled *linear* ODE's
- Very simple to solve numerically (with very high precision!) Much simpler than normal numerical relativity which involves solving nonlinear PDE's!!

- At each order we have a set of coupled linear ODE's
- Very simple to solve numerically (with very high precision!) Much simpler than normal numerical relativity which involves solving nonlinear PDE's!!

- At each order we have a set of coupled *linear* ODE's
- Very simple to solve numerically (with very high precision!) Much simpler than normal numerical relativity which involves solving nonlinear PDE's!!

Bjorken '83

Assume a flow that is invariant under longitudinal boosts and does not depend on the transverse coordinates.

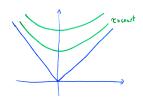


- In a conformal theory, T^μ_μ = 0 and ∂_μT^{μν} = 0 determine, under the above assumptions, the energy-momentum tensor completely in terms of a single function ε(τ), the energy density at mid-rapidity.
- ► The assumptions of symmetry fix uniquely the flow velocity
- Gradient expansion coincides with an expansion in

$$\frac{1}{\tau^{\frac{2}{3}}}$$

Bjorken '83

Assume a flow that is invariant under longitudinal boosts and does not depend on the transverse coordinates.

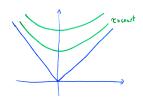


- In a conformal theory, T^μ_μ = 0 and ∂_μT^{μν} = 0 determine, under the above assumptions, the energy-momentum tensor completely in terms of a single function ε(τ), the energy density at mid-rapidity.
- ► The assumptions of symmetry fix uniquely the flow velocity
- Gradient expansion coincides with an expansion in

 $\frac{1}{\tau^{\frac{2}{3}}}$

Bjorken '83

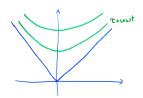
Assume a flow that is invariant under longitudinal boosts and does not depend on the transverse coordinates.



- ▶ In a conformal theory, $T^{\mu}_{\mu} = 0$ and $\partial_{\mu} T^{\mu\nu} = 0$ determine, under the above assumptions, the energy-momentum tensor completely in terms of a single function $\varepsilon(\tau)$, the energy density at mid-rapidity.
- The assumptions of symmetry fix uniquely the flow velocity
- Gradient expansion coincides with an expansion in

Bjorken '83

Assume a flow that is invariant under longitudinal boosts and does not depend on the transverse coordinates.

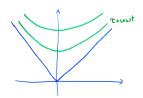


- ▶ In a conformal theory, $T^{\mu}_{\mu} = 0$ and $\partial_{\mu} T^{\mu\nu} = 0$ determine, under the above assumptions, the energy-momentum tensor completely in terms of a single function $\varepsilon(\tau)$, the energy density at mid-rapidity.
- ► The assumptions of symmetry fix uniquely the flow velocity
- Gradient expansion coincides with an expansion in

$\frac{1}{\tau^{\frac{2}{3}}}$

Bjorken '83

Assume a flow that is invariant under longitudinal boosts and does not depend on the transverse coordinates.



- In a conformal theory, T^μ_μ = 0 and ∂_μ T^{μν} = 0 determine, under the above assumptions, the energy-momentum tensor completely in terms of a single function ε(τ), the energy density at mid-rapidity.
- ► The assumptions of symmetry fix uniquely the flow velocity
- Gradient expansion coincides with an expansion in

$\frac{1}{\tau^{\frac{2}{3}}}$

• Structure of the analytical result for large τ :

$$\varepsilon(\tau) = \frac{1}{\tau^{\frac{4}{3}}} - \frac{2}{2^{\frac{1}{2}}3^{\frac{3}{4}}} \frac{1}{\tau^2} + \frac{1 + 2\log 2}{12\sqrt{3}} \frac{1}{\tau^{\frac{8}{3}}} + \frac{-3 + 2\pi^2 + 24\log 2 - 24\log^2 2}{324 \cdot 2^{\frac{1}{2}}3^{\frac{1}{4}}} \frac{1}{\tau^{\frac{10}{3}}} + \dots$$

RJ, Peschanski; Nakamura, S-J Sin; RJ; RJ, Heller; Heller

Leading term — perfect fluid behaviour second term — 1st order viscous hydrodynamics third term — 2nd order viscous hydrodynamics fourth term — 3rd order viscous hydrodynamics...

► In general:

$$\varepsilon(\tau) = \sum_{n=2}^{\infty} \frac{\varepsilon_n}{\tau^{\frac{2n}{3}}}$$

• Structure of the analytical result for large τ :

$$\varepsilon(\tau) = \frac{1}{\tau^{\frac{4}{3}}} - \frac{2}{2^{\frac{1}{2}}3^{\frac{3}{4}}} \frac{1}{\tau^2} + \frac{1 + 2\log 2}{12\sqrt{3}} \frac{1}{\tau^{\frac{8}{3}}} + \frac{-3 + 2\pi^2 + 24\log 2 - 24\log^2 2}{324 \cdot 2^{\frac{1}{2}}3^{\frac{1}{4}}} \frac{1}{\tau^{\frac{10}{3}}} + \dots$$

RJ, Peschanski; Nakamura, S-J Sin; RJ; RJ, Heller; Heller

Leading term — perfect fluid behaviour

second term — 1^{st} order viscous hydrodynamics third term — 2^{nd} order viscous hydrodynamics fourth term — 3^{rd} order viscous hydrodynamics...

► In general:

$$\varepsilon(\tau) = \sum_{n=2}^{\infty} \frac{\varepsilon_n}{\tau^{\frac{2n}{3}}}$$

Structure of the analytical result for large τ:

$$\varepsilon(\tau) = \frac{1}{\tau^{\frac{4}{3}}} - \frac{2}{2^{\frac{1}{2}}3^{\frac{3}{4}}} \frac{1}{\tau^2} + \frac{1 + 2\log 2}{12\sqrt{3}} \frac{1}{\tau^{\frac{8}{3}}} + \frac{-3 + 2\pi^2 + 24\log 2 - 24\log^2 2}{324 \cdot 2^{\frac{1}{2}}3^{\frac{1}{4}}} \frac{1}{\tau^{\frac{10}{3}}} + \dots$$

RJ, Peschanski; Nakamura, S-J Sin; RJ; RJ, Heller; Heller

 Leading term — perfect fluid behaviour second term — 1st order viscous hydrodynamics

third term — 2^{na} order viscous hydrodynamics fourth term — 3rd order viscous hydrodynamics...

► In general:

$$\varepsilon(\tau) = \sum_{n=2}^{\infty} \frac{\varepsilon_n}{\tau^{\frac{2n}{3}}}$$

• Structure of the analytical result for large τ :

$$\varepsilon(\tau) = \frac{1}{\tau^{\frac{4}{3}}} - \frac{2}{2^{\frac{1}{2}}3^{\frac{3}{4}}} \frac{1}{\tau^2} + \frac{1 + 2\log 2}{12\sqrt{3}} \frac{1}{\tau^{\frac{8}{3}}} + \frac{-3 + 2\pi^2 + 24\log 2 - 24\log^2 2}{324 \cdot 2^{\frac{1}{2}}3^{\frac{1}{4}}} \frac{1}{\tau^{\frac{10}{3}}} + \dots$$

RJ, Peschanski; Nakamura, S-J Sin; RJ; RJ, Heller; Heller

 Leading term — perfect fluid behaviour second term — 1st order viscous hydrodynamics third term — 2nd order viscous hydrodynamics fourth term — 3rd order viscous hydrodynamics...

► In general:

$$\varepsilon(\tau) = \sum_{n=2}^{\infty} \frac{\varepsilon_n}{\tau^{\frac{2n}{3}}}$$

• Structure of the analytical result for large τ :

$$\varepsilon(\tau) = \frac{1}{\tau^{\frac{4}{3}}} - \frac{2}{2^{\frac{1}{2}}3^{\frac{3}{4}}} \frac{1}{\tau^2} + \frac{1 + 2\log 2}{12\sqrt{3}} \frac{1}{\tau^{\frac{8}{3}}} + \frac{-3 + 2\pi^2 + 24\log 2 - 24\log^2 2}{324 \cdot 2^{\frac{1}{2}}3^{\frac{1}{4}}} \frac{1}{\tau^{\frac{10}{3}}} + \dots$$

RJ, Peschanski; Nakamura, S-J Sin; RJ; RJ, Heller; Heller

 Leading term — perfect fluid behaviour second term — 1st order viscous hydrodynamics third term — 2nd order viscous hydrodynamics fourth term — 3rd order viscous hydrodynamics...

▶ In general:



• Structure of the analytical result for large τ :

$$\varepsilon(\tau) = \frac{1}{\tau^{\frac{4}{3}}} - \frac{2}{2^{\frac{1}{2}}3^{\frac{3}{4}}} \frac{1}{\tau^2} + \frac{1 + 2\log 2}{12\sqrt{3}} \frac{1}{\tau^{\frac{8}{3}}} + \frac{-3 + 2\pi^2 + 24\log 2 - 24\log^2 2}{324 \cdot 2^{\frac{1}{2}}3^{\frac{1}{4}}} \frac{1}{\tau^{\frac{10}{3}}} + \dots$$

RJ, Peschanski; Nakamura, S-J Sin; RJ; RJ, Heller; Heller

Leading term — perfect fluid behaviour second term — 1st order viscous hydrodynamics third term — 2nd order viscous hydrodynamics fourth term — 3rd order viscous hydrodynamics...

► In general:

$$\varepsilon(\tau) = \sum_{n=2}^{\infty} \frac{\varepsilon_n}{\tau^{\frac{2n}{3}}}$$

• Structure of the analytical result for large τ :

$$\varepsilon(\tau) = \frac{1}{\tau^{\frac{4}{3}}} - \frac{2}{2^{\frac{1}{2}}3^{\frac{3}{4}}} \frac{1}{\tau^2} + \frac{1 + 2\log 2}{12\sqrt{3}} \frac{1}{\tau^{\frac{8}{3}}} + \frac{-3 + 2\pi^2 + 24\log 2 - 24\log^2 2}{324 \cdot 2^{\frac{1}{2}}3^{\frac{1}{4}}} \frac{1}{\tau^{\frac{10}{3}}} + \dots$$

RJ, Peschanski; Nakamura, S-J Sin; RJ; RJ, Heller; Heller

 Leading term — perfect fluid behaviour second term — 1st order viscous hydrodynamics third term — 2nd order viscous hydrodynamics fourth term — 3rd order viscous hydrodynamics...

