

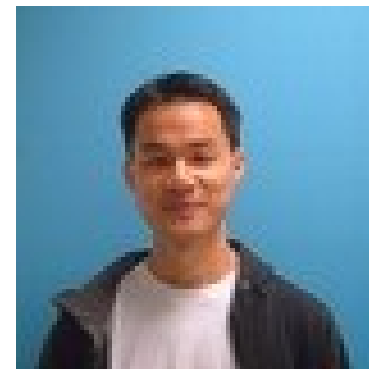
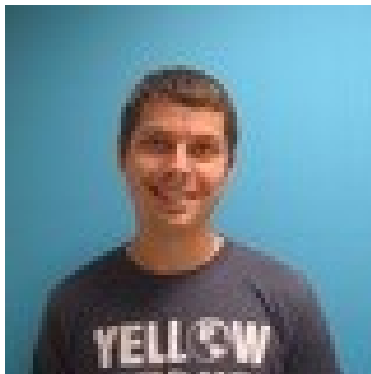
Today's Talk

Peter Arnold

- I. A fun problem in gravity
- II. What it's good for

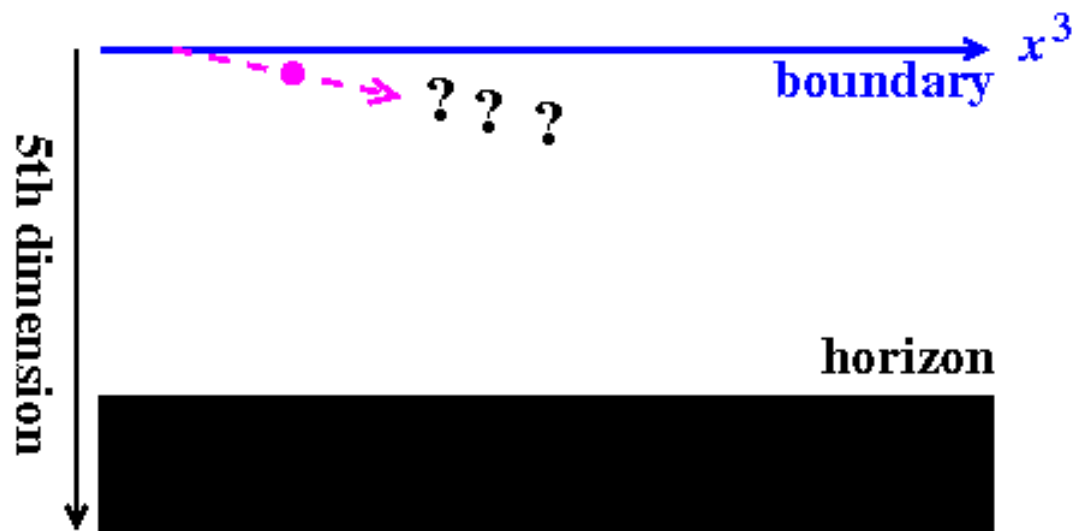
arXiv: 1212.3321

with Philip Szepietowski, Diana Vaman, and Gabriel Wong



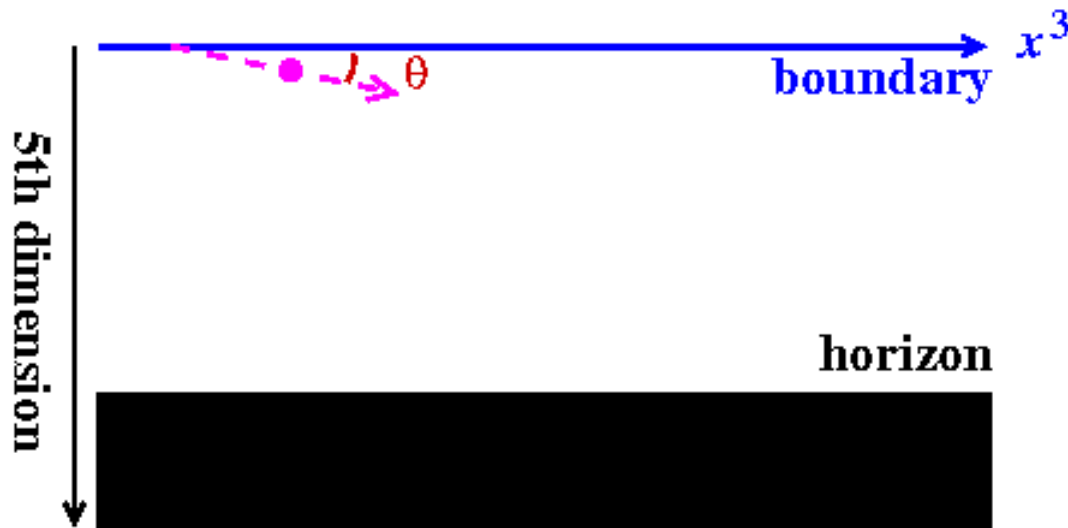
A fun problem in gravity

Suppose a **graviton** is launched from the boundary of $(\text{AdS}_5\text{-Schwarzschild}) \times S^5$.
What happens to it?



A fun problem in gravity

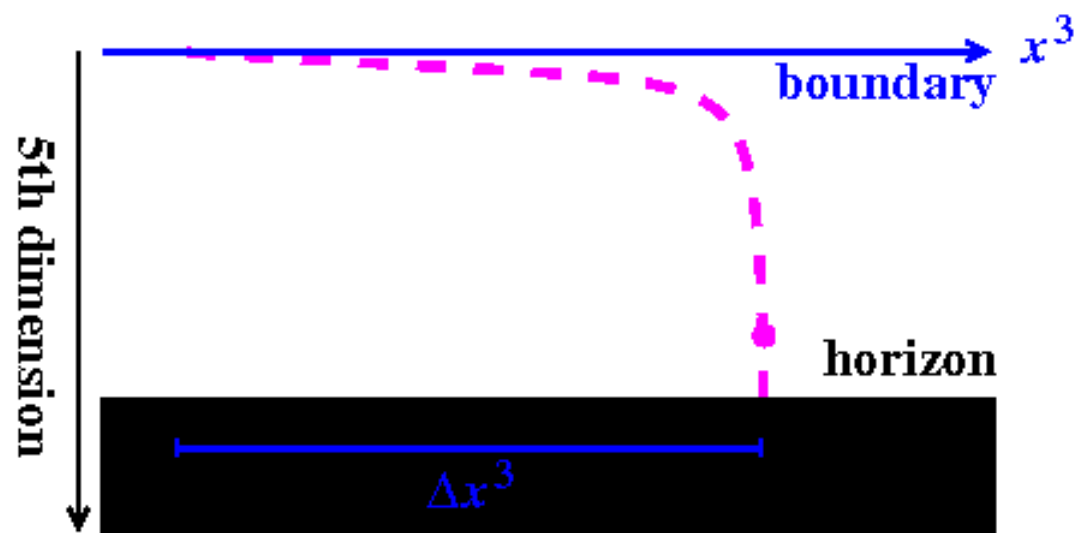
Suppose a **graviton** is launched from the boundary of $\text{AdS}_5\text{-Schwarzschild} \times S^5$.
What happens to it?



I'm going to be interested in gravitons (or photons or whatever)

- that start out moving at a small angle θ relative to the boundary
- with large momentum q_3 in the x^3 direction
- with a localized wave function in $\text{AdS}_5\text{-Schwarzschild}$


First answer




Note for later:

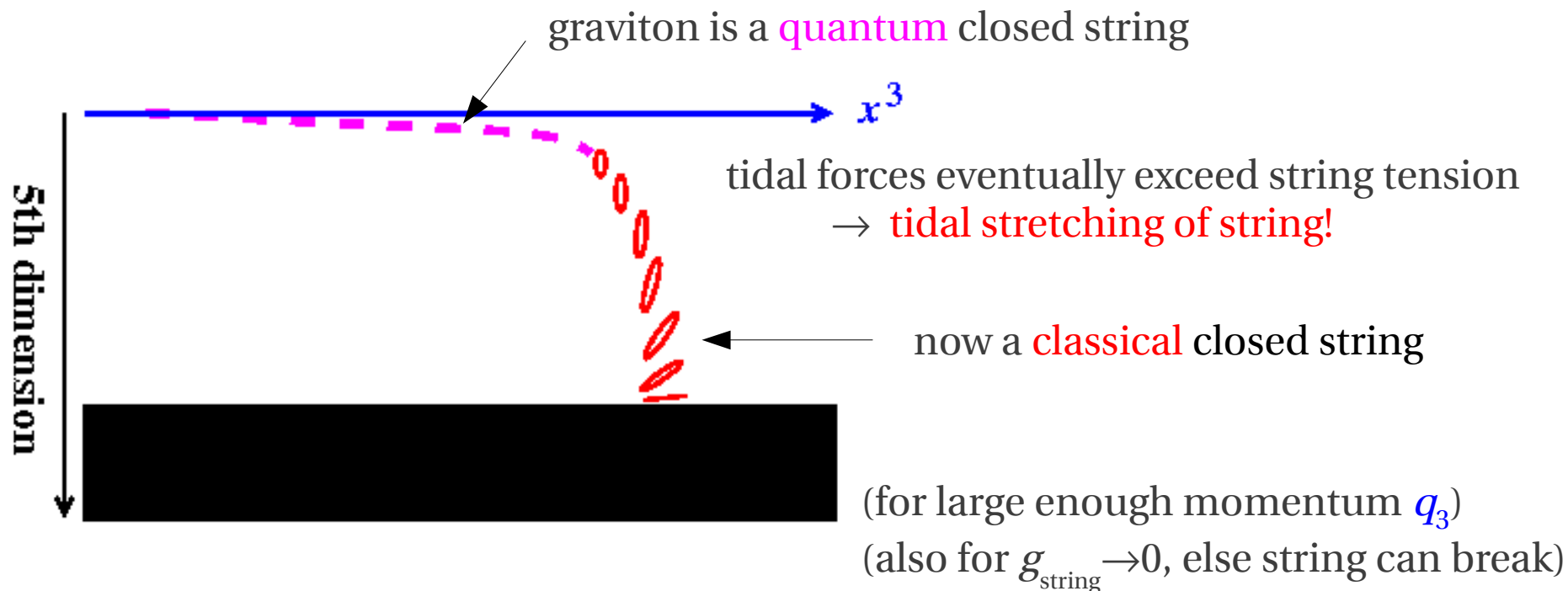
- graviton travels a finite distance Δx^3 in x^3
- takes infinite boundary-time x^0 to fall into black hole

But really

- =  w/ internal degrees of freedom in ground state
 — proper size $\sim (\text{string tension})^{-1/2} \sim (\alpha')^{1/2}$


But really

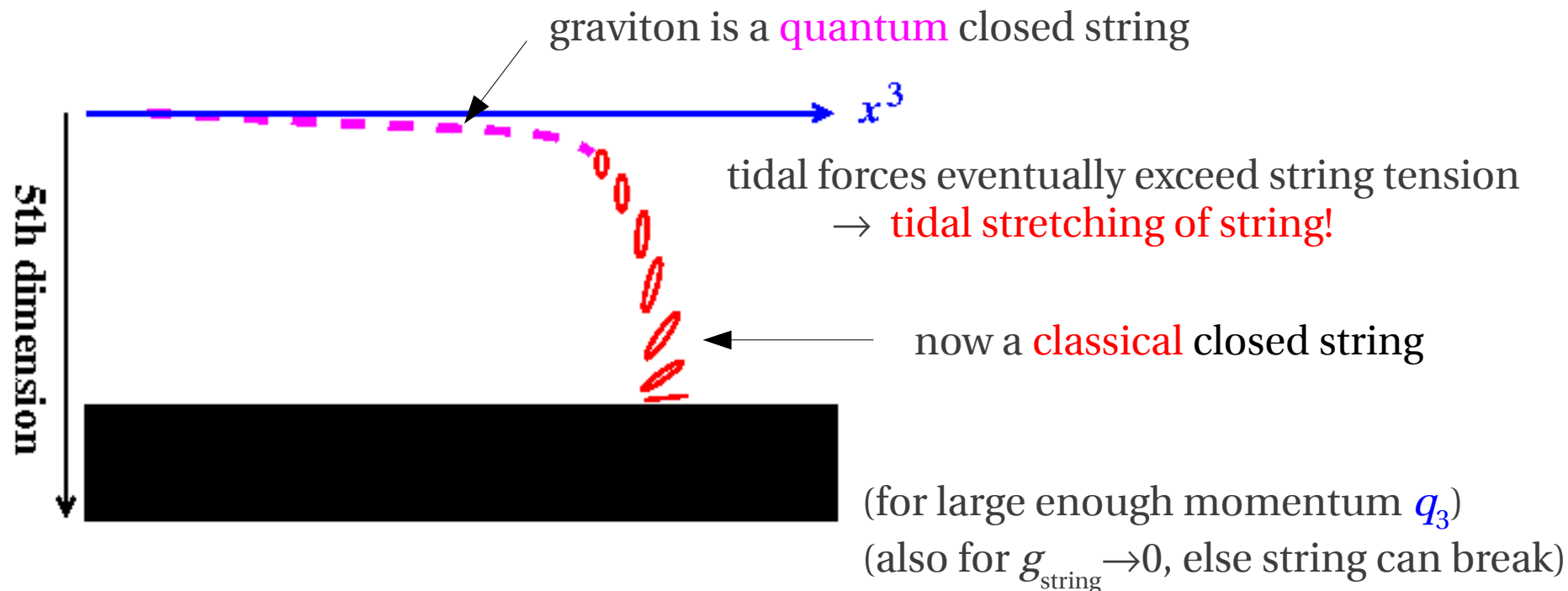
- =  w/ internal degrees of freedom in ground state
 — proper size $\sim (\text{string tension})^{-1/2} \sim (\alpha')^{1/2}$



Q: Is it possible to quantitatively calculate the late-time probability distribution of classical string configurations?

But really

- =  w/ internal degrees of freedom in ground state
 — proper size $\sim (\text{string tension})^{-1/2} \sim (\alpha')^{1/2}$




Q: Is it possible to quantitatively calculate the late-time probability distribution of classical string configurations?