► In general:

$$\varepsilon(\tau) = \sum_{n=2}^{\infty} \frac{\varepsilon_n}{\tau^{\frac{2n}{3}}}$$

▶ By iteratively solving numerically the linear ODE's from fluid/gravity duality, we obtained 240 coefficients in the gradient expansion

• chief complication – generate the r.h.s. of the equations

- to get to so high orders we need very high precision computations
- first couple of orders easy and fast

• Introduce $u \equiv 1/\tau^{2/3}$

 $\varepsilon(u) = \sum_{n=2}^{242} \varepsilon_n u^n$

Convergence

By iteratively solving numerically the linear ODE's from fluid/gravity duality, we obtained 240 coefficients in the gradient expansion

 $\varepsilon(\tau) = \sum_{n=2}^{242} \frac{\varepsilon_n}{\tau^{\frac{2n}{3}}}$

• chief complication – generate the r.h.s. of the equations

• to get to so high orders we need very high precision computations

• first couple of orders - easy and fast

• Introduce $u \equiv 1/\tau^{2/3}$

 $\varepsilon(u) = \sum_{n=2}^{242} \varepsilon_n u^n$

Convergence

 $\varepsilon(\tau) = \sum_{n=2}^{242} \frac{\varepsilon_n}{\tau^{\frac{2n}{3}}}$

By iteratively solving numerically the linear ODE's from fluid/gravity duality, we obtained 240 coefficients in the gradient expansion

• chief complication – generate the r.h.s. of the equations

• to get to so high orders we need very high precision computations

• first couple of orders - easy and fast

• Introduce $u \equiv 1/\tau^{2/3}$

 $\varepsilon(u) = \sum_{n=2}^{242} \varepsilon_n u^n$

Convergence

 $\varepsilon(\tau) = \sum_{n=2}^{242} \frac{\varepsilon_n}{\tau^{\frac{2n}{3}}}$

- By iteratively solving numerically the linear ODE's from fluid/gravity duality, we obtained 240 coefficients in the gradient expansion
 - chief complication generate the r.h.s. of the equations
 - to get to so high orders we need very high precision computations

first couple of orders – easy and fast

• Introduce $u \equiv 1/\tau^{2/3}$

 $\varepsilon(u) = \sum_{n=2}^{242} \varepsilon_n u^n$

Convergence

 $\varepsilon(\tau) = \sum_{n=2}^{242} \frac{\varepsilon_n}{\tau^{\frac{2n}{3}}}$

▶ Introduce $u \equiv 1/\tau^{2/3}$

By iteratively solving numerically the linear ODE's from fluid/gravity duality, we obtained 240 coefficients in the gradient expansion

• chief complication – generate the r.h.s. of the equations

- to get to so high orders we need very high precision computations
- first couple of orders easy and fast

$$\varepsilon(u) = \sum_{n=2}^{242} \varepsilon_n u^n$$

Convergence

 $\varepsilon(\tau) = \sum_{n=2}^{242} \frac{\varepsilon_n}{\tau^{\frac{2n}{3}}}$

By iteratively solving numerically the linear ODE's from fluid/gravity duality, we obtained 240 coefficients in the gradient expansion

• chief complication – generate the r.h.s. of the equations

- to get to so high orders we need very high precision computations
- first couple of orders easy and fast

• Introduce $u \equiv 1/\tau^{2/3}$

 $\varepsilon(u) = \sum_{n=2}^{242} \varepsilon_n u^n$

Convergence

 $\varepsilon(\tau) = \sum_{n=2}^{242} \frac{\varepsilon_n}{\tau^{\frac{2n}{3}}}$

By iteratively solving numerically the linear ODE's from fluid/gravity duality, we obtained 240 coefficients in the gradient expansion

• chief complication – generate the r.h.s. of the equations

- to get to so high orders we need very high precision computations
- first couple of orders easy and fast

• Introduce $u \equiv 1/\tau^{2/3}$

 $\varepsilon(u) = \sum_{n=2}^{242} \varepsilon_n u^n$

Convergence

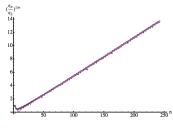
Large τ behaviour of $\varepsilon(\tau)$

 $\varepsilon(\tau) = \sum_{n=2}^{242} \frac{\varepsilon_n}{\tau^{\frac{2n}{3}}}$

- By iteratively solving numerically the linear ODE's from fluid/gravity duality, we obtained 240 coefficients in the gradient expansion
 - chief complication generate the r.h.s. of the equations
 - to get to so high orders we need very high precision computations
 - first couple of orders easy and fast

• Introduce $u \equiv 1/\tau^{2/3}$

$$\varepsilon(u) = \sum_{n=2}^{242} \varepsilon_n u^n$$



Zero radius of convergence Asymptotic series...

Convergence

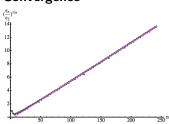
Large τ behaviour of $\varepsilon(\tau)$

 $\varepsilon(\tau) = \sum_{n=2}^{242} \frac{\varepsilon_n}{\tau^{\frac{2n}{3}}}$

- By iteratively solving numerically the linear ODE's from fluid/gravity duality, we obtained 240 coefficients in the gradient expansion
 - chief complication generate the r.h.s. of the equations
 - to get to so high orders we need very high precision computations
 - first couple of orders easy and fast

• Introduce $u \equiv 1/\tau^{2/3}$

$$\varepsilon(u) = \sum_{n=2}^{242} \varepsilon_n u^n$$



Zero radius of convergence

Asymptotic series...

Convergence

Large τ behaviour of $\varepsilon(\tau)$

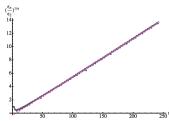
 $\varepsilon(\tau) = \sum_{n=2}^{242} \frac{\varepsilon_n}{\tau^{\frac{2n}{3}}}$

Convergence

- By iteratively solving numerically the linear ODE's from fluid/gravity duality, we obtained 240 coefficients in the gradient expansion
 - chief complication generate the r.h.s. of the equations
 - to get to so high orders we need very high precision computations
 - first couple of orders easy and fast

• Introduce $u \equiv 1/\tau^{2/3}$

$$\varepsilon(u) = \sum_{n=2}^{242} \varepsilon_n u^n$$



Zero radius of convergence Asymptotic series...

Define the Borel transform

$$\tilde{\varepsilon}(u) = \sum_{n=2}^{242} \frac{\varepsilon_n}{n!} u^n$$

$$arepsilon_{resum}(u) = \int_0^\infty e^{-s} \widetilde{arepsilon}(su) \, ds \qquad ext{where } u = au^{-rac{2}{3}}$$

- Look at poles ≡ zeroes of the denominator of the Pade approximant...

Define the Borel transform

$$\tilde{\varepsilon}(u) = \sum_{n=2}^{242} \frac{\varepsilon_n}{n!} u^n$$

$$arepsilon_{resum}(u) = \int_0^\infty e^{-s} \widetilde{arepsilon}(su) \, ds \qquad ext{where } u = au^{-rac{2}{3}}$$

- Look at poles ≡ zeroes of the denominator of the Pade approximant...

Define the Borel transform

$$\tilde{\varepsilon}(u) = \sum_{n=2}^{242} \frac{\varepsilon_n}{n!} u^n$$

$$arepsilon_{resum}(u) = \int_0^\infty e^{-s} \widetilde{arepsilon}(su) \, ds \qquad ext{where } u = au^{-rac{2}{3}}$$

- Look at poles ≡ zeroes of the denominator of the Pade approximant...

Define the Borel transform

$$\tilde{\varepsilon}(u) = \sum_{n=2}^{242} \frac{\varepsilon_n}{n!} u^n$$

$$arepsilon_{resum}(u) = \int_0^\infty e^{-s} \widetilde{arepsilon}(su) \, ds \qquad ext{where } u = au^{-rac{2}{3}}$$

- *ε̃*(u) has only a finite radius of convergence. In order to locate singularities in the Borel plane, we perform a symmetric Pade approximation...
- Look at poles ≡ zeroes of the denominator of the Pade approximant...

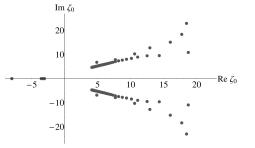
Define the Borel transform

$$\tilde{\varepsilon}(u) = \sum_{n=2}^{242} \frac{\varepsilon_n}{n!} u^n$$

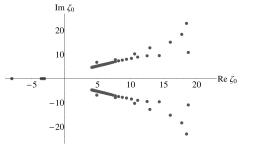
$$\varepsilon_{resum}(u) = \int_0^\infty e^{-s} \tilde{\varepsilon}(su) \, ds \qquad ext{where } u = \tau^{-rac{2}{3}}$$

- *ε̃*(u) has only a finite radius of convergence. In order to locate singularities in the Borel plane, we perform a symmetric Pade approximation...
- Look at poles ≡ zeroes of the denominator of the Pade approximant...

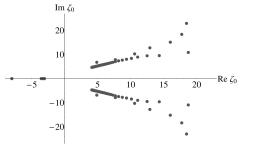
- The poles on the negative real axis are spurious
- The zeroes of the denominator of the Pade approximant coincide with the zeroes of the numerator up to 10⁻¹⁰⁰ accuracy



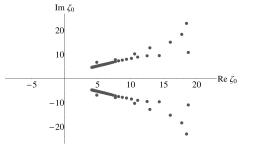
- The poles on the negative real axis are spurious
- The zeroes of the denominator of the Pade approximant coincide with the zeroes of the numerator up to 10⁻¹⁰⁰ accuracy



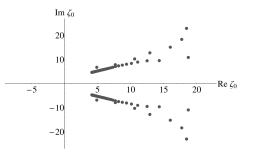
- The poles on the negative real axis are spurious
- The zeroes of the denominator of the Pade approximant coincide with the zeroes of the numerator up to 10⁻¹⁰⁰ accuracy



- The poles on the negative real axis are spurious
- The zeroes of the denominator of the Pade approximant coincide with the zeroes of the numerator up to 10⁻¹⁰⁰ accuracy

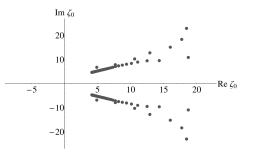


- The poles on the negative real axis are spurious
- The zeroes of the denominator of the Pade approximant coincide with the zeroes of the numerator up to 10⁻¹⁰⁰ accuracy



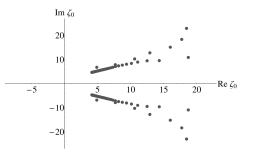
Question:

- Branch cuts on the Borel plane
- Branch points set the radius of convergence of the Borel transform
- Apparently no poles on the real axis!
 Borel resummation should be possible...



Question:

- Branch cuts on the Borel plane
- Branch points set the radius of convergence of the Borel transform
- Apparently no poles on the real axis!
 Borel resummation should be possible...

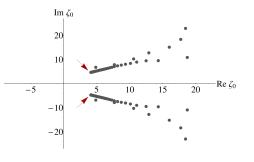


Question:

What is the physical interpretation of the branch cut singularities?

Branch cuts on the Borel plane

- Branch points set the radius of convergence of the Borel transform
- Apparently no poles on the real axis!
 Borel resummation should be possible...

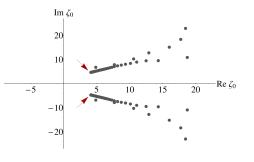


Branch cuts on the Borel plane

 Branch points set the radius of convergence of the Borel transform

 Apparently no poles on the real axis!
 Borel resummation should be possible...

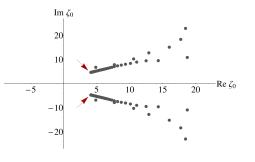
Question:



Branch cuts on the Borel plane

- Branch points set the radius of convergence of the Borel transform
- Apparently no poles on the real axis!
 Borel resummation should be possible...

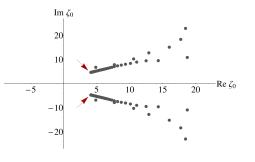
Question:



Branch cuts on the Borel plane

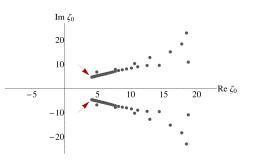
- Branch points set the radius of convergence of the Borel transform
- Apparently no poles on the real axis!
 Borel resummation should be possible...