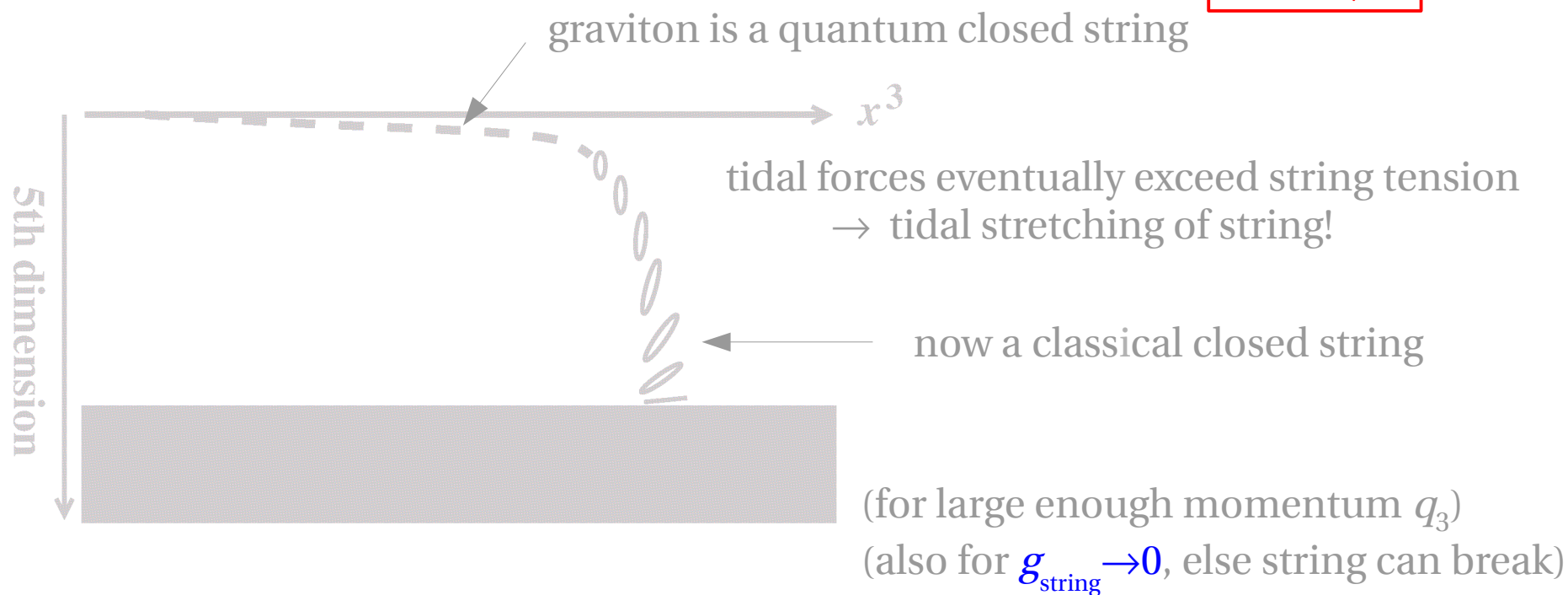
A: Yes, for a certain range of parameters.

Uses Penrose limit and quantization of strings in pp-wave backgrounds

But really

- =  w/ internal degrees of freedom in ground state
- proper size $\sim (\text{string tension})^{-1/2} \sim (\alpha')^{1/2}$

$$\alpha' \propto \frac{1}{\sqrt{\lambda}}$$



Q: Is it possible to quantitatively calculate the late-time probability distribution of classical string configurations?

A: Yes, for a certain range of parameters.

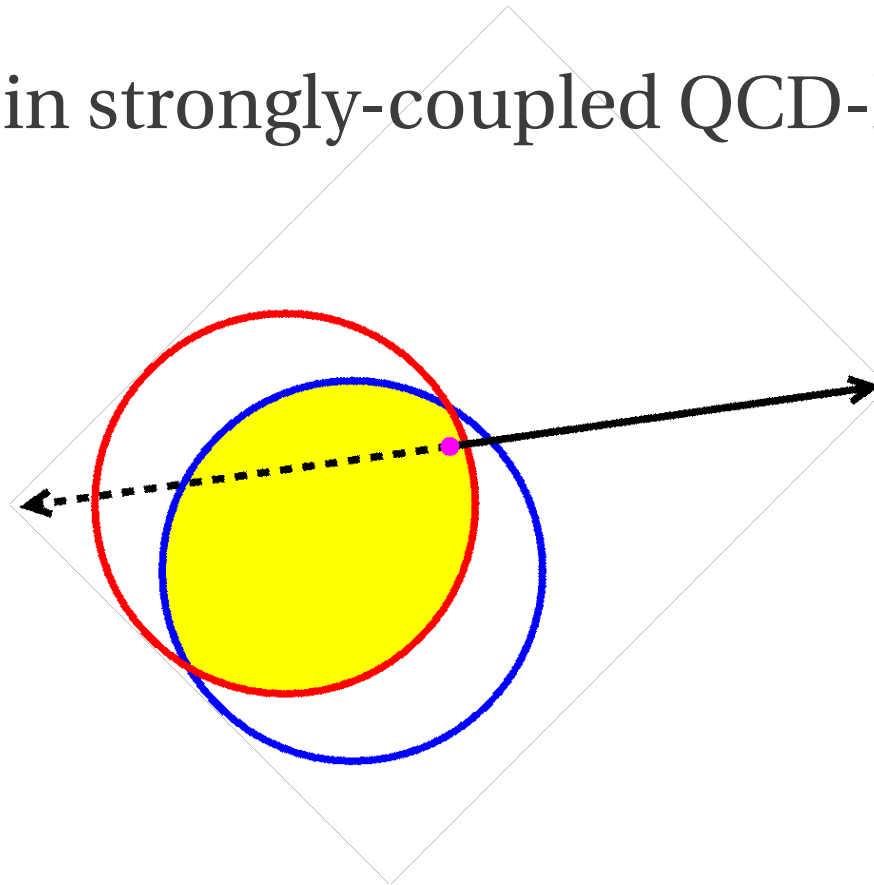
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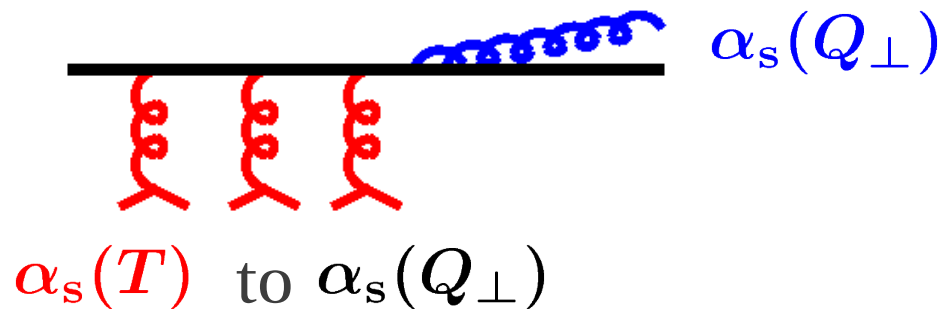
$$N_c \rightarrow \infty$$

What's it good for?

Applied Holography: Problem arises in the theory of

Jet quenching in strongly-coupled QCD-like plasmas





$$Q_\perp \sim (\hat{q}E)^{1/4}$$

typical transverse momentum transfer during formation time

How stopping length scales with energy (massless case)

weak coupling: $\alpha_s \sim \alpha_s$ small $\ell_{\text{stop}} \propto E^{1/2}$ (up to logs)
 [this scaling a corollary of BDMPS and Z (1996)]

mixed coupling: $\left. \begin{array}{l} \alpha_s \text{ BIG} \\ \alpha_s \text{ small} \end{array} \right\} \ell_{\text{stop}} \propto E^{1/2}$ (believed)
 $\ell_{\text{stop}} \sim \alpha_s^{-1} (E/\hat{q})^{1/2}$
 [e.g. Liu, Rajagopal, Wiedeman (2006)]

all strong coupling: $\alpha_s = \alpha_s$ BIG $\ell_{\text{stop}} \propto E^{1/3}$
 (Large- N_c $\mathcal{N}=4$ SYM, etc.)

[Gubser, Gollota, Pufu, Rocha; Hatta, Iancu, Mueller; Chesler, Jensen, Karch, Yaffe (2008)]

Interesting: Exponent in $\ell_{\text{stop}} \propto E^\nu$ can depend on α_s .

Defining Stopping Distance

BIG $\alpha_s = \alpha_s$: Large- N_c $\mathcal{N}=4$ SYM, etc. with $N_c \alpha_s \rightarrow \infty$



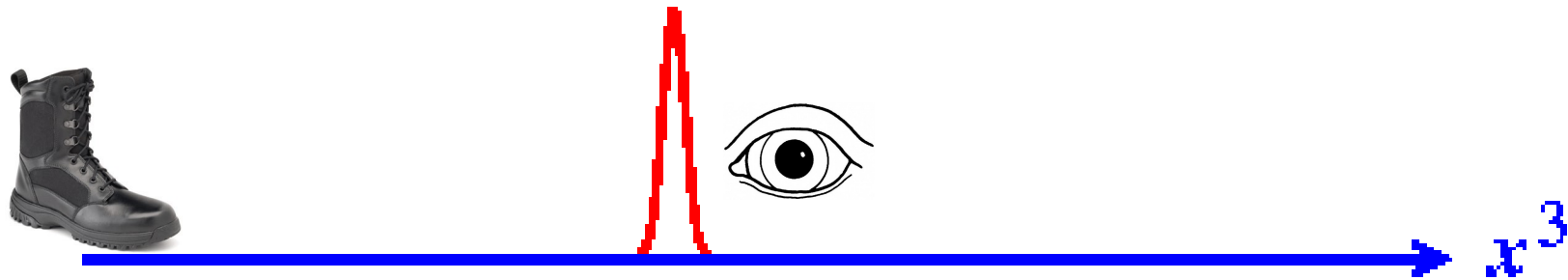
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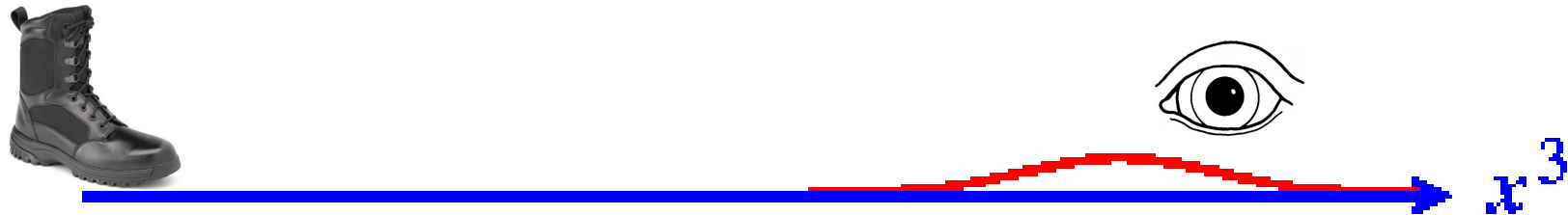
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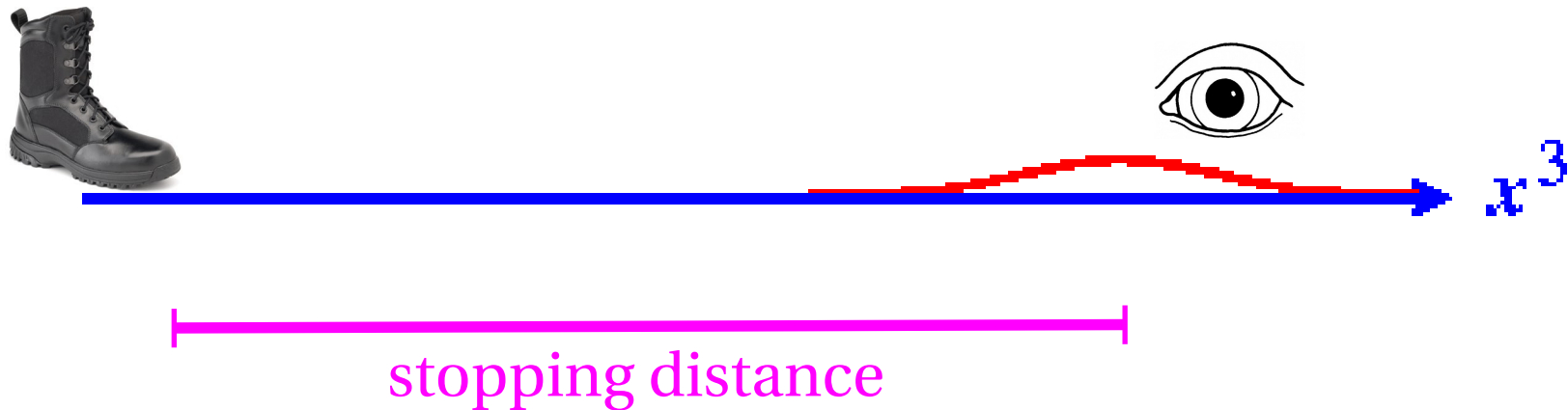
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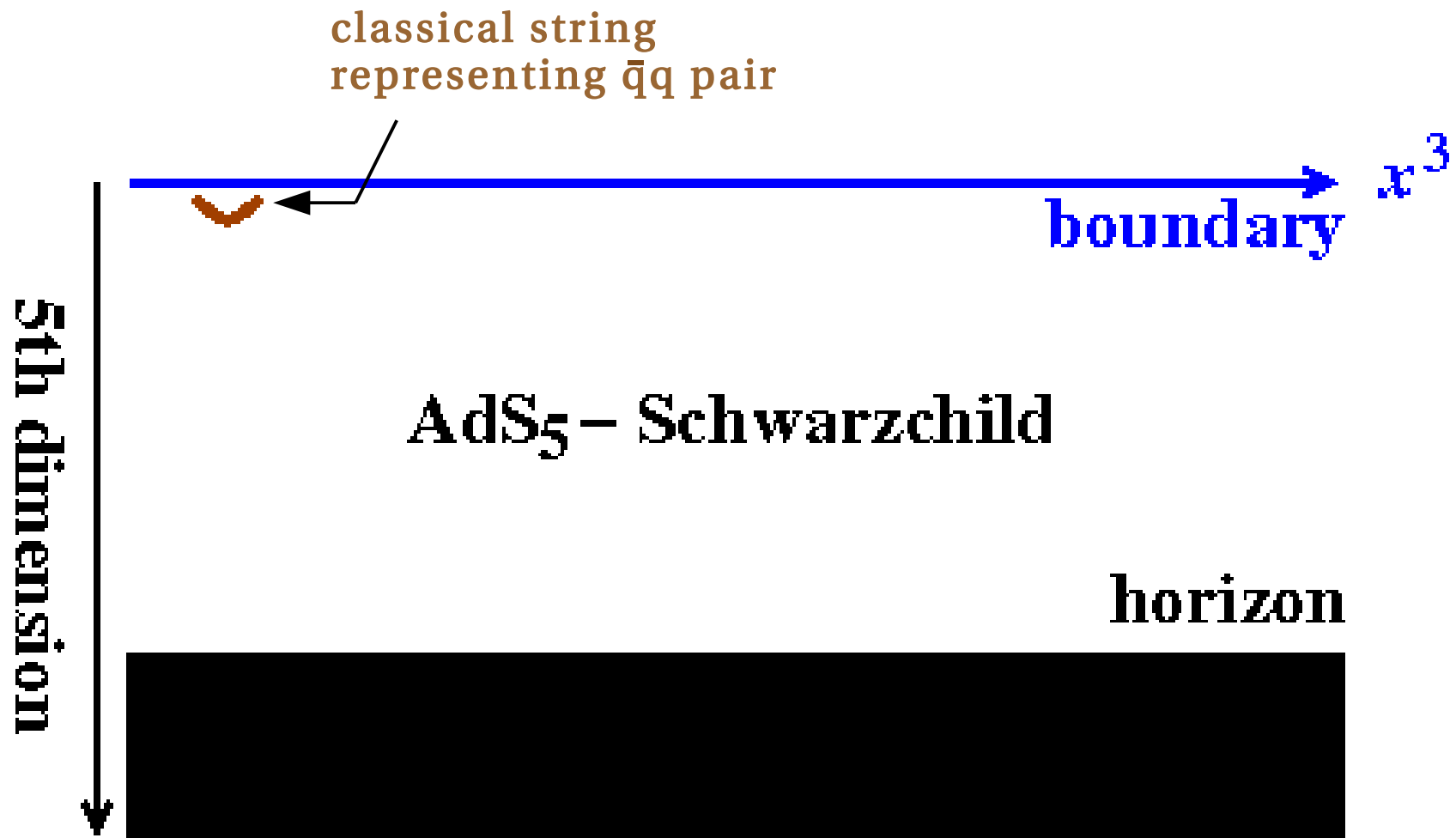
AdS/CFT Results

A. Results for $\lambda = \infty$ and $N_c = \infty$

B. Results for $1 \ll \lambda < \infty$ and $N_c = \infty$

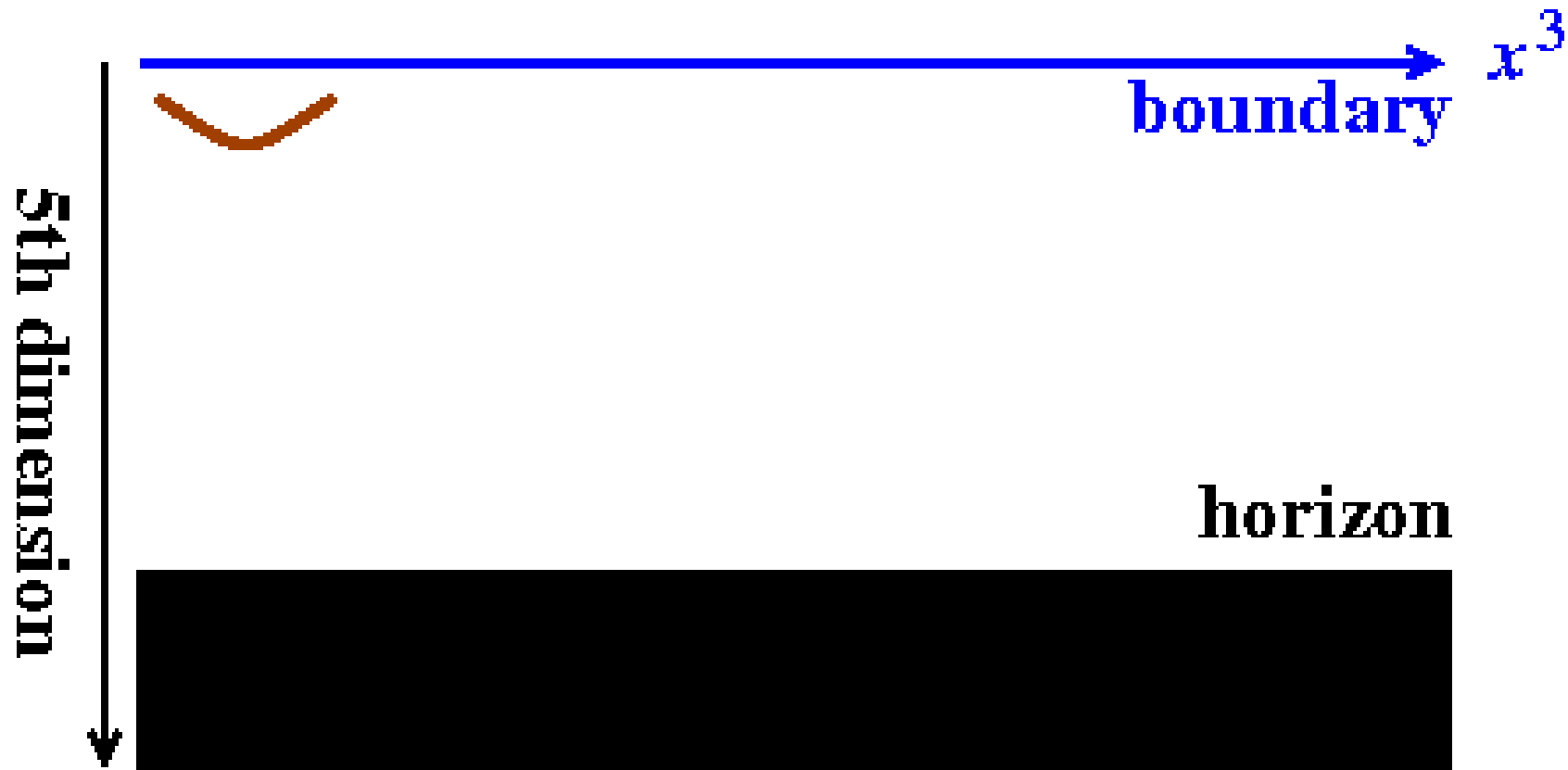
Early AdS Calculations

Example: Classical string calculation of
Chesler, Jensen, Karch, Yaffe (2008)