Question:



- Branch cuts on the Borel plane
- Branch points set the radius of convergence of the Borel transform
- Apparently no poles on the real axis!
 Borel resummation should be possible...

Question:



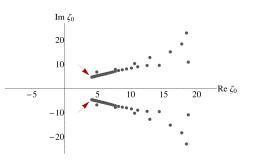
 Deform the contour of the inverse Borel transform

 $arepsilon_{resum}(au) = \int_{0}^{\infty} e^{-\zeta} \widetilde{arepsilon} \left(\zeta/ au^{2}_{3}
ight) \, d\zeta$

• The pole at the edge of the cut $(\zeta_0 = 4.12065 + 4.67895 i)$ will contribute as

- This is exactly the first lowest non-hydrodynamic quasi-normal mode!
- It is simply related to the scalar QNM of the planar black hole through
 RJ, Peschanski

$$-i\underbrace{(3.1195 - 2.7467 i)}_{planar BH QNM} \int \underbrace{\pi T(\tau)}_{1/\tau^{\frac{1}{3}}} d\tau = \underbrace{-i\frac{3}{2}(3.1195 - 2.7467 i)}_{-4.12005 - 4.67925 i} \tau^{\frac{2}{3}}$$



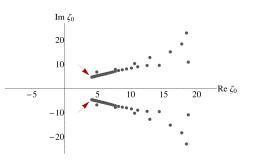
 Deform the contour of the inverse Borel transform

 $arepsilon_{resum}(au) = \int_{0}^{\infty} e^{-\zeta} \widetilde{arepsilon} \left(\zeta/ au^{2}_{3}
ight) \, d\zeta$

• The pole at the edge of the cut $(\zeta_0 = 4.12065 + 4.67895 i)$ will contribute as

- This is exactly the first lowest non-hydrodynamic quasi-normal mode!
- It is simply related to the scalar QNM of the planar black hole through
 RJ, Peschanski

$$-i\underbrace{(3.1195 - 2.7467 i)}_{planar BH QNM} \int \underbrace{\pi T(\tau)}_{1/\tau^{\frac{1}{3}}} d\tau = \underbrace{-i\frac{3}{2}(3.1195 - 2.7467 i)}_{-4.12005 - 4.67925 i} \tau^{\frac{2}{3}}$$



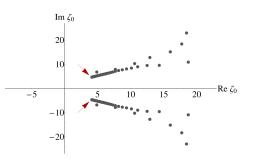
 Deform the contour of the inverse Borel transform

 $arepsilon_{resum}(au) = \int_{0}^{\infty} e^{-\zeta} \widetilde{arepsilon} \left(\zeta/ au^{rac{2}{3}}
ight) \, d\zeta$

• The pole at the edge of the cut $(\zeta_0 = 4.12065 + 4.67895 i)$ will contribute as

- This is exactly the first lowest non-hydrodynamic quasi-normal mode!
- It is simply related to the scalar QNM of the planar black hole through
 RJ, Peschanski

$$-i\underbrace{(3.1195 - 2.7467 i)}_{planar BH QNM} \int \underbrace{\pi T(\tau)}_{1/\tau^{\frac{1}{3}}} d\tau = \underbrace{-i\frac{3}{2}(3.1195 - 2.7467 i)}_{-4.12005 - 4.67925 i} \tau^{\frac{2}{3}}$$



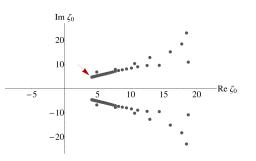
 Deform the contour of the inverse Borel transform

$$arepsilon_{resum}(au) = \int_{0}^{\infty} e^{-\zeta} \widetilde{arepsilon} \left(\zeta/ au^{rac{2}{3}}
ight) \, d\zeta$$

• The pole at the edge of the cut $(\zeta_0 = 4.12065 + 4.67895 i)$ will contribute as

- This is exactly the first lowest non-hydrodynamic quasi-normal mode!
- It is simply related to the scalar QNM of the planar black hole through
 RJ, Peschanski

$$-i\underbrace{(3.1195 - 2.7467 i)}_{planar BH QNM} \int \underbrace{\pi T(\tau)}_{1/\tau^{\frac{1}{3}}} d\tau = \underbrace{-i\frac{3}{2}(3.1195 - 2.7467 i)}_{-4.12005 - 4.67925 i} \tau^{\frac{2}{3}}$$



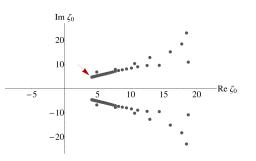
 Deform the contour of the inverse Borel transform

$$arepsilon_{\textit{resum}}(au) = \int_{0}^{\infty} e^{-\zeta} \widetilde{arepsilon} \left(\zeta/ au^{rac{2}{3}}
ight) \, d\zeta$$

► The pole at the edge of the cut (ζ₀ = 4.12065 + 4.67895 i) will contribute as

- This is exactly the first lowest non-hydrodynamic quasi-normal mode!
- It is simply related to the scalar QNM of the planar black hole through
 RJ, Peschanski

$$-i\underbrace{(3.1195 - 2.7467 i)}_{planar BH QNM} \int \underbrace{\pi T(\tau)}_{1/\tau^{\frac{1}{3}}} d\tau = \underbrace{-i\frac{3}{2}(3.1195 - 2.7467 i)}_{-4.12005 - 4.67925 i} \tau^{\frac{2}{3}}$$



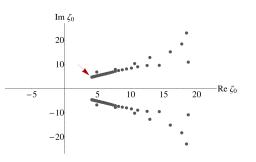
 Deform the contour of the inverse Borel transform

$$arepsilon_{\textit{resum}}(au) = \int_{0}^{\infty} e^{-\zeta} \widetilde{arepsilon} \left(\zeta/ au^{rac{2}{3}}
ight) \, d\zeta$$

► The pole at the edge of the cut (ζ₀ = 4.12065 + 4.67895 i) will contribute as

- This is exactly the first lowest non-hydrodynamic quasi-normal mode!
- It is simply related to the scalar QNM of the planar black hole through
 RJ, Peschanski

$$-i\underbrace{(3.1195 - 2.7467 i)}_{planar BH QNM} \int \underbrace{\pi T(\tau)}_{1/\tau^{\frac{1}{3}}} d\tau = \underbrace{-i\frac{3}{2}(3.1195 - 2.7467 i)}_{-4.12005 - 4.67925 i} \tau^{\frac{2}{3}}$$



 Deform the contour of the inverse Borel transform

$$arepsilon_{\textit{resum}}(au) = \int_{0}^{\infty} e^{-\zeta} \widetilde{arepsilon} \left(\zeta/ au^{rac{2}{3}}
ight) \, d\zeta$$

► The pole at the edge of the cut (ζ₀ = 4.12065 + 4.67895 i) will contribute as

- This is exactly the first lowest non-hydrodynamic quasi-normal mode!
- It is simply related to the scalar QNM of the planar black hole through
 RJ, Peschanski

$$-i\underbrace{(3.1195 - 2.7467 i)}_{planar \ BH \ QNM} \int \underbrace{\pi T(\tau)}_{1/\tau^{\frac{1}{3}}} d\tau = \underbrace{-i\frac{3}{2}(3.1195 - 2.7467 i)}_{-4.12005 - 4.67925 i} \tau^{\frac{2}{3}}$$

What is the interpretation of the whole branch cut?

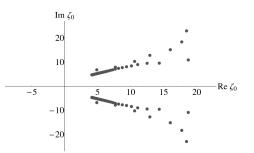
- Deform the contour of the inverse Borel transform to encircle the cut and extract the large \(\tau\) behaviour
- We obtain a preexponential power law factor

 $\tau^{-1.5426+0.5192 i} \cdot e^{-i \frac{3}{2}(3.1193-2.7471 i)\tau^{\frac{2}{3}}}$

- The late time geometry is an evolving black hole deformed by viscous corrections (first gradient terms in the fluid/gravity duality)
- This modification of the geometry generates a nontrivial power-law modification of the quasi-normal mode which is

```
\tau^{-1.5422+0.5199 \, i} \cdot e^{-i \frac{3}{2}(3.1195-2.7467 \, i)\tau^{\frac{4}{3}}}
```

What is the interpretation of the whole branch cut?



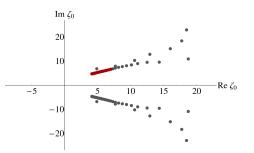
- Deform the contour of the inverse Borel transform to encircle the cut and extract the large \(\tau\) behaviour
- We obtain a preexponential power law factor

 $\tau^{-1.5426+0.5192\,i}$, $e^{-i\frac{3}{2}(3.1193-2.7471\,i)\tau^{\frac{2}{3}}}$

The late time geometry is an evolving black hole deformed by viscous corrections (first gradient terms in the fluid/gravity duality)
 This modification of the geometry generates a nontrivial power-law modification of the quasi-normal mode which is

 $\tau^{-1.5422+0.5199 \, i} \cdot e^{-i \frac{3}{2} (3.1195-2.7467 \, i) \tau^{\frac{4}{3}}}$

What is the interpretation of the whole branch cut?



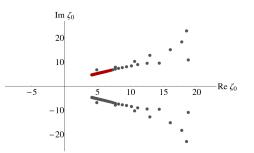
- Deform the contour of the inverse Borel transform to encircle the cut and extract the large \(\tau\) behaviour
- We obtain a preexponential power law factor

 $\tau^{-1.5426+0.5192 i}$, $-i\frac{3}{2}(3.1193-2.7471 i)\tau^{\frac{2}{3}}$

The late time geometry is an evolving black hole deformed by viscous corrections (first gradient terms in the fluid/gravity duality)
 This modification of the geometry generates a nontrivial power-law modification of the quasi-normal mode which is

 $\tau^{-1.5422+0.5199 \, i} \cdot e^{-i \frac{3}{2}(3.1195-2.7467 \, i)\tau^{\frac{5}{3}}}$

What is the interpretation of the whole branch cut?



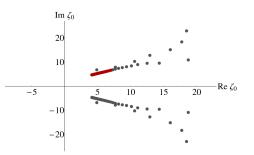
- Deform the contour of the inverse Borel transform to encircle the cut and extract the large \(\tau\) behaviour
- We obtain a preexponential power law factor

 $\tau^{-1.5426+0.5192\,i} \cdot e^{-i\frac{3}{2}(3.1193-2.7471\,i)\tau^{\frac{2}{3}}}$

The late time geometry is an evolving black hole deformed by viscous corrections (first gradient terms in the fluid/gravity duality)
 This modification of the geometry generates a nontrivial power-law modification of the quasi-normal mode which is

 $\tau^{-1.5422+0.5199 i} \cdot e^{-i\frac{3}{2}(3.1195-2.7467 i)\tau^{\frac{5}{3}}}$

What is the interpretation of the whole branch cut?