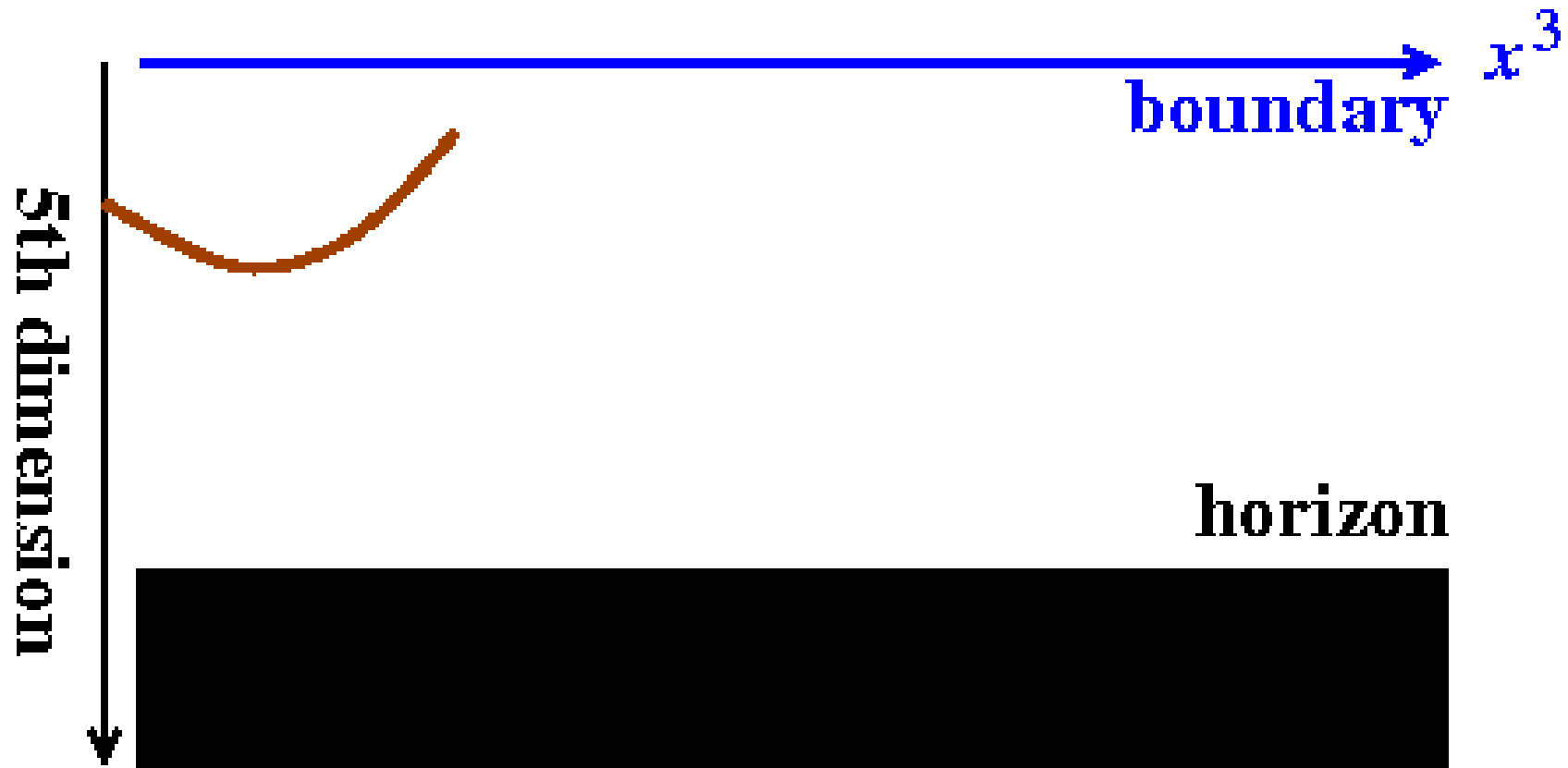
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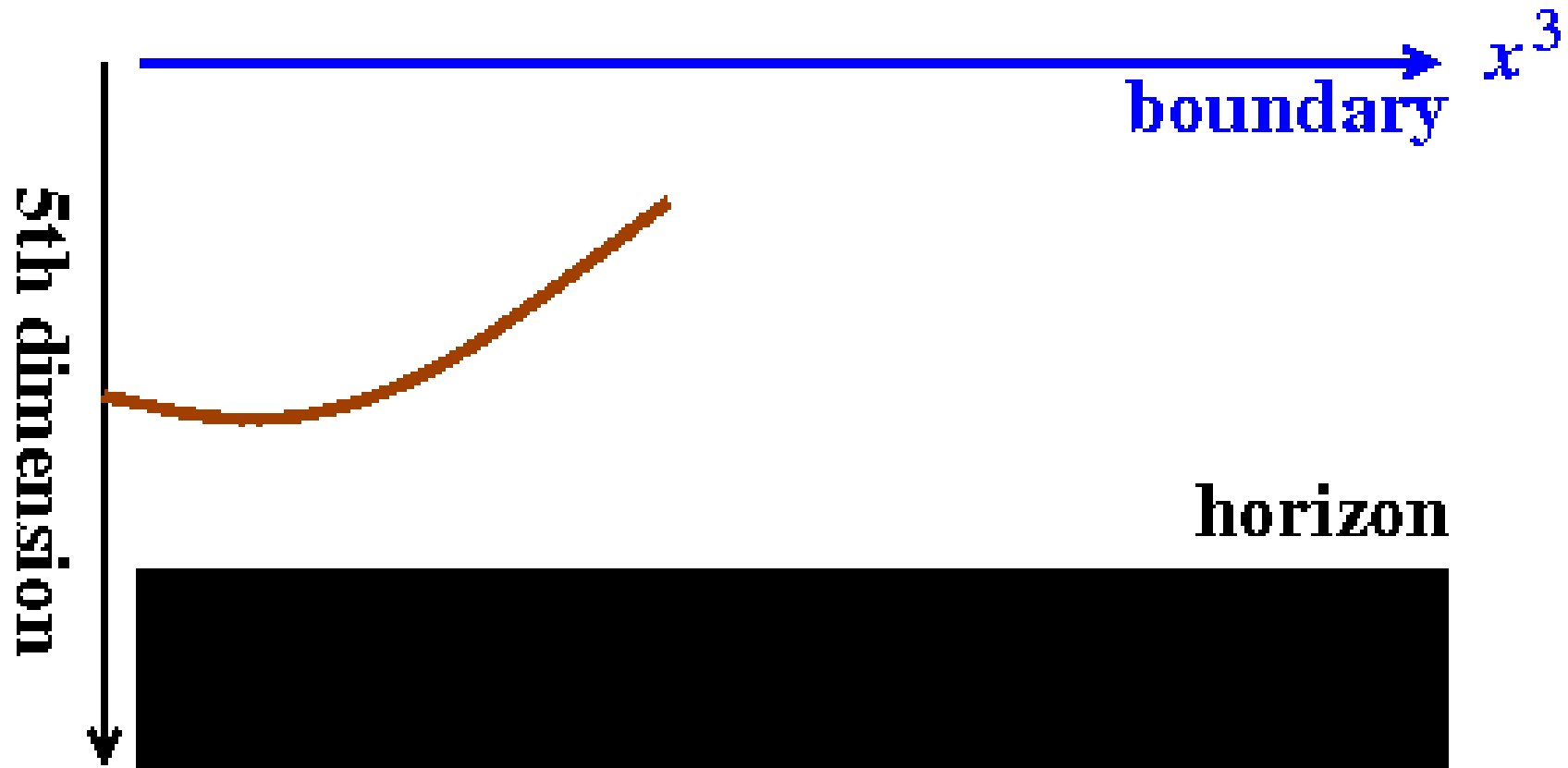
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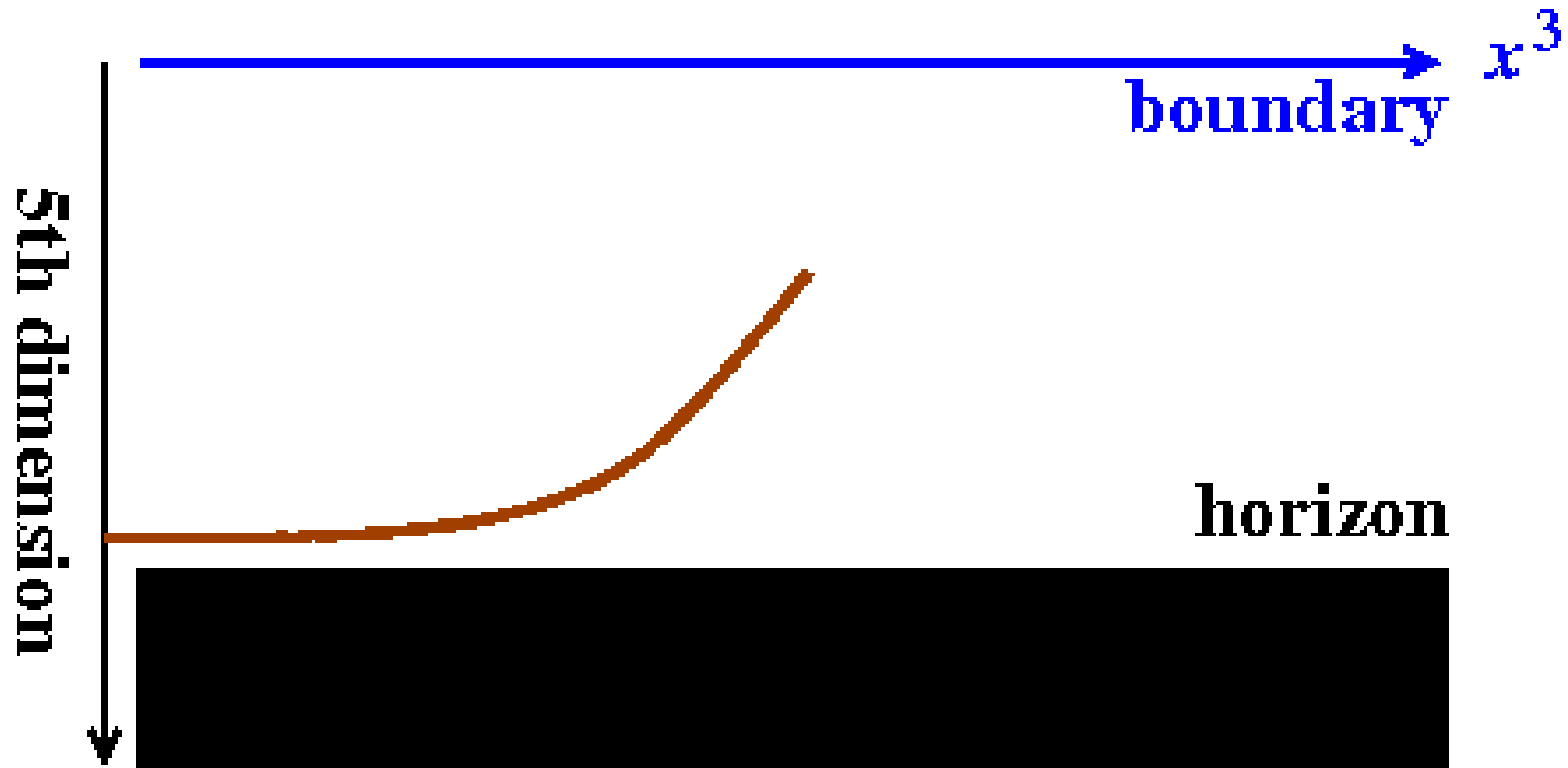
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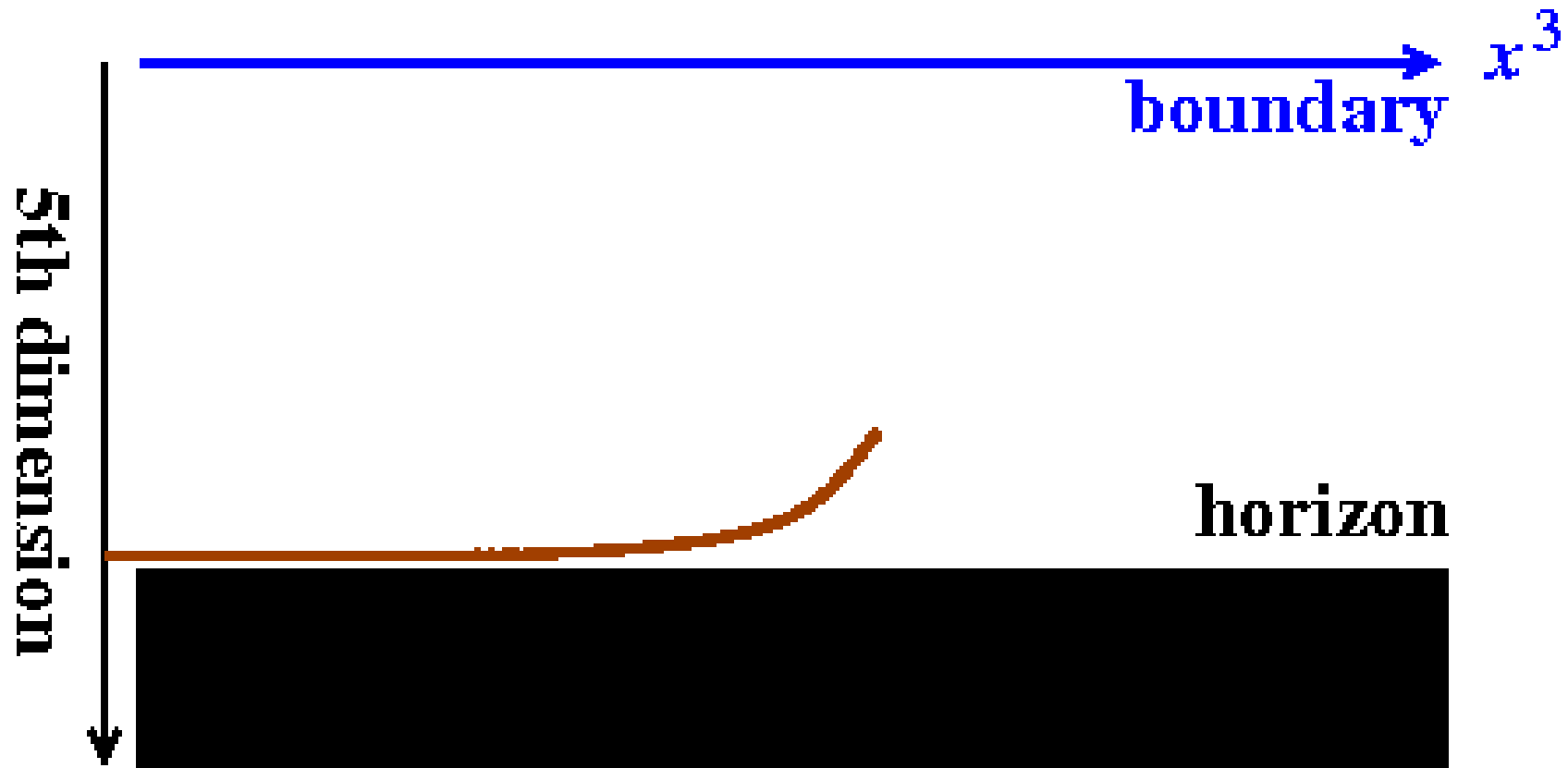
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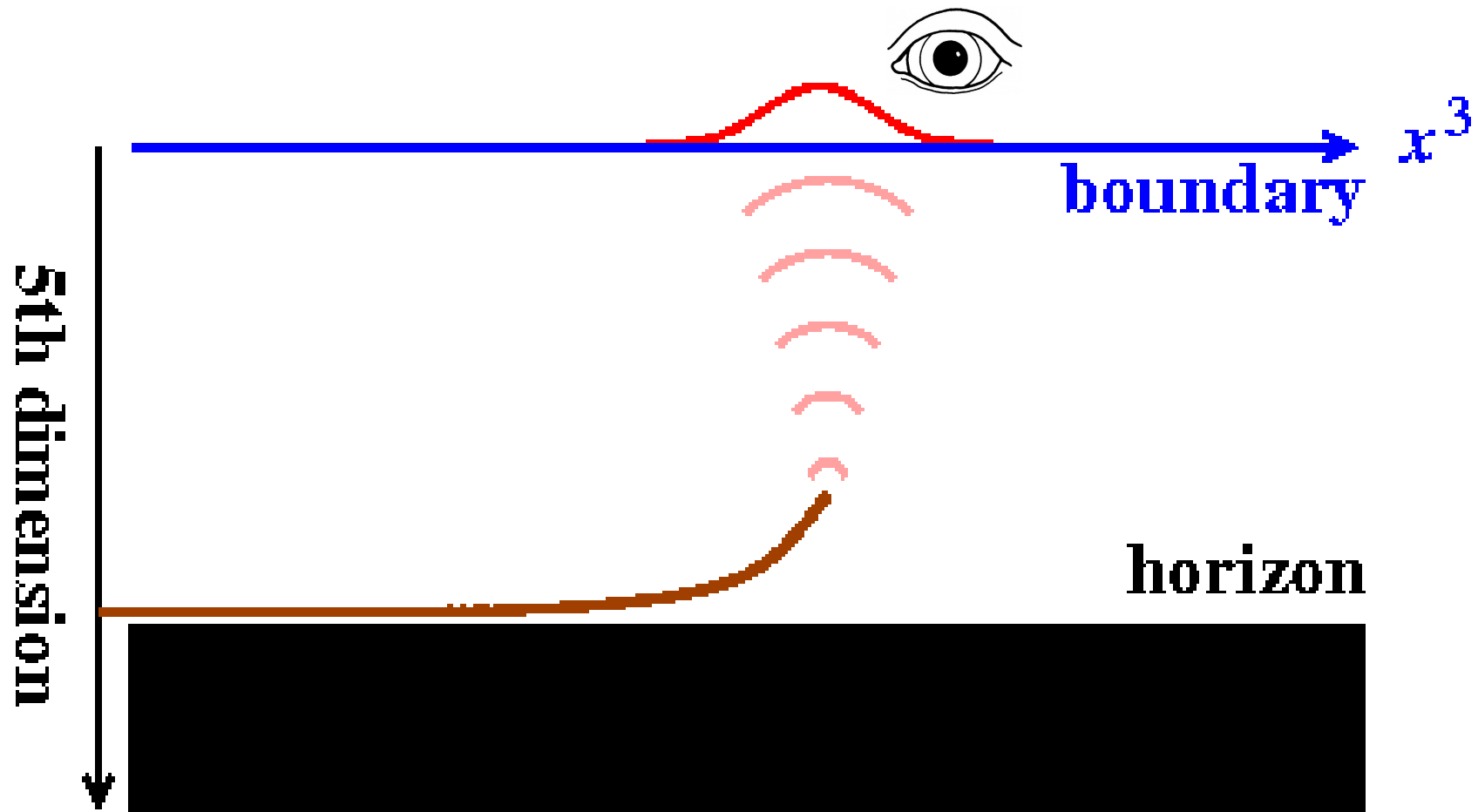
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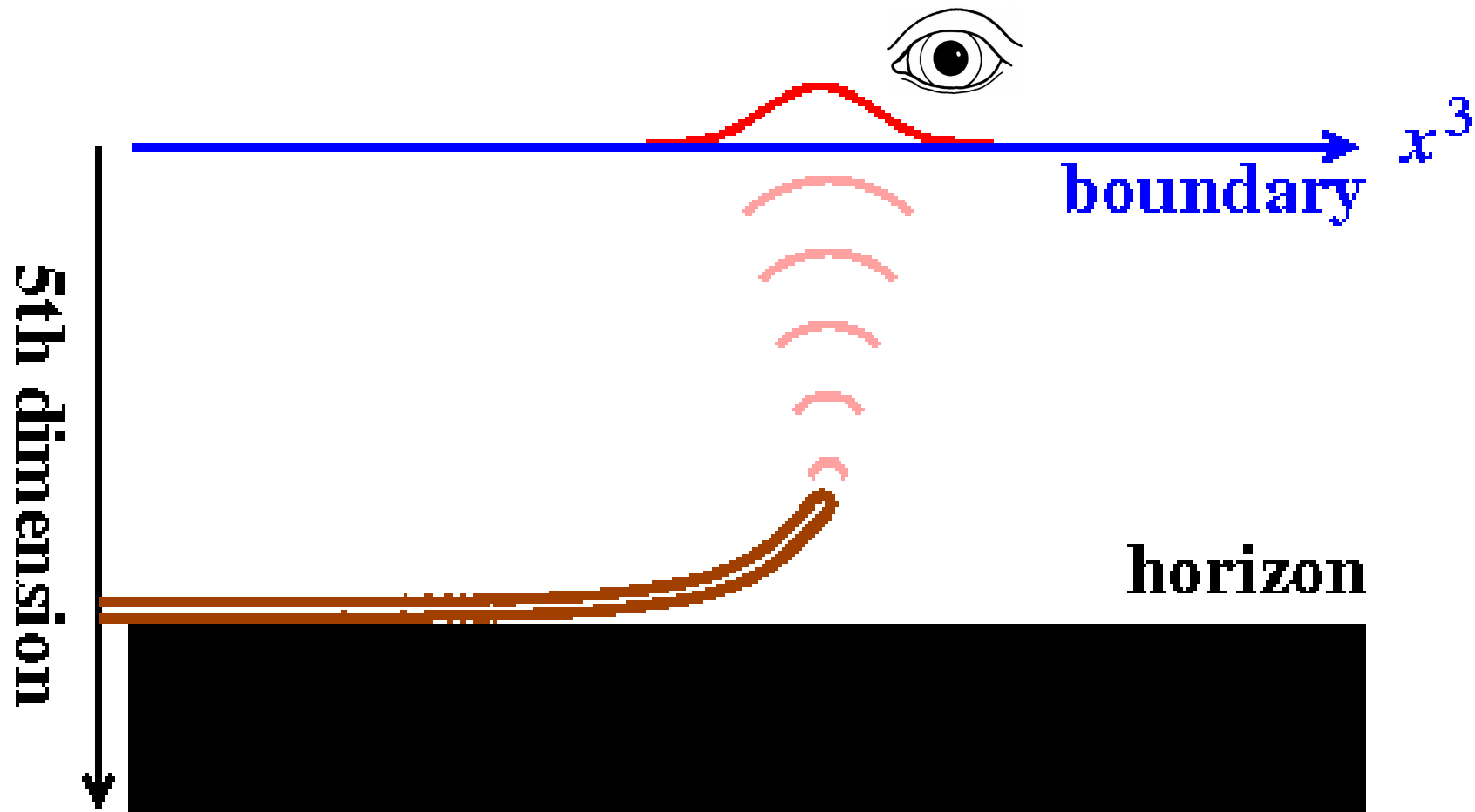
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Early AdS Calculations

Gubser, Gulotta, Pufu, Rocha (2008) had a roughly similar picture but with a folded string, representing a gluon jet.