- Deform the contour of the inverse Borel transform to encircle the cut and extract the large \(\tau\) behaviour
- We obtain a preexponential power law factor

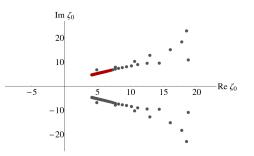
 $\tau^{-1.5426+0.5192\,i} \cdot e^{-i\frac{3}{2}(3.1193-2.7471\,i)\tau^{\frac{2}{3}}}$

 The late time geometry is an evolving black hole deformed by viscous corrections (first gradient terms in the fluid/gravity duality)

This modification of the geometry generates a nontrivial power-law modification of the quasi-normal mode which is

 $\tau^{-1.5422+0.5199\,i} \cdot e^{-i\frac{3}{2}(3.1195-2.7467\,i)\tau^{\frac{5}{3}}}$

What is the interpretation of the whole branch cut?



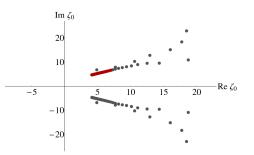
- Deform the contour of the inverse Borel transform to encircle the cut and extract the large \(\tau\) behaviour
- We obtain a preexponential power law factor

 $\tau^{-1.5426+0.5192\,i} \cdot e^{-i\frac{3}{2}(3.1193-2.7471\,i)\tau^{\frac{2}{3}}}$

- The late time geometry is an evolving black hole deformed by viscous corrections (first gradient terms in the fluid/gravity duality)
- This modification of the geometry generates a nontrivial power-law modification of the quasi-normal mode which is

```
\tau^{-1.5422+0.5199\,i} \cdot e^{-i\frac{3}{2}(3.1195-2.7467\,i)\tau^{\frac{2}{3}}}
```

What is the interpretation of the whole branch cut?



- Deform the contour of the inverse Borel transform to encircle the cut and extract the large \(\tau\) behaviour
- We obtain a preexponential power law factor

 $\tau^{-1.5426+0.5192 i} \cdot e^{-i\frac{3}{2}(3.1193-2.7471 i)\tau^{\frac{2}{3}}}$

- The late time geometry is an evolving black hole deformed by viscous corrections (first gradient terms in the fluid/gravity duality)
- This modification of the geometry generates a nontrivial power-law modification of the quasi-normal mode which is

```
\tau^{-1.5422+0.5199\,i} \cdot e^{-i\frac{3}{2}(3.1195-2.7467\,i)\tau^{\frac{2}{3}}}
```

- In fact the quasinormal mode preexponential factor has a clear interpretation
- Recall the hydrodynamic expression for the temperature $T(\tau)$:

$$\pi T(\tau) = \frac{1}{\tau^{\frac{1}{3}}} \left(1 - \frac{1}{6\tau^{\frac{2}{3}}} + \ldots \right)$$

• Then using the substitution $\pi Tt \longrightarrow \int \pi T(\tau) d\tau$ we get

$$\int \pi T(\tau) d\tau \sim \frac{3}{2}\tau^{\frac{2}{3}} - \frac{1}{6}\log \tau + \dots$$

So the QNM including viscous corrections which we derived earlier follows from the simple formula

$$\tau^{-2}e^{-i\omega_{QNM}\int \pi T(\tau)d\tau}$$

as

$$-2 - i\underbrace{(3.1195 - 2.7467i)}_{\omega_{QNM}} \cdot \left(-\frac{1}{6}\right) = -1.54222 + 0.519917i$$

- In fact the quasinormal mode preexponential factor has a clear interpretation
- Recall the hydrodynamic expression for the temperature $T(\tau)$:

$$\pi T(\tau) = \frac{1}{\tau^{\frac{1}{3}}} \left(1 - \frac{1}{6\tau^{\frac{2}{3}}} + \ldots \right)$$

▶ Then using the substitution $\pi Tt \longrightarrow \int \pi T(\tau) d\tau$ we get

$$\int \pi T(\tau) d\tau \sim \frac{3}{2}\tau^{\frac{2}{3}} - \frac{1}{6}\log\tau + \dots$$

So the QNM including viscous corrections which we derived earlier follows from the simple formula

$$\tau^{-2}e^{-i\omega_{QNM}\int \pi T(\tau)d\tau}$$

as

$$-2 - i\underbrace{(3.1195 - 2.7467i)}_{\omega_{QNM}} \cdot \left(-\frac{1}{6}\right) = -1.54222 + 0.519917i$$

- In fact the quasinormal mode preexponential factor has a clear interpretation
- Recall the hydrodynamic expression for the temperature $T(\tau)$:

$$\pi T(\tau) = rac{1}{ au^{rac{1}{3}}} \left(1 - rac{1}{6 au^{rac{2}{3}}} + \ldots
ight)$$

▶ Then using the substitution $\pi Tt \longrightarrow \int \pi T(\tau) d\tau$ we get

$$\int \pi T(\tau) d\tau \sim \frac{3}{2}\tau^{\frac{2}{3}} - \frac{1}{6}\log \tau + \dots$$

So the QNM including viscous corrections which we derived earlier follows from the simple formula

$$\tau^{-2}e^{-i\omega_{QNM}\int \pi T(\tau)d\tau}$$

$$-2 - i\underbrace{(3.1195 - 2.7467i)}_{\omega_{QNM}} \cdot \left(-\frac{1}{6}\right) = -1.54222 + 0.519917i$$

- In fact the quasinormal mode preexponential factor has a clear interpretation
- Recall the hydrodynamic expression for the temperature $T(\tau)$:

$$\pi T(\tau) = \frac{1}{\tau^{\frac{1}{3}}} \left(1 - \frac{1}{6\tau^{\frac{2}{3}}} + \ldots \right)$$

• Then using the substitution $\pi Tt \longrightarrow \int \pi T(\tau) d\tau$ we get

$$\int \pi T(\tau) d\tau \sim \frac{3}{2}\tau^{\frac{2}{3}} - \frac{1}{6}\log \tau + \dots$$

So the QNM including viscous corrections which we derived earlier follows from the simple formula

$$\tau^{-2} e^{-i\omega_{QNM}\int \pi T(\tau)d\tau}$$

$$-2 - i\underbrace{(3.1195 - 2.7467i)}_{\omega_{QNM}} \cdot \left(-\frac{1}{6}\right) = -1.54222 + 0.519917i$$

- In fact the quasinormal mode preexponential factor has a clear interpretation
- Recall the hydrodynamic expression for the temperature $T(\tau)$:

$$\pi T(\tau) = rac{1}{ au^{rac{1}{3}}} \left(1 - rac{1}{6 au^{rac{2}{3}}} + \ldots
ight)$$

▶ Then using the substitution $\pi Tt \longrightarrow \int \pi T(\tau) d\tau$ we get

$$\int \pi T(\tau) d\tau \sim \frac{3}{2}\tau^{\frac{2}{3}} - \frac{1}{6}\log\tau + \dots$$

 So the QNM including viscous corrections which we derived earlier follows from the simple formula

$$\tau^{-2} e^{-i\omega_{QNM}\int \pi T(\tau)d\tau}$$

$$-2 - i\underbrace{(3.1195 - 2.7467i)}_{\omega_{QNM}} \cdot \left(-\frac{1}{6}\right) = -1.54222 + 0.519917i$$

- In fact the quasinormal mode preexponential factor has a clear interpretation
- Recall the hydrodynamic expression for the temperature $T(\tau)$:

$$\pi T(\tau) = \frac{1}{\tau^{\frac{1}{3}}} \left(1 - \frac{1}{6\tau^{\frac{2}{3}}} + \ldots \right)$$

• Then using the substitution $\pi Tt \longrightarrow \int \pi T(\tau) d\tau$ we get

$$\int \pi T(\tau) d\tau \sim \frac{3}{2} \tau^{\frac{2}{3}} - \frac{1}{6} \log \tau + \dots$$

 So the QNM including viscous corrections which we derived earlier follows from the simple formula

$$\tau^{-2} e^{-i\omega_{QNM}\int \pi T(\tau)d\tau}$$

$$-2 - i\underbrace{(3.1195 - 2.7467i)}_{\omega_{QNM}} \cdot \left(-\frac{1}{6}\right) = -1.54222 + 0.519917i$$

- The singularities in the Borel plane have a clear physical origin they correspond to the lowest non-hydrodynamic modes/degrees of freedom
- All-order hydrodynamics knows about its UV completion!

No singularities on the positive real axis!

- The singularities in the Borel plane have a clear physical origin they correspond to the lowest non-hydrodynamic modes/degrees of freedom
- All-order hydrodynamics knows about its UV completion!

No singularities on the positive real axis!

- The singularities in the Borel plane have a clear physical origin they correspond to the lowest non-hydrodynamic modes/degrees of freedom
- All-order hydrodynamics knows about its UV completion!

No singularities on the positive real axis!

- The singularities in the Borel plane have a clear physical origin they correspond to the lowest non-hydrodynamic modes/degrees of freedom
- All-order hydrodynamics knows about its UV completion!

No singularities on the positive real axis!

- The singularities in the Borel plane have a clear physical origin they correspond to the lowest non-hydrodynamic modes/degrees of freedom
- All-order hydrodynamics knows about its UV completion!

No singularities on the positive real axis!

We need to evaluate the integral

$$arepsilon_{resum}(au) = \int_{0}^{\infty} e^{-\zeta} \widetilde{arepsilon} \left(\zeta/ au^{rac{2}{3}}
ight) \, d\zeta$$

- For this we need to have an analytical continuation of the Borel transform *ε̃*(..) on the *whole* positive real axis → use Pade approximant
- What type of Pade??
- ▶ Naively (n, n) would be natural as it would lead to $\varepsilon(\tau) \rightarrow const$ as $\tau \rightarrow 0$
- ► However the hydrodynamic series suggests different asymptotics:

$$u \frac{\partial_u \tilde{\varepsilon}(u)}{\tilde{\varepsilon}(u)} \longrightarrow (n, n) \text{ Pade } \longrightarrow \lim_{u \to \infty} u \frac{\partial_u \tilde{\varepsilon}(u)}{\tilde{\varepsilon}(u)} = 3.00000022...$$

We need to evaluate the integral

$$arepsilon_{\textit{resum}}(au) = \int_{0}^{\infty} e^{-\zeta} \widetilde{arepsilon} \left(\zeta/ au^{2}_{3}
ight) \, d\zeta$$

- For this we need to have an analytical continuation of the Borel transform *ε̃*(..) on the *whole* positive real axis → use Pade approximant
- What type of Pade??
- ▶ Naively (n, n) would be natural as it would lead to $\varepsilon(\tau) \rightarrow const$ as $\tau \rightarrow 0$
- However the hydrodynamic series suggests different asymptotics:

$$u \frac{\partial_u \tilde{\varepsilon}(u)}{\tilde{\varepsilon}(u)} \longrightarrow (n, n) \text{ Pade } \longrightarrow \lim_{u \to \infty} u \frac{\partial_u \tilde{\varepsilon}(u)}{\tilde{\varepsilon}(u)} = 3.00000022...$$

We need to evaluate the integral

$$arepsilon_{resum}(au) = \int_{0}^{\infty} e^{-\zeta} \widetilde{arepsilon} \left(\zeta/ au^{2}_{3}
ight) \, d\zeta$$

For this we need to have an analytical continuation of the Borel transform *ε̃*(..) on the *whole* positive real axis → use Pade approximant