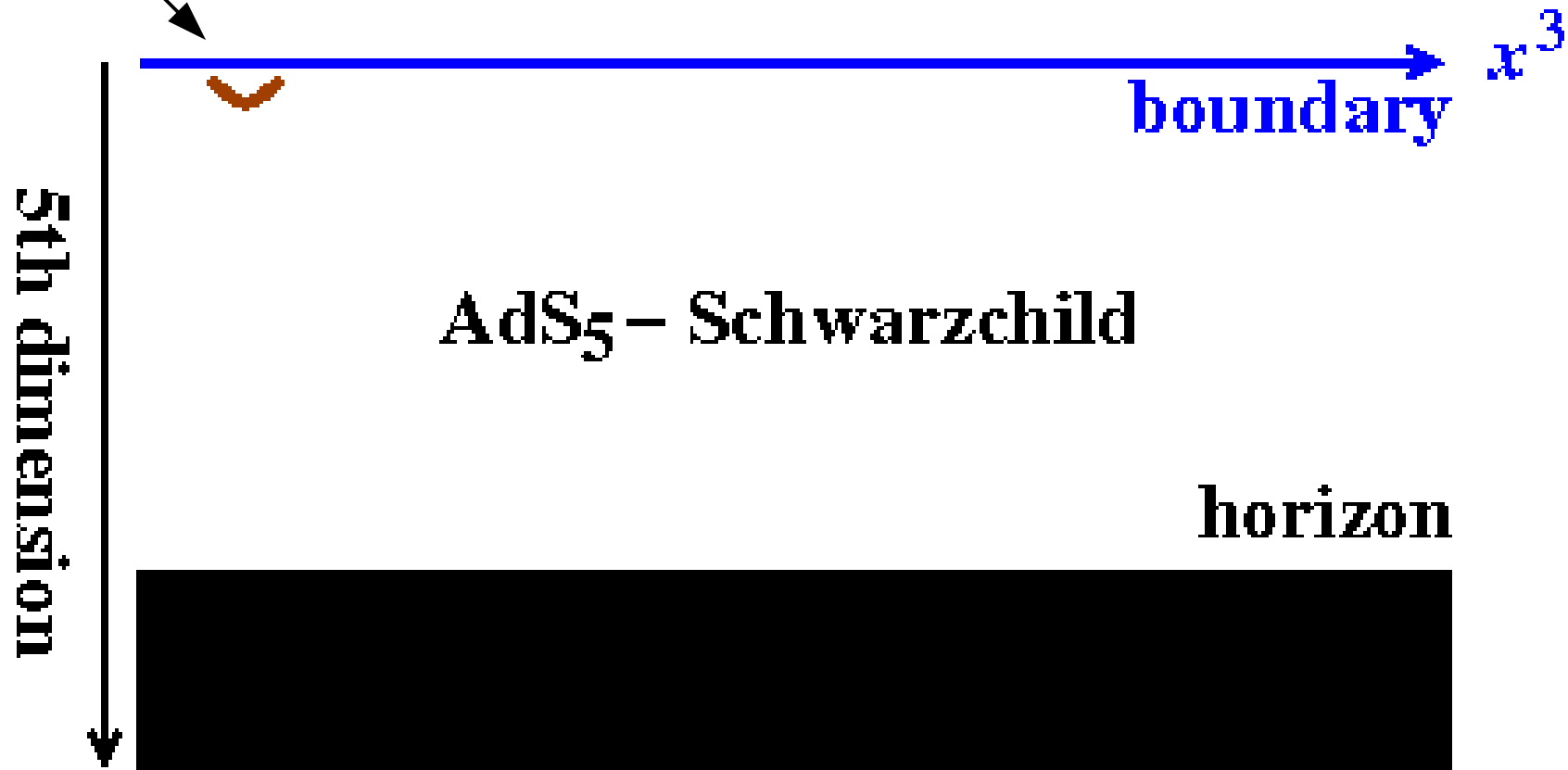


Something slightly dissatisfying: What is the



?

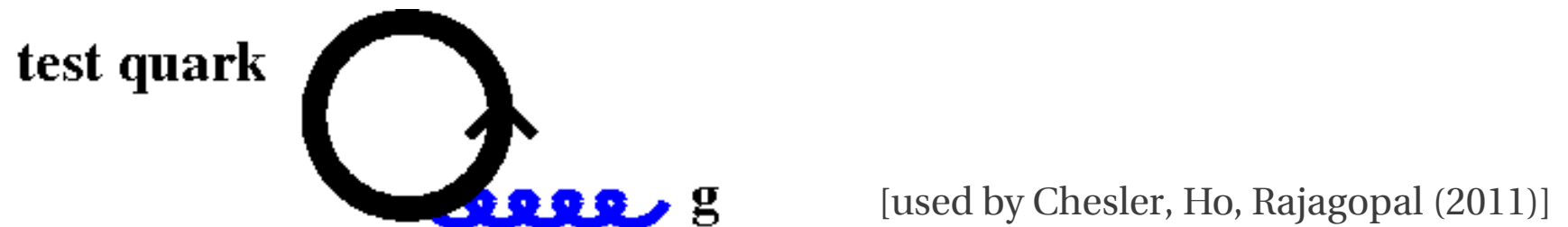
The initial configuration has been expressed in the gravity dual.
How precisely do I set up the problem in the 3+1 dim. field theory?



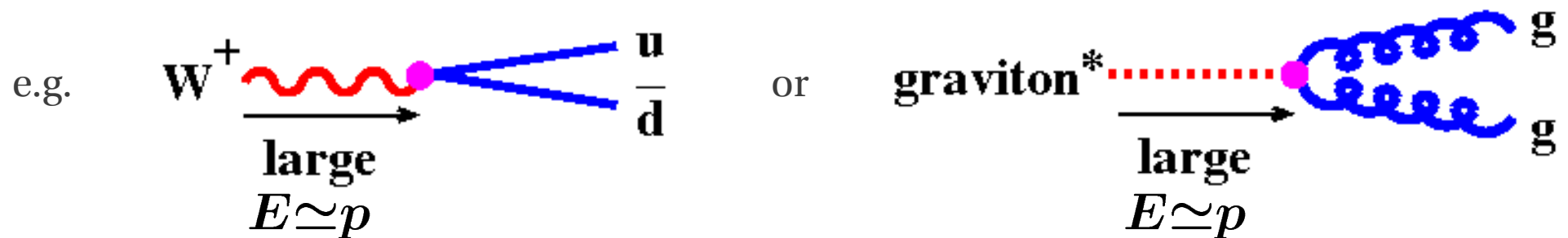
Choice of in the QFT

Gedanken experiments for creating localized, very high-momentum excitations in the plasma.

Synchrotron method: Drag a heavy test quark around in a circle to make a beam of gluon synchrotron radiation.



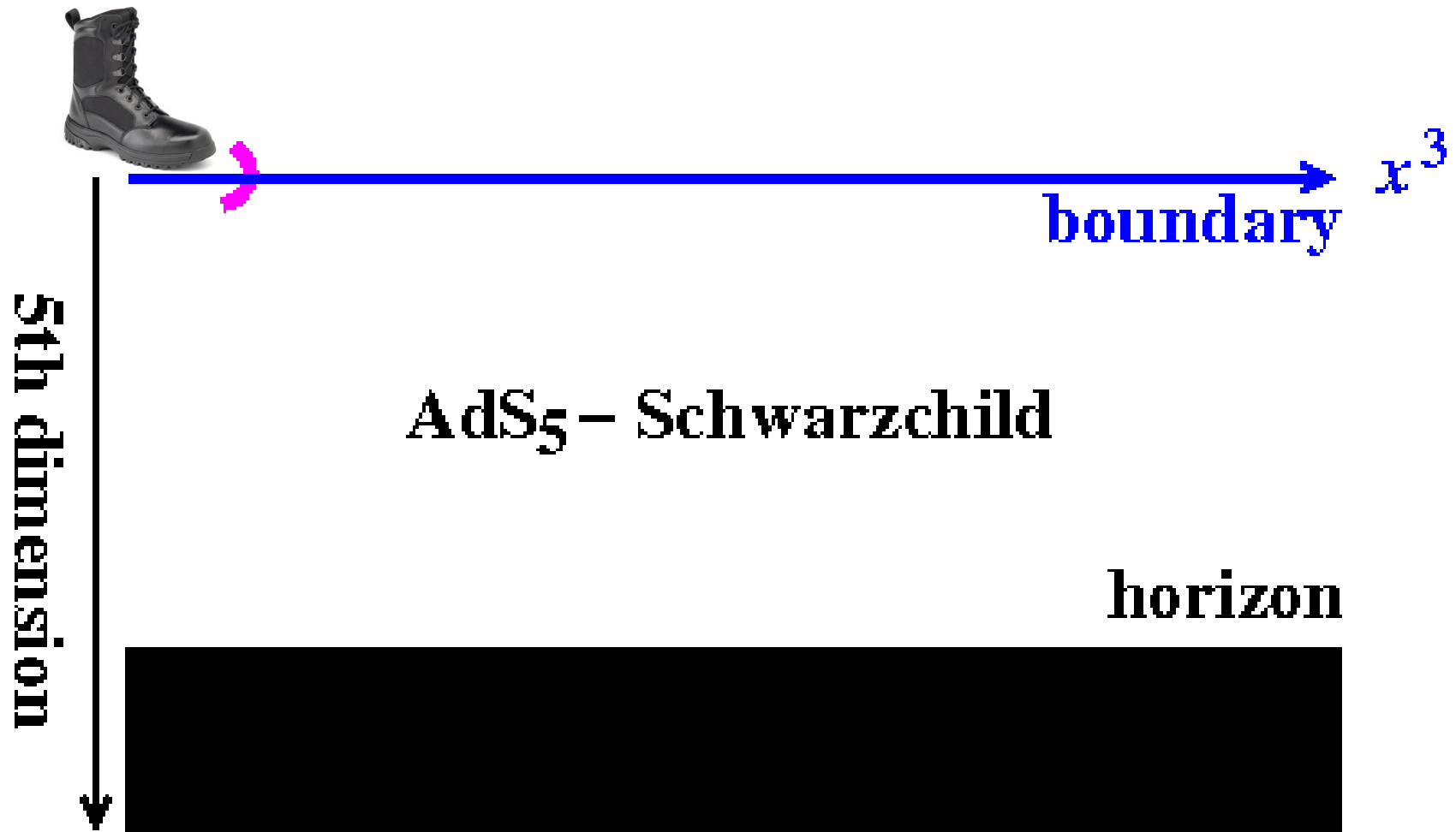
Our method: Analogous to considering the hadronic decay of some very high-momentum, unstable particle in a QCD plasma.



[Arnold & Vaman (2010)]

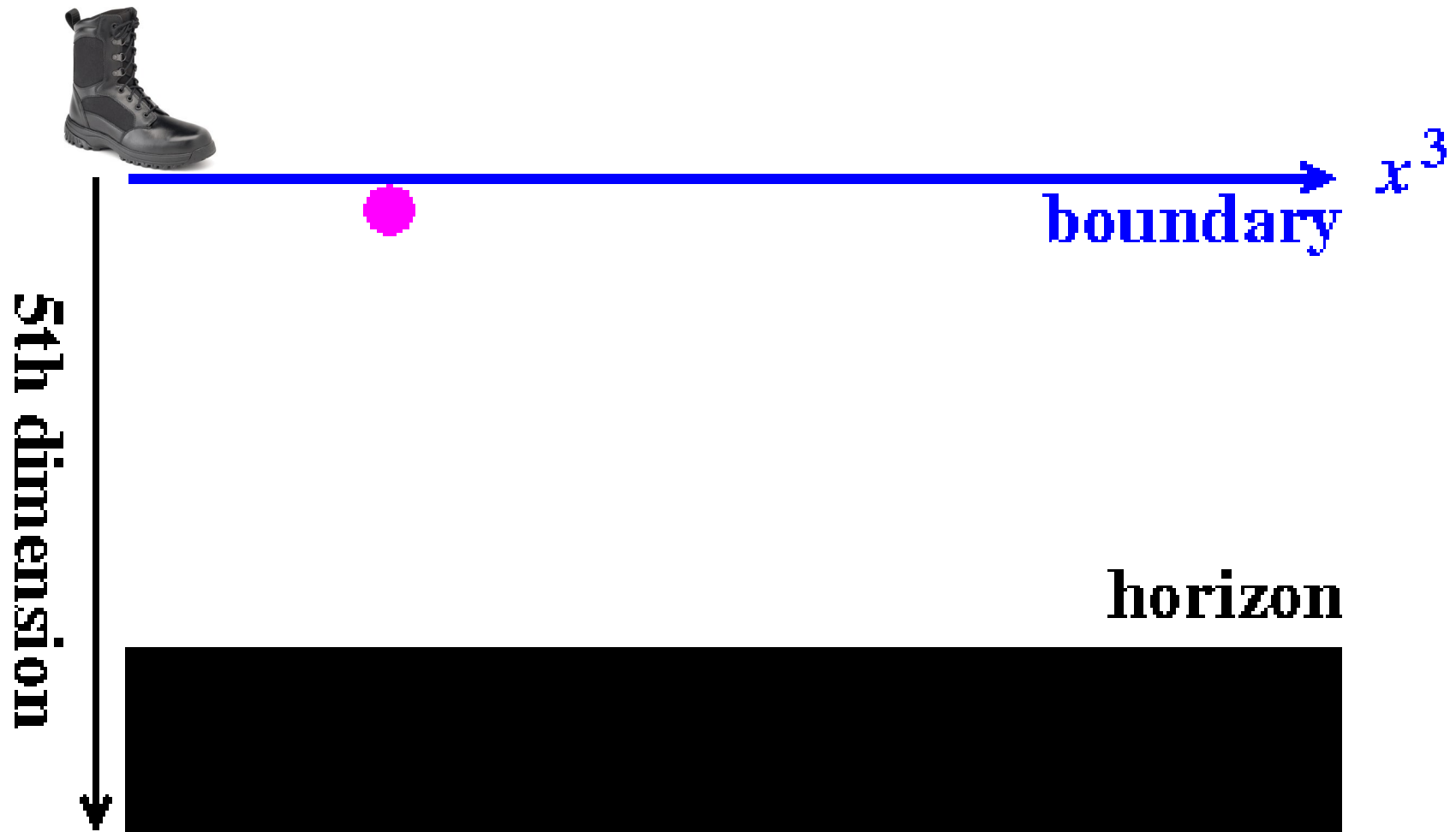
Stopping Distance using AdS/CFT

BIG $\alpha_s = \alpha_s$: Large- N_c $\mathcal{N}=4$ SYM, etc. with $N_c \alpha_s \rightarrow \infty$