- What type of Pade??
- ▶ Naively (n, n) would be natural as it would lead to $\varepsilon(\tau) \rightarrow const$ as $\tau \rightarrow 0$
- ► However the hydrodynamic series suggests different asymptotics:

$$u \frac{\partial_u \widetilde{\varepsilon}(u)}{\widetilde{\varepsilon}(u)} \longrightarrow (n, n) \text{ Pade } \longrightarrow \lim_{u \to \infty} u \frac{\partial_u \widetilde{\varepsilon}(u)}{\widetilde{\varepsilon}(u)} = 3.00000022...$$

We need to evaluate the integral

$$arepsilon_{resum}(au) = \int_{0}^{\infty} e^{-\zeta} \widetilde{arepsilon} \left(\zeta/ au^{2}_{3}
ight) \, d\zeta$$

- For this we need to have an analytical continuation of the Borel transform *ε̃*(..) on the *whole* positive real axis → use Pade approximant
- What type of Pade??
- ▶ Naively (n, n) would be natural as it would lead to $\varepsilon(\tau) \rightarrow const$ as $\tau \rightarrow 0$
- However the hydrodynamic series suggests different asymptotics:

$$u \frac{\partial_u \widetilde{\varepsilon}(u)}{\widetilde{\varepsilon}(u)} \longrightarrow (n, n) \text{ Pade } \longrightarrow \lim_{u \to \infty} u \frac{\partial_u \widetilde{\varepsilon}(u)}{\widetilde{\varepsilon}(u)} = 3.00000022...$$

We need to evaluate the integral

$$arepsilon_{resum}(au) = \int_{0}^{\infty} e^{-\zeta} \widetilde{arepsilon} \left(\zeta/ au^{2}_{3}
ight) \, d\zeta$$

- For this we need to have an analytical continuation of the Borel transform *ε̃*(..) on the *whole* positive real axis → use Pade approximant
- What type of Pade??
- ▶ Naively (n, n) would be natural as it would lead to $\varepsilon(\tau) \rightarrow const$ as $\tau \rightarrow 0$
- However the hydrodynamic series suggests different asymptotics:

$$u \frac{\partial_u \widetilde{\varepsilon}(u)}{\widetilde{\varepsilon}(u)} \longrightarrow (n, n) \text{ Pade } \longrightarrow \lim_{u \to \infty} u \frac{\partial_u \widetilde{\varepsilon}(u)}{\widetilde{\varepsilon}(u)} = 3.00000022...$$

▶ We need to evaluate the integral

$$arepsilon_{resum}(au) = \int_{0}^{\infty} e^{-\zeta} \widetilde{arepsilon} \left(\zeta/ au^{2}_{3}
ight) \, d\zeta$$

- For this we need to have an analytical continuation of the Borel transform *ε̃*(..) on the *whole* positive real axis → use Pade approximant
- What type of Pade??
- ▶ Naively (n, n) would be natural as it would lead to $\varepsilon(\tau) \to const$ as $\tau \to 0$
- However the hydrodynamic series suggests different asymptotics:

$$u \frac{\partial_u \widetilde{\varepsilon}(u)}{\widetilde{\varepsilon}(u)} \longrightarrow (n, n) \text{ Pade } \longrightarrow \lim_{u \to \infty} u \frac{\partial_u \widetilde{\varepsilon}(u)}{\widetilde{\varepsilon}(u)} = 3.00000022...$$

We need to evaluate the integral

$$arepsilon_{resum}(au) = \int_{0}^{\infty} e^{-\zeta} \widetilde{arepsilon} \left(\zeta/ au^{2}_{3}
ight) \, d\zeta$$

- For this we need to have an analytical continuation of the Borel transform *ε̃*(..) on the *whole* positive real axis → use Pade approximant
- What type of Pade??
- ▶ Naively (n, n) would be natural as it would lead to $\varepsilon(\tau) \rightarrow const$ as $\tau \rightarrow 0$
- However the hydrodynamic series suggests different asymptotics:

$$u rac{\partial_u \widetilde{\varepsilon}(u)}{\widetilde{\varepsilon}(u)} \longrightarrow (n, n) \text{ Pade } \longrightarrow \lim_{u \to \infty} u rac{\partial_u \widetilde{\varepsilon}(u)}{\widetilde{\varepsilon}(u)} = 3.00000022...$$

We need to evaluate the integral

$$arepsilon_{resum}(au) = \int_{0}^{\infty} e^{-\zeta} \widetilde{arepsilon} \left(\zeta/ au^{2}_{3}
ight) \, d\zeta$$

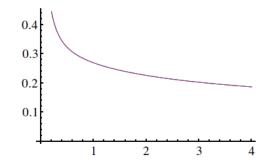
- For this we need to have an analytical continuation of the Borel transform *ε̃*(..) on the *whole* positive real axis → use Pade approximant
- What type of Pade??
- ▶ Naively (n, n) would be natural as it would lead to $\varepsilon(\tau) \rightarrow const$ as $\tau \rightarrow 0$
- However the hydrodynamic series suggests different asymptotics:

$$u \frac{\partial_u \widetilde{\varepsilon}(u)}{\widetilde{\varepsilon}(u)} \longrightarrow (n, n) \text{ Pade } \longrightarrow \lim_{u \to \infty} u \frac{\partial_u \widetilde{\varepsilon}(u)}{\widetilde{\varepsilon}(u)} = 3.00000022...$$

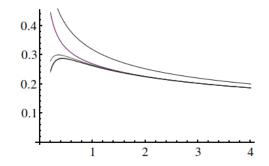
We constructed the Borel resummed $T(\tau)$ for $\tau > 0.2$ (units set by $T(\tau) \sim \frac{1}{\pi} \tau^{-\frac{1}{3}}$ as $\tau \to \infty$)

We constructed the Borel resummed $T(\tau)$ for $\tau > 0.2$ (units set by $T(\tau) \sim \frac{1}{\pi} \tau^{-\frac{1}{3}}$ as $\tau \to \infty$)

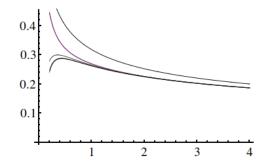
We constructed the Borel resummed $T(\tau)$ for $\tau > 0.2$ (units set by $T(\tau) \sim \frac{1}{\pi} \tau^{-\frac{1}{3}}$ as $\tau \to \infty$)



We constructed the Borel resummed $T(\tau)$ for $\tau > 0.2$ (units set by $T(\tau) \sim \frac{1}{\pi} \tau^{-\frac{1}{3}}$ as $\tau \to \infty$)



We constructed the Borel resummed $T(\tau)$ for $\tau > 0.2$ (units set by $T(\tau) \sim \frac{1}{\pi} \tau^{-\frac{1}{3}}$ as $\tau \to \infty$)



M. Heller, RJ, P, Witaszczyk, 1103.3452, 1203.0755

Describe the time dependent evolving boost-invariant strongly coupled plasma system

Describe it in terms of lightest degrees of freedom on the AdS side which are relevant at strong coupling

$$ds^{2} = \frac{g_{\mu\nu}(\tau, z)dx^{\mu}dx^{\nu} + dz^{2}}{z^{2}} \equiv g_{\alpha\beta}^{5D}dx^{\alpha}dx^{\beta}$$

$$\downarrow$$

Compute the time-evolution by solving (numerically) 5D Einstein's equations

$$R_{lphaeta} - rac{1}{2}g^{5D}_{lphaeta}R - 6\,g^{5D}_{lphaeta} = 0$$

M. Heller, RJ, P, Witaszczyk, 1103.3452, 1203.0755

Describe the time dependent evolving boost-invariant strongly coupled plasma system

Describe it in terms of lightest degrees of freedom on the AdS side which are relevant at strong coupling

$$ds^{2} = \frac{g_{\mu\nu}(\tau, z)dx^{\mu}dx^{\nu} + dz^{2}}{z^{2}} \equiv g_{\alpha\beta}^{5D}dx^{\alpha}dx^{\beta}$$

$$\downarrow$$

Compute the time-evolution by solving (numerically) 5D Einstein's equations

$$R_{lphaeta} - rac{1}{2}g^{5D}_{lphaeta}R - 6\,g^{5D}_{lphaeta} = 0$$

M. Heller, RJ, P, Witaszczyk, 1103.3452, 1203.0755

Describe the time dependent evolving boost-invariant strongly coupled plasma system

Describe it in terms of lightest degrees of freedom on the AdS side which are relevant at strong coupling

 \downarrow

$$ds^{2} = \frac{g_{\mu\nu}(\tau, z)dx^{\mu}dx^{\nu} + dz^{2}}{z^{2}} \equiv g_{\alpha\beta}^{5D}dx^{\alpha}dx^{\beta}$$

Compute the time-evolution by solving (numerically) 5D Einstein's equations

$$R_{lphaeta} - rac{1}{2} g^{5D}_{lphaeta} R - 6 \, g^{5D}_{lphaeta} = 0$$

M. Heller, RJ, P, Witaszczyk, 1103.3452, 1203.0755

Describe the time dependent evolving boost-invariant strongly coupled plasma system

Describe it in terms of lightest degrees of freedom on the AdS side which are relevant at strong coupling

 \downarrow

$$ds^{2} = \frac{g_{\mu\nu}(\tau, z)dx^{\mu}dx^{\nu} + dz^{2}}{z^{2}} \equiv g_{\alpha\beta}^{5D}dx^{\alpha}dx^{\beta}$$

Compute the time-evolution by solving (numerically) 5D Einstein's equations

$$R_{\alpha\beta} - \frac{1}{2}g^{5D}_{\alpha\beta}R - 6\,g^{5D}_{\alpha\beta} = 0$$

M. Heller, RJ, P, Witaszczyk, 1103.3452, 1203.0755

Describe the time dependent evolving boost-invariant strongly coupled plasma system

Describe it in terms of lightest degrees of freedom on the AdS side which are relevant at strong coupling

 \downarrow

$$ds^{2} = \frac{g_{\mu\nu}(\tau, z)dx^{\mu}dx^{\nu} + dz^{2}}{z^{2}} \equiv g_{\alpha\beta}^{5D}dx^{\alpha}dx^{\beta}$$

$$\downarrow$$

Compute the time-evolution by solving (numerically) 5D Einstein's equations

$$R_{lphaeta}-rac{1}{2}g^{5D}_{lphaeta}R-6\,g^{5D}_{lphaeta}=0$$

M. Heller, RJ, P, Witaszczyk, 1103.3452, 1203.0755

Describe the time dependent evolving boost-invariant strongly coupled plasma system

Describe it in terms of lightest degrees of freedom on the AdS side which are relevant at strong coupling

 \downarrow

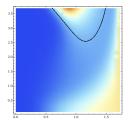
$$ds^{2} = \frac{g_{\mu\nu}(\tau, z)dx^{\mu}dx^{\nu} + dz^{2}}{z^{2}} \equiv g_{\alpha\beta}^{5D}dx^{\alpha}dx^{\beta}$$

$$\downarrow$$

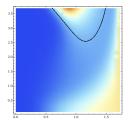
Compute the time-evolution by solving (numerically) 5D Einstein's equations

$$R_{lphaeta}-rac{1}{2}g^{5D}_{lphaeta}R-6\,g^{5D}_{lphaeta}=0$$

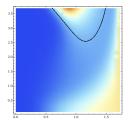
- Asymptotics of $g_{\mu\nu}(x^{\rho}, z)$ at $z \sim 0$ gives the energy-momentum tensor $T_{\mu\nu}(x^{\rho})$ of the plasma system
- We can test whether T_{μν}(x^ρ) is of a hydrodynamic form...
- We can check for local thermal equilibrium
- The area of the apparent horizon defines for us the entropy density
- We observe some *initial* entropy
- It is convenient to organize initial data according to their initial entropy