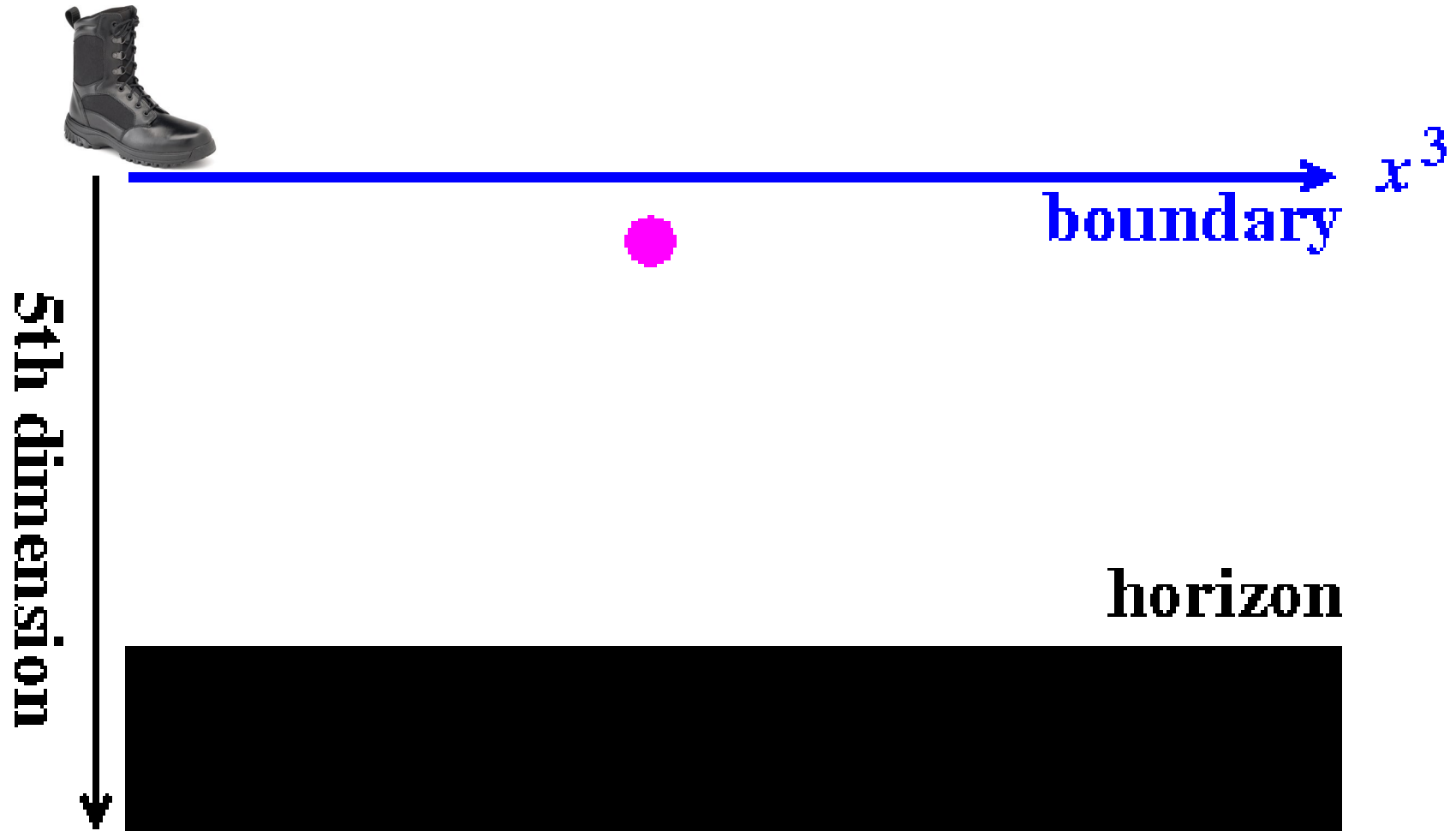
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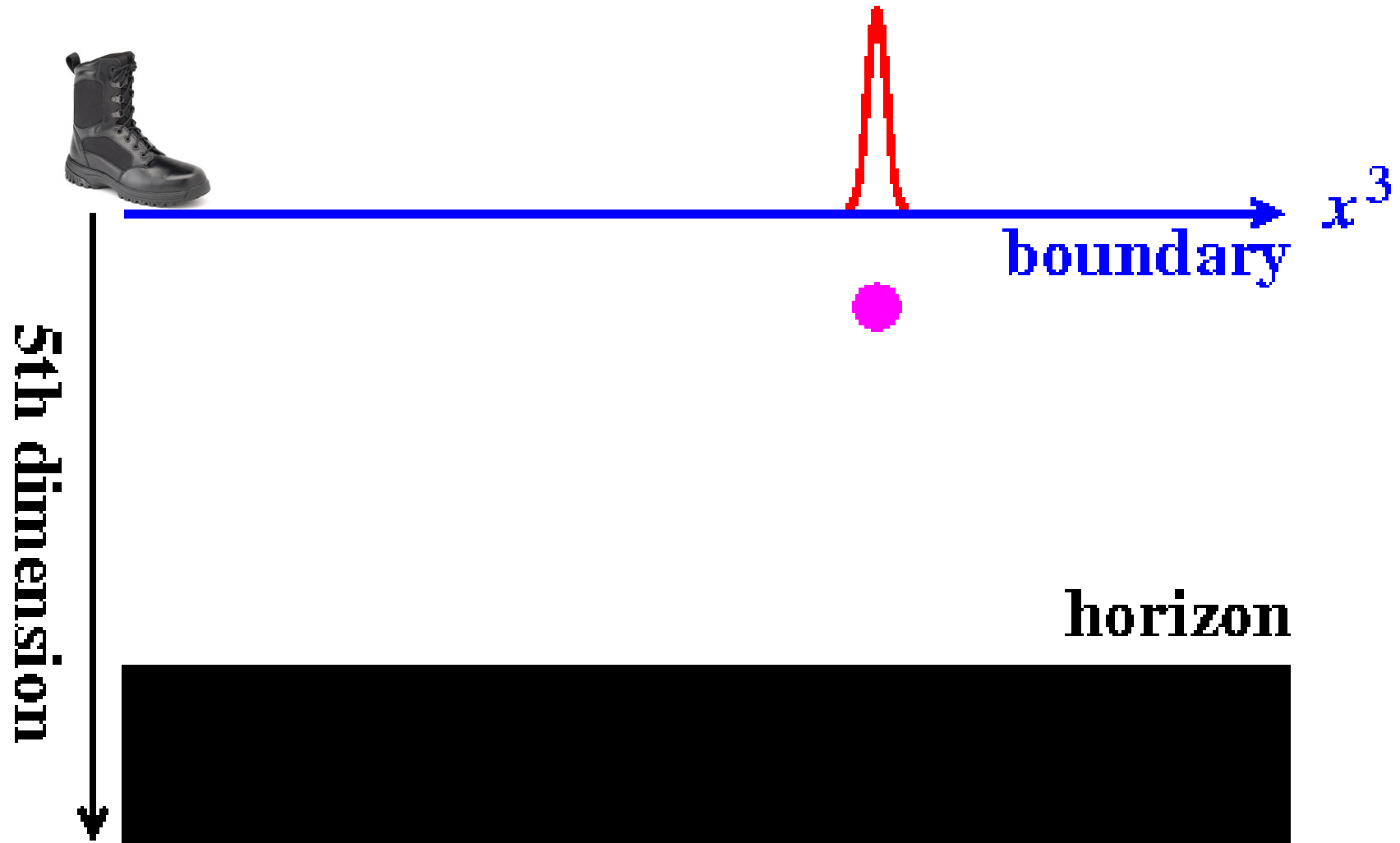
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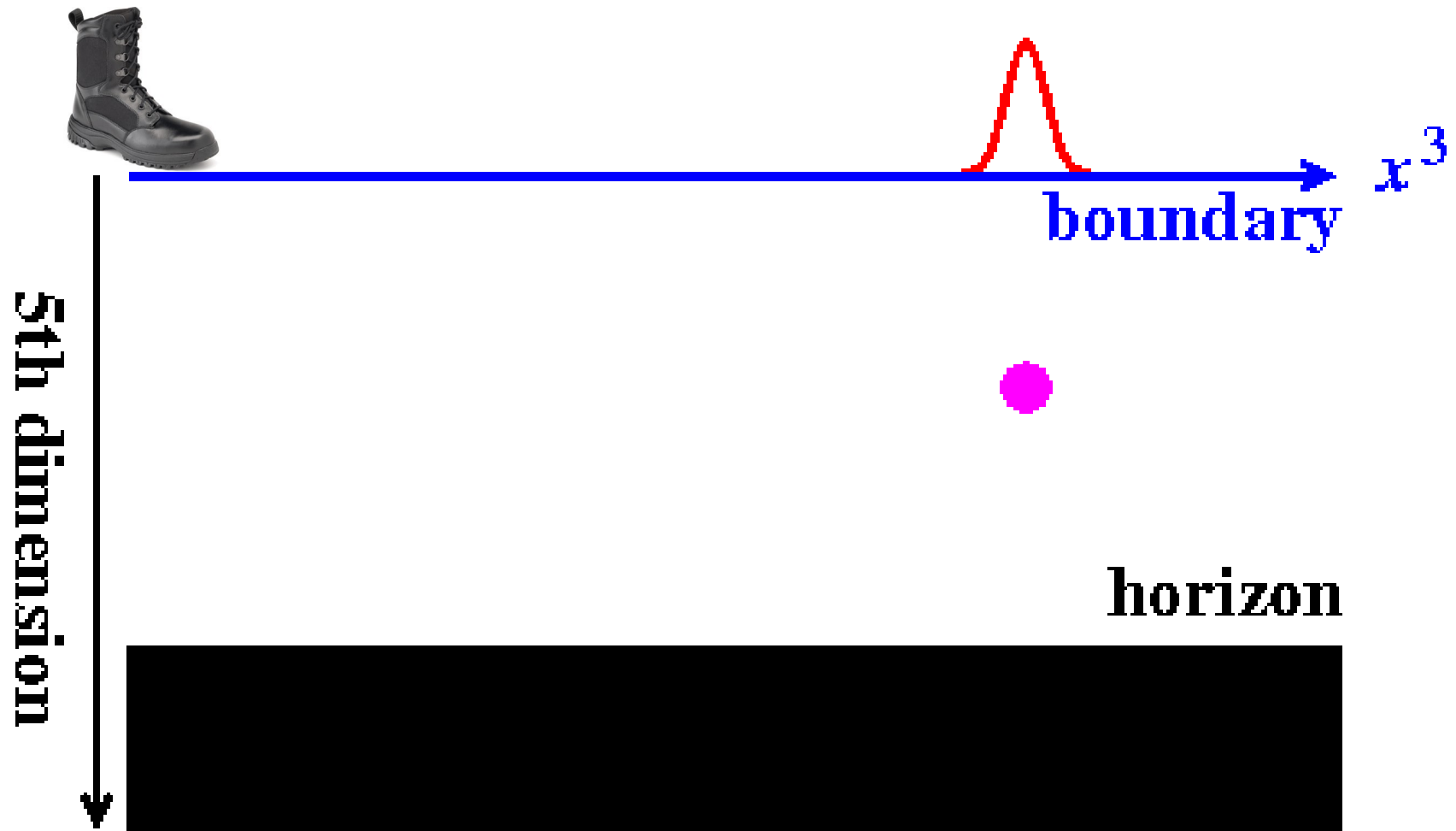
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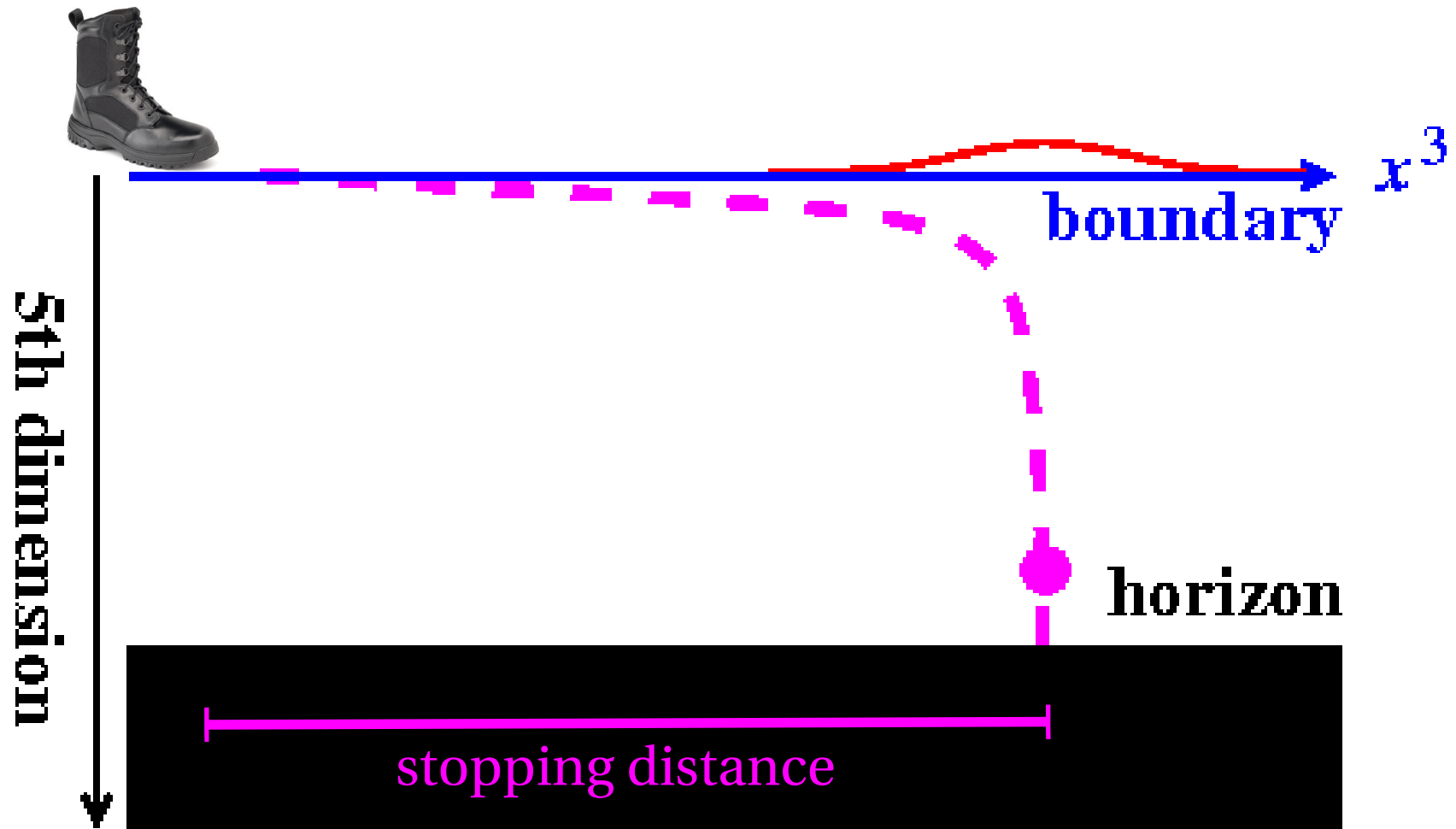
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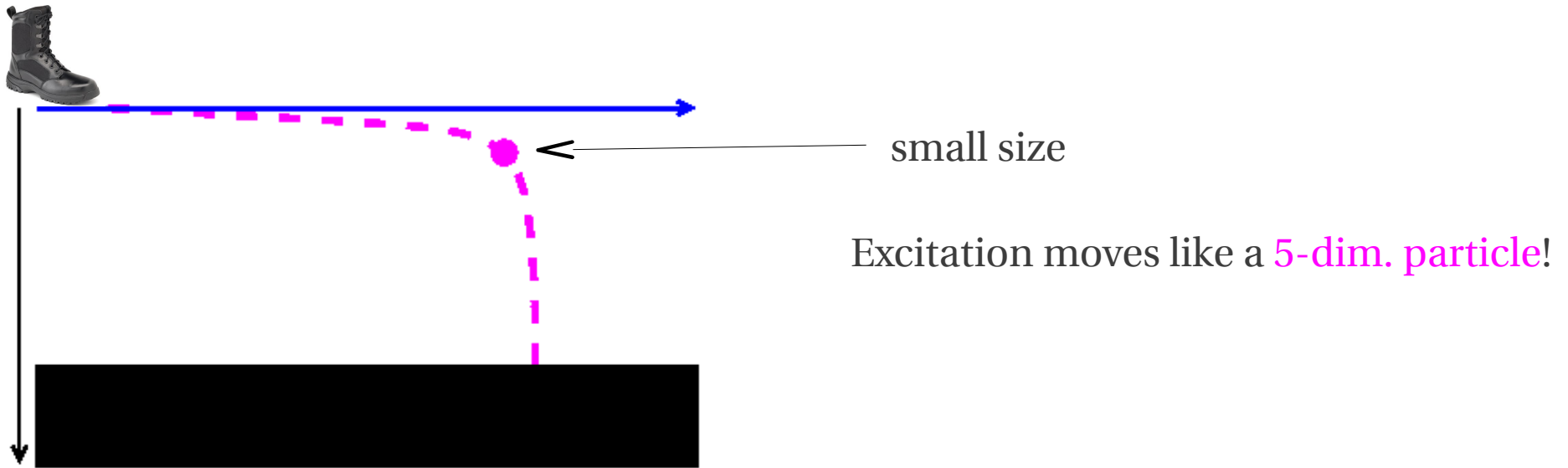
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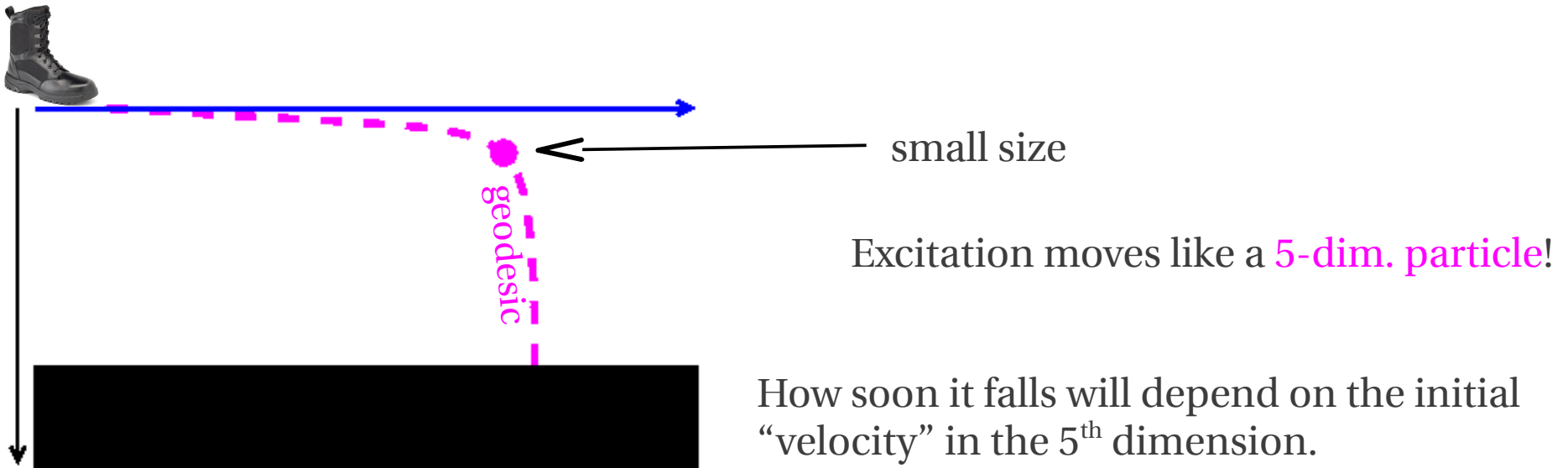
$\lambda = \infty$ result

$$l_{\text{stop}} \gtrsim l_{\text{max}} \propto E^{1/3}$$


A simplified picture for $l_{\text{stop}} \ll l_{\text{max}}$



A simplified picture for $l_{\text{stop}} \ll l_{\text{max}}$



Q: What determines l_{stop} ?

A: the 4-virtuality $q^2 \equiv q_\mu \eta^{\mu\nu} q_\nu$ of the source 

Why?

Consider massless 5-dim. particle near the boundary:

$$0 = q_\mu q^\mu + q_5 q^5$$

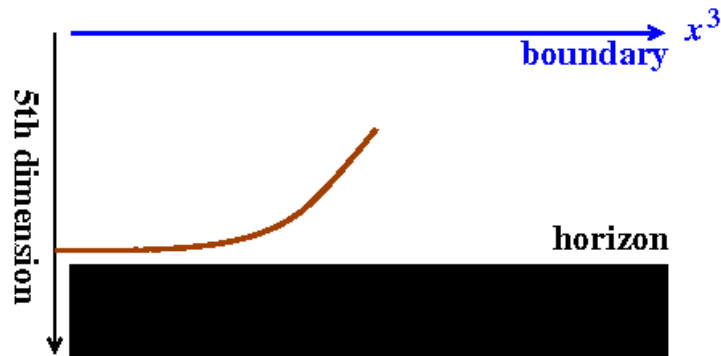
Bigger $-q_\mu q^\mu \rightarrow$ bigger $q^5 \rightarrow$ falls sooner !

The result:

$$l_{\text{stop}} \simeq \frac{\Gamma^2(\frac{1}{4})}{\sqrt{4\pi}} \left(\frac{E^2}{-q^2} \right)^{1/4}$$

A difference between different treatments

“Jets” represented as classical strings in dual theory

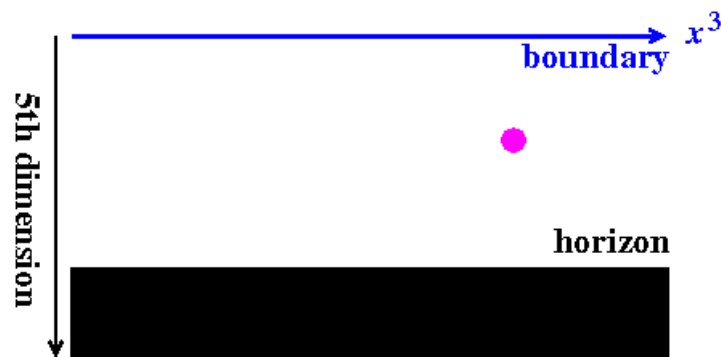
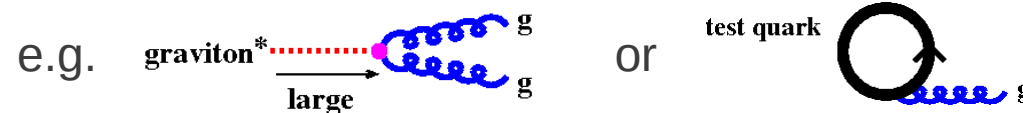


$$\ell_{\max} \sim T^{-4/3} (E / \sqrt{\lambda})^{1/3}$$

Gubser, Gulotta, Pufu, Rocha (2008)

Chesler, Jensen, Karch, Yaffe (2008)

“Jets” that are not



$$\ell_{\max} \sim T^{-4/3} E^{1/3}$$

Hatta, Iancu, Mueller (2008)

Arnold & Vaman (2010)

Chesler, Ho, Rajagopal (2011)

AdS/CFT Results

A. Results for $\lambda = \infty$ and $N_c = \infty$

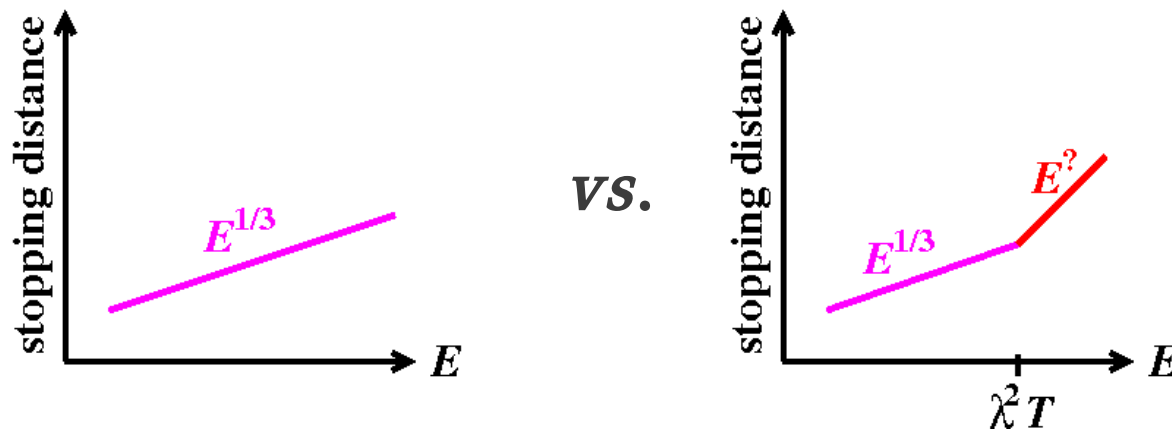
B. Results for $1 \ll \lambda < \infty$ and $N_c = \infty$

Previous “ $\alpha_s = \alpha_s$ BIG” result $\ell_{\text{stop}} \propto E^{1/3}$ has only been derived for $\lambda = N_c \alpha = \infty$.

What could we learn by also studying
 λ BIG but $< \infty$?

Answer: Is the high-energy behavior really $E^{1/3}$?

e.g.



(Note: will assume $N_c = \infty$ throughout.)

Our first look at this question

Arnold, Vaman, Szepietowski, 1203.6658

1. Assume $1/\lambda^n$ corrections are small.
2. Estimate their size.
3. See if anything goes wrong as $E \rightarrow \infty$!