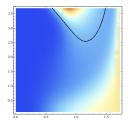
- Asymptotics of $g_{\mu\nu}(x^{\rho}, z)$ at $z \sim 0$ gives the energy-momentum tensor $T_{\mu\nu}(x^{\rho})$ of the plasma system
- We can test whether T_{μν}(x^ρ) is of a hydrodynamic form...
- We can check for local thermal equilibrium
- The area of the apparent horizon defines for us the entropy density
- We observe some *initial* entropy
- It is convenient to organize initial data according to their initial entropy



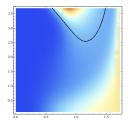
- Asymptotics of g_{µν}(x^ρ, z) at z ~ 0 gives the energy-momentum tensor T_{µν}(x^ρ) of the plasma system
- We can test whether T_{µν}(x^ρ) is of a hydrodynamic form...
- We can check for local thermal equilibrium
- The area of the apparent horizon defines for us the entropy density
- We observe some *initial* entropy
- It is convenient to organize initial data according to their initial entropy



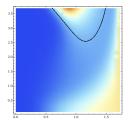
- Asymptotics of $g_{\mu\nu}(x^{\rho}, z)$ at $z \sim 0$ gives the energy-momentum tensor $T_{\mu\nu}(x^{\rho})$ of the plasma system
- We can test whether T_{µν}(x^ρ) is of a hydrodynamic form...
- We can check for local thermal equilibrium
- The area of the apparent horizon defines for us the entropy density
- We observe some *initial* entropy
- It is convenient to organize initial data according to their initial entropy



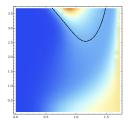
- Asymptotics of $g_{\mu\nu}(x^{\rho}, z)$ at $z \sim 0$ gives the energy-momentum tensor $T_{\mu\nu}(x^{\rho})$ of the plasma system
- We can test whether T_{µν}(x^ρ) is of a hydrodynamic form...
- We can check for local thermal equilibrium
- The area of the apparent horizon defines for us the entropy density
- We observe some *initial* entropy
- It is convenient to organize initial data according to their initial entropy

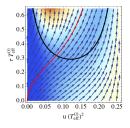


- Asymptotics of $g_{\mu\nu}(x^{\rho}, z)$ at $z \sim 0$ gives the energy-momentum tensor $T_{\mu\nu}(x^{\rho})$ of the plasma system
- We can test whether T_{µν}(x^ρ) is of a hydrodynamic form...
- We can check for local thermal equilibrium
- The area of the apparent horizon defines for us the entropy density
- We observe some *initial* entropy
- It is convenient to organize initial data according to their initial entropy



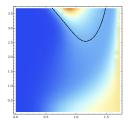
- Asymptotics of $g_{\mu\nu}(x^{\rho}, z)$ at $z \sim 0$ gives the energy-momentum tensor $T_{\mu\nu}(x^{\rho})$ of the plasma system
- We can test whether T_{µν}(x^ρ) is of a hydrodynamic form...
- We can check for local thermal equilibrium
- The area of the apparent horizon defines for us the entropy density
- We observe some *initial* entropy
- It is convenient to organize initial data according to their initial entropy

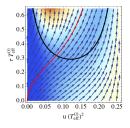




- Asymptotics of $g_{\mu\nu}(x^{\rho}, z)$ at $z \sim 0$ gives the energy-momentum tensor $T_{\mu\nu}(x^{\rho})$ of the plasma system
- We can test whether T_{µν}(x^ρ) is of a hydrodynamic form...
- We can check for local thermal equilibrium
- The area of the apparent horizon defines for us the entropy density
- We observe some *initial* entropy
- It is convenient to organize initial data according to their initial entropy

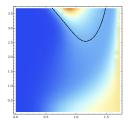
What physics can we extract?

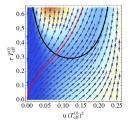




- Asymptotics of $g_{\mu\nu}(x^{\rho}, z)$ at $z \sim 0$ gives the energy-momentum tensor $T_{\mu\nu}(x^{\rho})$ of the plasma system
- We can test whether T_{µν}(x^ρ) is of a hydrodynamic form...
- We can check for local thermal equilibrium
- The area of the apparent horizon defines for us the entropy density
- We observe some *initial* entropy
- It is convenient to organize initial data according to their initial entropy

What physics can we extract?





- Asymptotics of g_{µν}(x^ρ, z) at z ~ 0 gives the energy-momentum tensor T_{µν}(x^ρ) of the plasma system
- We can test whether T_{µν}(x^ρ) is of a hydrodynamic form...
- We can check for local thermal equilibrium
- The area of the apparent horizon defines for us the entropy density
- We observe some *initial* entropy
- It is convenient to organize initial data according to their initial entropy

- As a measure of energy density introduce the effective temperature (≡ temperature of a thermal system with the same energy density)
- **2.** Form the dimensionless product $w \equiv T_{eff} \cdot \tau$
- 3. For all initial conditions considered, viscous hydrodynamics works very well for $w \equiv T_{eff} \cdot \tau > 0.7$

(natural values for RHIC: ($\tau_0 = 0.25 \text{ fm}$, $T_0 = 500 \text{ MeV}$) assumed in [Broniowski, Chojnacki, Florkowski, Kisiel] correspond to w = 0.63)

$$\Delta p_L \equiv 1 - \frac{p_L}{\varepsilon/3} \sim 0.7$$

- 1. As a measure of energy density introduce the effective temperature (\equiv temperature of a thermal system with the same energy density)
- **2.** Form the dimensionless product $w \equiv T_{eff} \cdot au$
- **3.** For all initial conditions considered, *viscous* hydrodynamics works very well for $w \equiv T_{eff} \cdot \tau > 0.7$

(natural values for RHIC: ($\tau_0 = 0.25 \text{ fm}$, $T_0 = 500 \text{ MeV}$) assumed in [Broniowski, Chojnacki, Florkowski, Kisiel] correspond to w = 0.63)

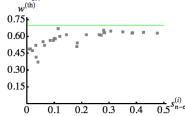
$$\Delta p_L \equiv 1 - \frac{p_L}{\varepsilon/3} \sim 0.7$$

- 1. As a measure of energy density introduce the effective temperature (\equiv temperature of a thermal system with the same energy density)
- **2.** Form the dimensionless product $w \equiv T_{eff} \cdot \tau$
- 3. For all initial conditions considered, viscous hydrodynamics works very well for $w \equiv T_{eff} \cdot \tau > 0.7$

(natural values for RHIC: ($\tau_0 = 0.25 \text{ fm}$, $T_0 = 500 \text{ MeV}$) assumed in [Broniowski, Chojnacki, Florkowski, Kisiel] correspond to w = 0.63)

$$\Delta p_L \equiv 1 - \frac{p_L}{\varepsilon/3} \sim 0.7$$

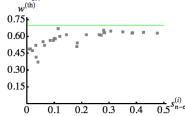
- 1. As a measure of energy density introduce the effective temperature (\equiv temperature of a thermal system with the same energy density)
- 2. Form the dimensionless product $w \equiv T_{eff} \cdot \tau$
- 3. For all initial conditions considered, viscous hydrodynamics works very well for $w \equiv T_{eff} \cdot \tau > 0.7$



(natural values for RHIC: ($\tau_0 = 0.25 \text{ fm}$, $T_0 = 500 \text{ MeV}$) assumed in [Broniowski, Chojnacki, Florkowski, Kisiel] correspond to w = 0.63)

$$\Delta p_L \equiv 1 - \frac{p_L}{\varepsilon/3} \sim 0.7$$

- 1. As a measure of energy density introduce the effective temperature (\equiv temperature of a thermal system with the same energy density)
- 2. Form the dimensionless product $w \equiv T_{eff} \cdot \tau$
- 3. For all initial conditions considered, viscous hydrodynamics works very well for $w \equiv T_{eff} \cdot \tau > 0.7$



(natural values for RHIC: ($\tau_0 = 0.25 \text{ fm}$, $T_0 = 500 \text{ MeV}$) assumed in [Broniowski, Chojnacki, Florkowski, Kisiel] correspond to w = 0.63)

$$\Delta p_L \equiv 1 - \frac{p_L}{\varepsilon/3} \sim 0.7$$

- 1. As a measure of energy density introduce the effective temperature (\equiv temperature of a thermal system with the same energy density)
- 2. Form the dimensionless product $w \equiv T_{eff} \cdot \tau$
- 3. For all initial conditions considered, viscous hydrodynamics works very well for $w \equiv T_{eff} \cdot \tau > 0.7$

$$\begin{array}{c} w^{(\text{in})} \\ 0.75 \\ 0.60 \\ 0.45 \\ 0.30 \\ 0.15 \\ 0 \\ 0 \\ 0.1 \\ 0 \\ 0.1 \\ 0 \\ 0.1 \\ 0.2 \\ 0.3 \\ 0.4 \\ 0.5 \\ s^{(i)}_{n-eq} \\ s^{(i)}_{n-$$

(natural values for RHIC: ($\tau_0 = 0.25 \text{ fm}$, $T_0 = 500 \text{ MeV}$) assumed in [Broniowski, Chojnacki, Florkowski, Kisiel] correspond to w = 0.63)

$$\Delta p_L \equiv 1 - \frac{p_L}{\varepsilon/3} \sim 0.7$$

- 1. As a measure of energy density introduce the effective temperature (\equiv temperature of a thermal system with the same energy density)
- 2. Form the dimensionless product $w \equiv T_{eff} \cdot \tau$
- 3. For all initial conditions considered, viscous hydrodynamics works very well for $w \equiv T_{eff} \cdot \tau > 0.7$

$$\begin{array}{c} w^{(h)} \\ 0.75 \\ 0.60 \\ 0.45 \\ 0.30 \\ 0.15 \\ 0 \\ 0 \\ 0.1 \\ 0.2 \\ 0.3 \\ 0.4 \\ 0.5 \\ s^{(i)}_{n-eq} \\ s$$

(natural values for RHIC: ($\tau_0 = 0.25 \text{ fm}$, $T_0 = 500 \text{ MeV}$) assumed in [Broniowski, Chojnacki, Florkowski, Kisiel] correspond to w = 0.63)

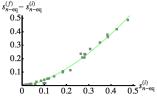
$$\Delta p_L \equiv 1 - \frac{p_L}{\varepsilon/3} \sim 0.7$$

Initial entropy turns out to be a key characterization of the initial state

1. There is a clear correlation of produced entropy with the initial entropy...

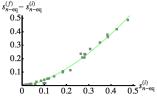
Initial entropy turns out to be a key characterization of the initial state

1. There is a clear correlation of produced entropy with the initial entropy...



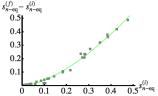
Initial entropy turns out to be a key characterization of the initial state

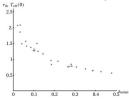
1. There is a clear correlation of produced entropy with the initial entropy...



Initial entropy turns out to be a key characterization of the initial state

1. There is a clear correlation of produced entropy with the initial entropy...





As mentioned earlier, we get rid of the dependence on the number of degrees of freedom by parametrizing the energy density through an effective temperature given by

$$\varepsilon(\tau) = \frac{3}{8} N_c^2 \pi^2 T_{eff}^4(\tau)$$

Previously, we normalized our initial data by setting

 $T_{eff}(\tau=0)=1$

but this is generically unknown in realistic heavy-ion collisions...

It is much better to fix the normalization through the hydrodynamic tail...

$$\pi T_{e\!f\!f}(au) \sim rac{1}{ au^{rac{1}{3}}} \qquad ext{in the } au o \infty ext{ limit}$$

The coefficient '1' fixes the units of τ .

As mentioned earlier, we get rid of the dependence on the number of degrees of freedom by parametrizing the energy density through an effective temperature given by

$$\varepsilon(\tau) = \frac{3}{8} N_c^2 \pi^2 T_{eff}^4(\tau)$$

Previously, we normalized our initial data by setting

 $T_{eff}(\tau=0)=1$

but this is generically unknown in realistic heavy-ion collisions...

It is much better to fix the normalization through the hydrodynamic tail...

$$\pi T_{e\!f\!f}(au) \sim rac{1}{ au^{rac{1}{3}}} \qquad ext{in the } au o \infty ext{ limit}$$

The coefficient '1' fixes the units of τ .

As mentioned earlier, we get rid of the dependence on the number of degrees of freedom by parametrizing the energy density through an effective temperature given by

$$\varepsilon(\tau) = \frac{3}{8} N_c^2 \pi^2 T_{eff}^4(\tau)$$

Previously, we normalized our initial data by setting

 $T_{eff}(\tau=0)=1$

but this is generically unknown in realistic heavy-ion collisions...

It is much better to fix the normalization through the hydrodynamic tail...

 $\pi T_{e\!f\!f}(au) \sim rac{1}{ au^{rac{1}{3}}} \qquad ext{in the } au o \infty ext{ limit}$

The coefficient '1' fixes the units of τ .

As mentioned earlier, we get rid of the dependence on the number of degrees of freedom by parametrizing the energy density through an effective temperature given by

$$\varepsilon(\tau) = \frac{3}{8} N_c^2 \pi^2 T_{eff}^4(\tau)$$

Previously, we normalized our initial data by setting

 $T_{eff}(\tau=0)=1$

but this is generically unknown in realistic heavy-ion collisions...

It is much better to fix the normalization through the hydrodynamic tail...

 $\pi T_{e\!f\!f}(au) \sim rac{1}{ au^{rac{1}{3}}} \qquad ext{in the } au o \infty ext{ limit}$

The coefficient '1' fixes the units of τ .

As mentioned earlier, we get rid of the dependence on the number of degrees of freedom by parametrizing the energy density through an effective temperature given by

$$\varepsilon(\tau) = \frac{3}{8} N_c^2 \pi^2 T_{eff}^4(\tau)$$

Previously, we normalized our initial data by setting

$$T_{eff}(\tau=0)=1$$

but this is generically unknown in realistic heavy-ion collisions...

It is much better to fix the normalization through the hydrodynamic tail...

$$\pi au_{e\!f\!f}(au) \sim rac{1}{ au^{rac{1}{3}}} \qquad ext{in the } au o \infty ext{ limit}$$

The coefficient '1' fixes the units of τ .

As mentioned earlier, we get rid of the dependence on the number of degrees of freedom by parametrizing the energy density through an effective temperature given by

$$\varepsilon(\tau) = \frac{3}{8} N_c^2 \pi^2 T_{eff}^4(\tau)$$

Previously, we normalized our initial data by setting

$$T_{eff}(\tau=0)=1$$

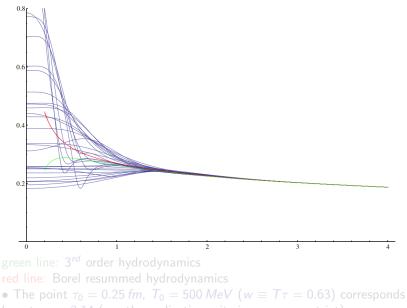
but this is generically unknown in realistic heavy-ion collisions...

It is much better to fix the normalization through the hydrodynamic tail...

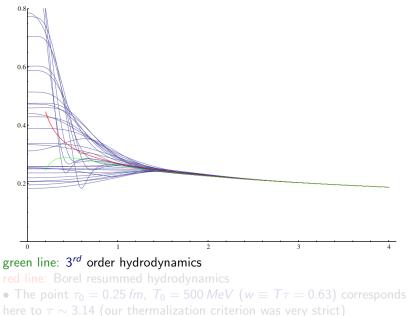
$$\pi au_{e\!f\!f}(au) \sim rac{1}{ au^{rac{1}{3}}} \qquad ext{in the } au o \infty ext{ limit}$$

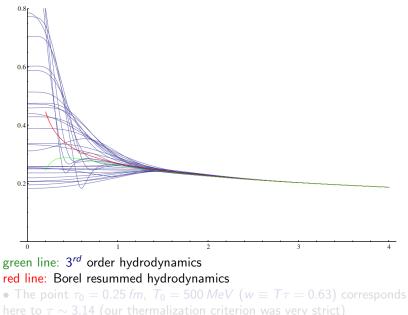
The coefficient '1' fixes the units of τ .

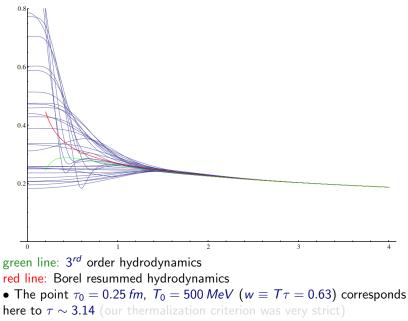
green line: 3^{rd} order hydrodynamics red line: Borel resummed hydrodynamics • The point $\tau_0 = 0.25 \text{ fm}$, $T_0 = 500 \text{ MeV}$ ($w \equiv T\tau = 0.63$) corresponds here to $\tau \sim 3.14$ (our thermalization criterion was very strict)

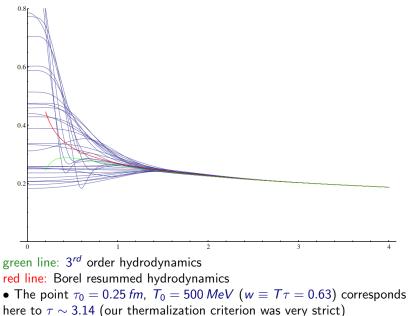


27 / 32









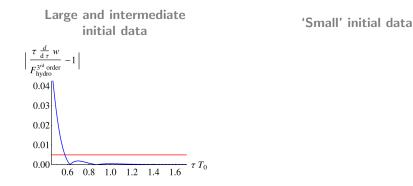
27 / 32

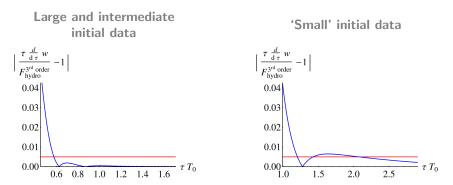
Large and intermediate initial data

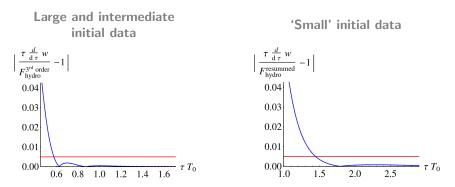
'Small' initial data

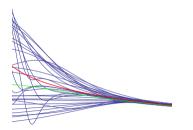
Large and intermediate initial data

'Small' initial data









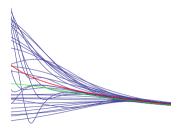
Is there a simple phenomenological model simpler than 5D Einstein's equations??

 $egin{array}{l} T_{\mu
u}(T,u^
ho) \ T_{\mu
u}(T,u^
ho,\ldots) \end{array}$

hydrodynamics

hydrodynamics + additional DOF

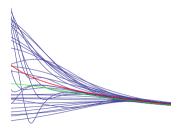
c.f. anisotropic hydrodynamics of Florkowski, Strickland and collaborators



Is there a simple phenomenological model simpler than 5D Einstein's equations??

 $egin{array}{lll} T_{\mu
u}({\mathcal T},u^
ho)&\sim&{
m hydrodynamics}\ T_{\mu
u}({\mathcal T},u^
ho,\ldots)&\sim&{
m hydrodynamics}+{
m additional}\;{
m DOF} \end{array}$

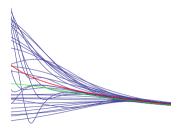
c.f. anisotropic hydrodynamics of Florkowski, Strickland and collaborators



Is there a simple phenomenological model simpler than 5D Einstein's equations??

 $T_{\mu
u}(T, u^{
ho}) \sim hydrodynamics$ $T_{\mu
u}(T, u^{
ho}, ...) \sim hydrodynamics + additional DOF$

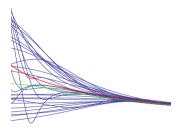
c.f. anisotropic hydrodynamics of Florkowski, Strickland and collaborators



Is there a simple phenomenological model simpler than 5D Einstein's equations??

 $T_{\mu
u}(T, u^{
ho}) \sim hydrodynamics$ $T_{\mu
u}(T, u^{
ho}, ...) \sim hydrodynamics + additional DOF$

c.f. anisotropic hydrodynamics of Florkowski, Strickland and collaborators



Is there a simple phenomenological model simpler than 5D Einstein's equations??

 $T_{\mu
u}(T, u^{
ho}) \sim hydrodynamics$ $T_{\mu
u}(T, u^{
ho}, ...) \sim hydrodynamics + additional DOF$

c.f. anisotropic hydrodynamics of Florkowski, Strickland and collaborators

Quasinormal modes — still very preliminary!

- ► Each quasinormal mode represents an independent degree of freedom from the 4D perspective...
- Recall the generic structure of QNM modes for a boost-invariant flow (including first viscous corrections)

 $\delta \varepsilon(\tau) \sim \tau^{-2} e^{-i\omega_{QNM}\int \pi T(\tau)d\tau}$

- One can estimate that $s = \int_0^\tau \pi T(\tau) d\tau$ at the transition to hydrodynamics would set the scale for how many QNM would be relevant there...
- ▶ It turns out that s = 1.6...3 and $Im \omega_{QNM} = 2.75, 4.76, 6.77, ...$
- A few additional DOF might suffice?

Each quasinormal mode represents an independent degree of freedom from the 4D perspective...

 Recall the generic structure of QNM modes for a boost-invariant flow (including first viscous corrections)

 $\delta \varepsilon(\tau) \sim \tau^{-2} e^{-i\omega_{QNM}\int \pi T(\tau)d\tau}$

- One can estimate that $s = \int_0^\tau \pi T(\tau) d\tau$ at the transition to hydrodynamics would set the scale for how many QNM would be relevant there...
- ▶ It turns out that s = 1.6...3 and $Im \omega_{QNM} = 2.75, 4.76, 6.77, ...$
- A few additional DOF might suffice?

- Each quasinormal mode represents an independent degree of freedom from the 4D perspective...
- Recall the generic structure of QNM modes for a boost-invariant flow (including first viscous corrections)

 $\delta \varepsilon(\tau) \sim \tau^{-2} e^{-i\omega_{\text{QNM}}\int \pi T(\tau) d\tau}$

- One can estimate that $s = \int_0^\tau \pi T(\tau) d\tau$ at the transition to hydrodynamics would set the scale for how many QNM would be relevant there...
- ▶ It turns out that s = 1.6...3 and $Im \omega_{QNM} = 2.75, 4.76, 6.77, ...$
- A few additional DOF might suffice?

- Each quasinormal mode represents an independent degree of freedom from the 4D perspective...
- Recall the generic structure of QNM modes for a boost-invariant flow (including first viscous corrections)

$$\delta \varepsilon(\tau) \sim \tau^{-2} e^{-i\omega_{QNM}\int \pi T(\tau)d\tau}$$

- One can estimate that $s = \int_0^\tau \pi T(\tau) d\tau$ at the transition to hydrodynamics would set the scale for how many QNM would be relevant there...
- ▶ It turns out that s = 1.6...3 and $Im \omega_{QNM} = 2.75, 4.76, 6.77, ...$
- A few additional DOF might suffice?

- Each quasinormal mode represents an independent degree of freedom from the 4D perspective...
- Recall the generic structure of QNM modes for a boost-invariant flow (including first viscous corrections)

$$\delta \varepsilon(\tau) \sim \tau^{-2} e^{-i\omega_{QNM}\int \pi T(\tau)d\tau}$$

- One can estimate that $s = \int_0^\tau \pi T(\tau) d\tau$ at the transition to hydrodynamics would set the scale for how many QNM would be relevant there...
- It turns out that s = 1.6...3 and $Im \omega_{QNM} = 2.75, 4.76, 6.77, ...$
- A few additional DOF might suffice?

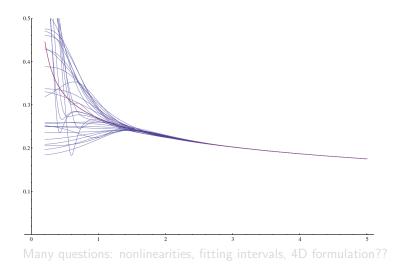
- Each quasinormal mode represents an independent degree of freedom from the 4D perspective...
- Recall the generic structure of QNM modes for a boost-invariant flow (including first viscous corrections)

$$\delta \varepsilon(\tau) \sim \tau^{-2} e^{-i\omega_{QNM}\int \pi T(\tau)d\tau}$$

- One can estimate that $s = \int_0^\tau \pi T(\tau) d\tau$ at the transition to hydrodynamics would set the scale for how many QNM would be relevant there...
- It turns out that s = 1.6...3 and $Im \omega_{QNM} = 2.75, 4.76, 6.77, ...$
- A few additional DOF might suffice?

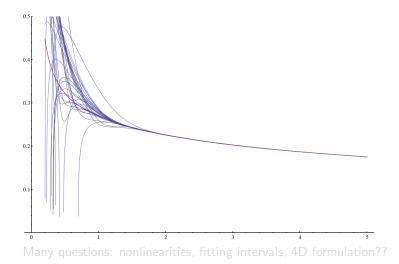
Many questions: nonlinearities, fitting intervals, 4D formulation??

 $T_{eff}(\tau)$ from numerical simulations

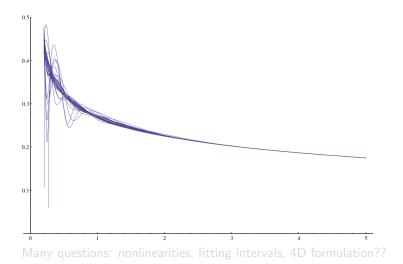


31 / 32

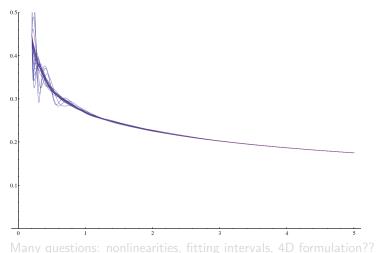
 $T_{\rm eff}(\tau)$ from numerical simulations with fitted 1st QNM subtracted



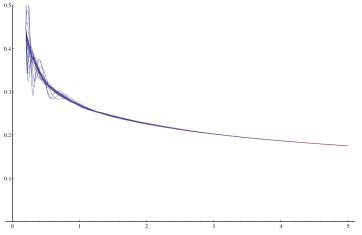
 $\mathcal{T}_{eff}(au)$ from numerical simulations with fitted $1^{st}+2^{nd}$ QNM subtracted



$T_{\rm eff}(\tau)$ from numerical simulations with fitted $1^{st}+2^{nd}+3^{rd}$ QNM subtracted



$T_{\rm eff}(\tau)$ from numerical simulations with fitted $1^{st}+2^{nd}+3^{rd}$ QNM subtracted



Many questions: nonlinearities, fitting intervals, 4D formulation??

- ▶ We calculated the gradient expansion to very high orders
- ► The hydrodynamic expansion has zero radius of convergence
- The singularities in the Borel plane have a clear physical origin they correspond to the lowest non-hydrodynamic modes/degrees of freedom
- ► Analogy with perturbative expansion in QFT and instanton effects
- Hydrodynamic expansion captures *quantitatively* fine details of these leading nonhydrodynamic modes
- We do not find poles on the positive real axis suggesting the existence of a Borel resummation, which can be constructed using Pade approximants
- ▶ Higher order hydrodynamics seems relevant for 'small' initial data...
- 4D perspective on deviations from hydro???
- QNM frequencies for QCD??? (lattice?)

- We calculated the gradient expansion to very high orders
- ▶ The hydrodynamic expansion has zero radius of convergence
- The singularities in the Borel plane have a clear physical origin they correspond to the lowest non-hydrodynamic modes/degrees of freedom
- ► Analogy with perturbative expansion in QFT and instanton effects
- Hydrodynamic expansion captures *quantitatively* fine details of these leading nonhydrodynamic modes
- We do not find poles on the positive real axis suggesting the existence of a Borel resummation, which can be constructed using Pade approximants
- ▶ Higher order hydrodynamics seems relevant for 'small' initial data...
- 4D perspective on deviations from hydro???
- QNM frequencies for QCD??? (lattice?)

- We calculated the gradient expansion to very high orders
- ► The hydrodynamic expansion has zero radius of convergence
- The singularities in the Borel plane have a clear physical origin they correspond to the lowest non-hydrodynamic modes/degrees of freedom
- ► Analogy with perturbative expansion in QFT and instanton effects
- Hydrodynamic expansion captures *quantitatively* fine details of these leading nonhydrodynamic modes
- We do not find poles on the positive real axis suggesting the existence of a Borel resummation, which can be constructed using Pade approximants
- ▶ Higher order hydrodynamics seems relevant for 'small' initial data...
- 4D perspective on deviations from hydro???
- QNM frequencies for QCD??? (lattice?)

- We calculated the gradient expansion to very high orders
- ► The hydrodynamic expansion has zero radius of convergence
- The singularities in the Borel plane have a clear physical origin they correspond to the lowest non-hydrodynamic modes/degrees of freedom
- ► Analogy with perturbative expansion in QFT and instanton effects
- Hydrodynamic expansion captures *quantitatively* fine details of these leading nonhydrodynamic modes
- We do not find poles on the positive real axis suggesting the existence of a Borel resummation, which can be constructed using Pade approximants
- ▶ Higher order hydrodynamics seems relevant for 'small' initial data...
- 4D perspective on deviations from hydro???
- QNM frequencies for QCD??? (lattice?)

- We calculated the gradient expansion to very high orders
- ► The hydrodynamic expansion has zero radius of convergence
- The singularities in the Borel plane have a clear physical origin they correspond to the lowest non-hydrodynamic modes/degrees of freedom
- Analogy with perturbative expansion in QFT and instanton effects
- Hydrodynamic expansion captures *quantitatively* fine details of these leading nonhydrodynamic modes
- We do not find poles on the positive real axis suggesting the existence of a Borel resummation, which can be constructed using Pade approximants
- ▶ Higher order hydrodynamics seems relevant for 'small' initial data...
- 4D perspective on deviations from hydro???
- QNM frequencies for QCD??? (lattice?)

- We calculated the gradient expansion to very high orders
- ► The hydrodynamic expansion has zero radius of convergence
- The singularities in the Borel plane have a clear physical origin they correspond to the lowest non-hydrodynamic modes/degrees of freedom
- Analogy with perturbative expansion in QFT and instanton effects
- Hydrodynamic expansion captures *quantitatively* fine details of these leading nonhydrodynamic modes
- We do not find poles on the positive real axis suggesting the existence of a Borel resummation, which can be constructed using Pade approximants
- ▶ Higher order hydrodynamics seems relevant for 'small' initial data...
- 4D perspective on deviations from hydro???
- QNM frequencies for QCD??? (lattice?)

- We calculated the gradient expansion to very high orders
- ► The hydrodynamic expansion has zero radius of convergence
- The singularities in the Borel plane have a clear physical origin they correspond to the lowest non-hydrodynamic modes/degrees of freedom
- Analogy with perturbative expansion in QFT and instanton effects
- Hydrodynamic expansion captures *quantitatively* fine details of these leading nonhydrodynamic modes
- We do not find poles on the positive real axis suggesting the existence of a Borel resummation, which can be constructed using Pade approximants
- ▶ Higher order hydrodynamics seems relevant for 'small' initial data...
- 4D perspective on deviations from hydro???
- QNM frequencies for QCD??? (lattice?)

- We calculated the gradient expansion to very high orders
- ► The hydrodynamic expansion has zero radius of convergence
- The singularities in the Borel plane have a clear physical origin they correspond to the lowest non-hydrodynamic modes/degrees of freedom
- Analogy with perturbative expansion in QFT and instanton effects
- Hydrodynamic expansion captures *quantitatively* fine details of these leading nonhydrodynamic modes
- We do not find poles on the positive real axis suggesting the existence of a Borel resummation, which can be constructed using Pade approximants
- Higher order hydrodynamics seems relevant for 'small' initial data...
- 4D perspective on deviations from hydro???
- QNM frequencies for QCD??? (lattice?)

- We calculated the gradient expansion to very high orders
- ► The hydrodynamic expansion has zero radius of convergence
- The singularities in the Borel plane have a clear physical origin they correspond to the lowest non-hydrodynamic modes/degrees of freedom
- Analogy with perturbative expansion in QFT and instanton effects
- Hydrodynamic expansion captures *quantitatively* fine details of these leading nonhydrodynamic modes
- We do not find poles on the positive real axis suggesting the existence of a Borel resummation, which can be constructed using Pade approximants
- Higher order hydrodynamics seems relevant for 'small' initial data...
- 4D perspective on deviations from hydro???
- QNM frequencies for QCD??? (lattice?)

- We calculated the gradient expansion to very high orders
- ► The hydrodynamic expansion has zero radius of convergence
- The singularities in the Borel plane have a clear physical origin they correspond to the lowest non-hydrodynamic modes/degrees of freedom
- Analogy with perturbative expansion in QFT and instanton effects
- Hydrodynamic expansion captures *quantitatively* fine details of these leading nonhydrodynamic modes
- We do not find poles on the positive real axis suggesting the existence of a Borel resummation, which can be constructed using Pade approximants
- Higher order hydrodynamics seems relevant for 'small' initial data...
- 4D perspective on deviations from hydro???
- QNM frequencies for QCD??? (lattice?)