Higher curvature corrections to gauge-gravity duality

AdS/CFT correspondence:

$\mathcal{N}=4$ SYM



string theory in $\text{AdS}_5 \times S^5$ background

Strong-coupling limit:

$\lambda \rightarrow \infty$



“low energy” string theory

= supergravity in $\text{AdS}_5 \times S^5$ background

(gravitons + other massless string modes)

$$\mathcal{L}_{\text{grav}} \sim R$$

Higher curvature corrections to gauge-gravity duality

AdS/CFT correspondence:

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string theory in $\text{AdS}_5 \times S^5$ background

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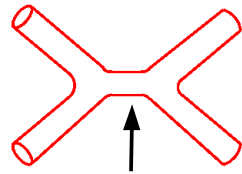
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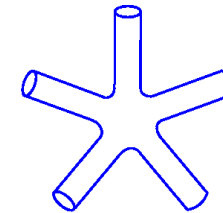
“low energy” string theory

= supergravity in $\text{AdS}_5 \times S^5$ background
(gravitons + other massless string modes)

$$\mathcal{L}_{\text{grav}} \sim R + \alpha'^3 R^4 + \alpha'^5 D^4 R^4 + \alpha'^6 D^6 R^4 + \dots + \alpha'^5 D^2 R^5 + \alpha'^6 D^4 R^5 + \dots + \dots$$



integrate out massive
intermediate string states



ditto

where

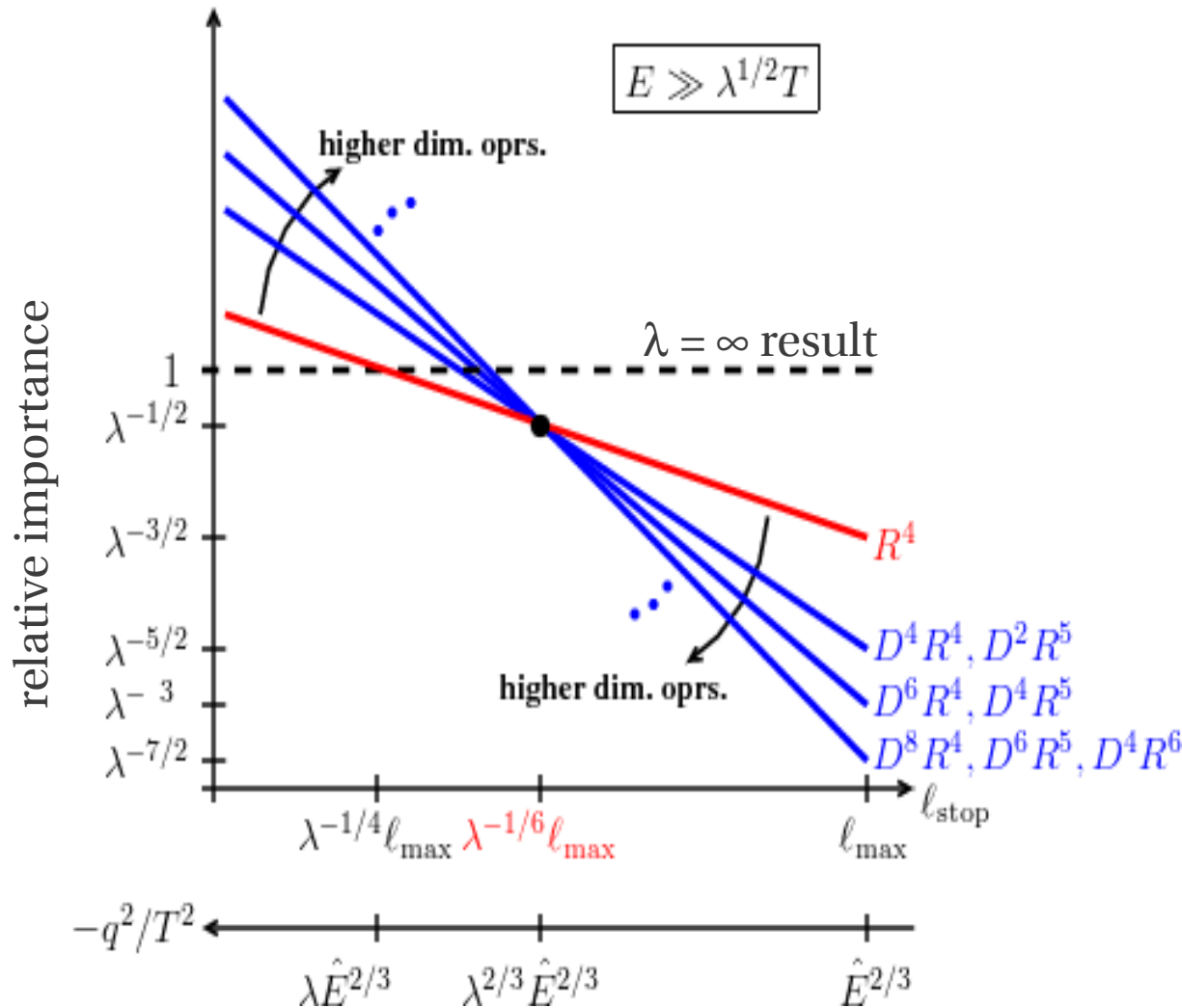
$$\frac{1}{\sqrt{\lambda}}$$



proportional to $\alpha' \sim \frac{1}{\text{string tension}}$

Note: Loops  are suppressed by $g_{\text{string}} \propto \frac{1}{N_c}$.

Importance to Jet Stopping



Reminder:

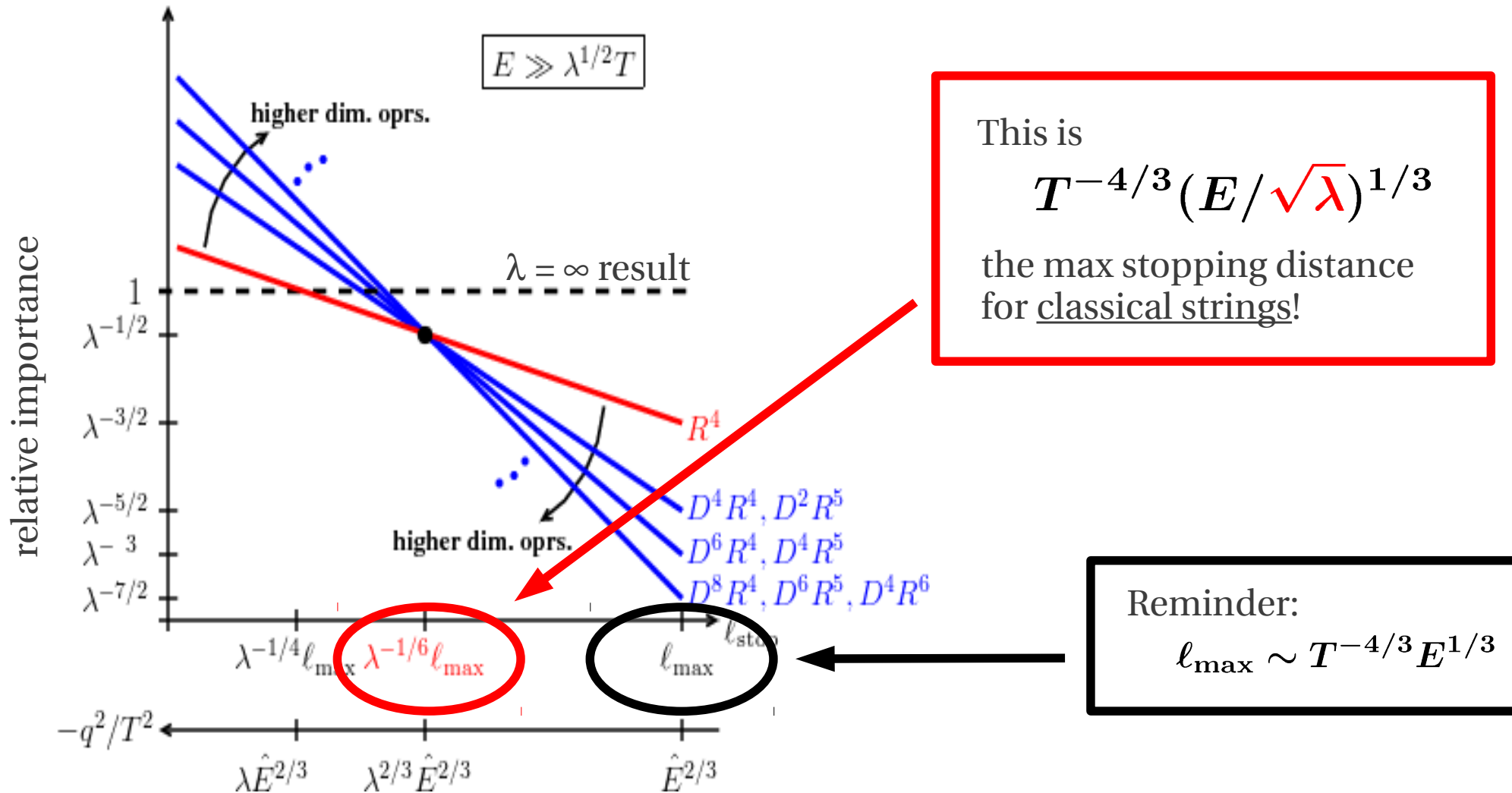
$$\ell_{\max} \propto E^{1/3}$$

Moral: Expansion in $1/\lambda$ is well-behaved for $\lambda^{-1/6} \ell_{\max} \ll \ell_{\text{stop}} \lesssim \ell_{\max}$

Expansion **breaks down** for $\ell_{\text{stop}} \lesssim \lambda^{-1/6} \ell_{\max}$

Note: Individual corrections all small ($\lambda^{-1/2}$) where expansion first breaks down.

Importance to Jet Stopping

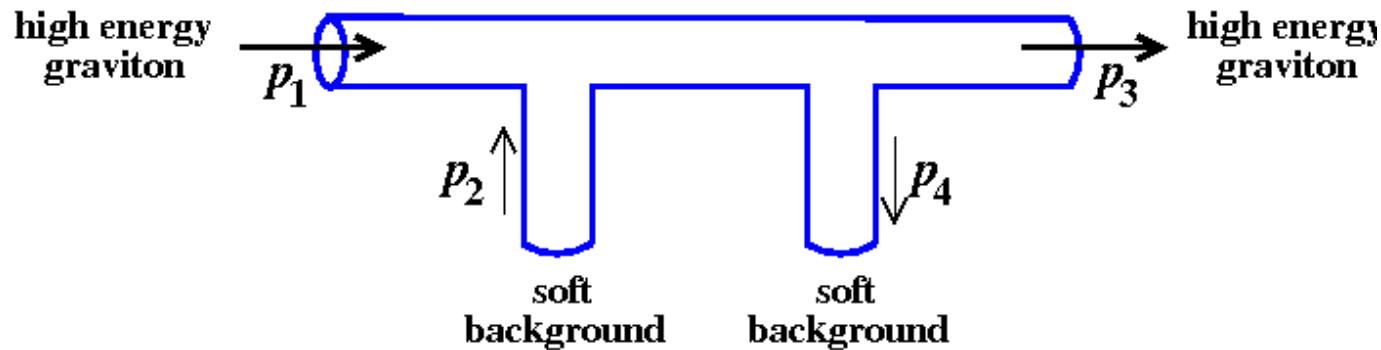


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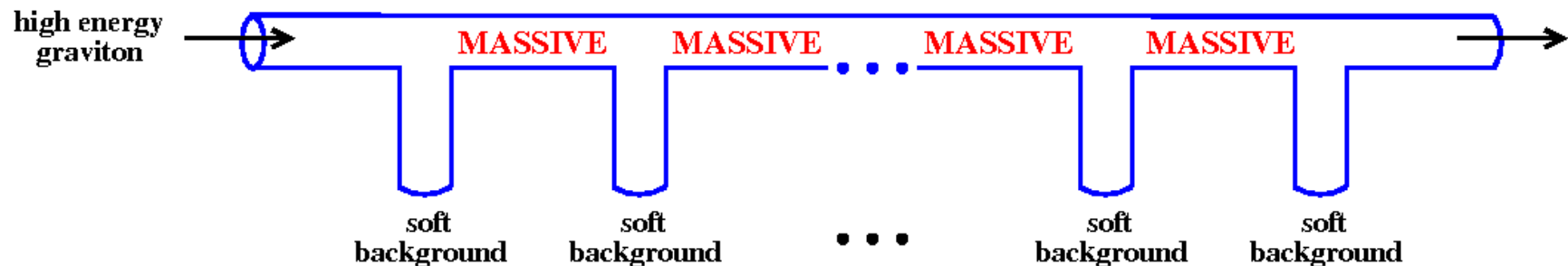
So what's gone wrong?



5-dimensional $\sqrt{s} \gtrsim M_{\text{stringy}}$

→ can excite real, massive, stringy internal states on-shell!

→ small p expansion breaks down, and also



is unsuppressed.

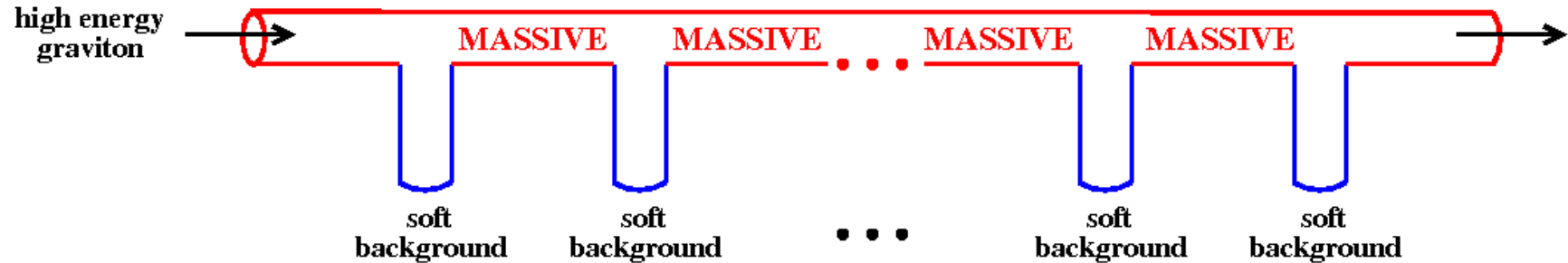
When might this possibly happen?

$$\sqrt{s} \propto \sqrt{E} \quad \text{and} \quad M_{\text{stringy}} \propto (\text{tension})^{1/2} \propto (\alpha')^{-1/2} \propto \lambda^{1/4} \quad \Rightarrow \quad E \gg \sqrt{\lambda} T$$

Details about l_{stop} condition comes from whether stringy modes excited before or after stopping distance reached.

What can we do?

Arnold, Vaman, Szepietowski, Wong, 1212.3321

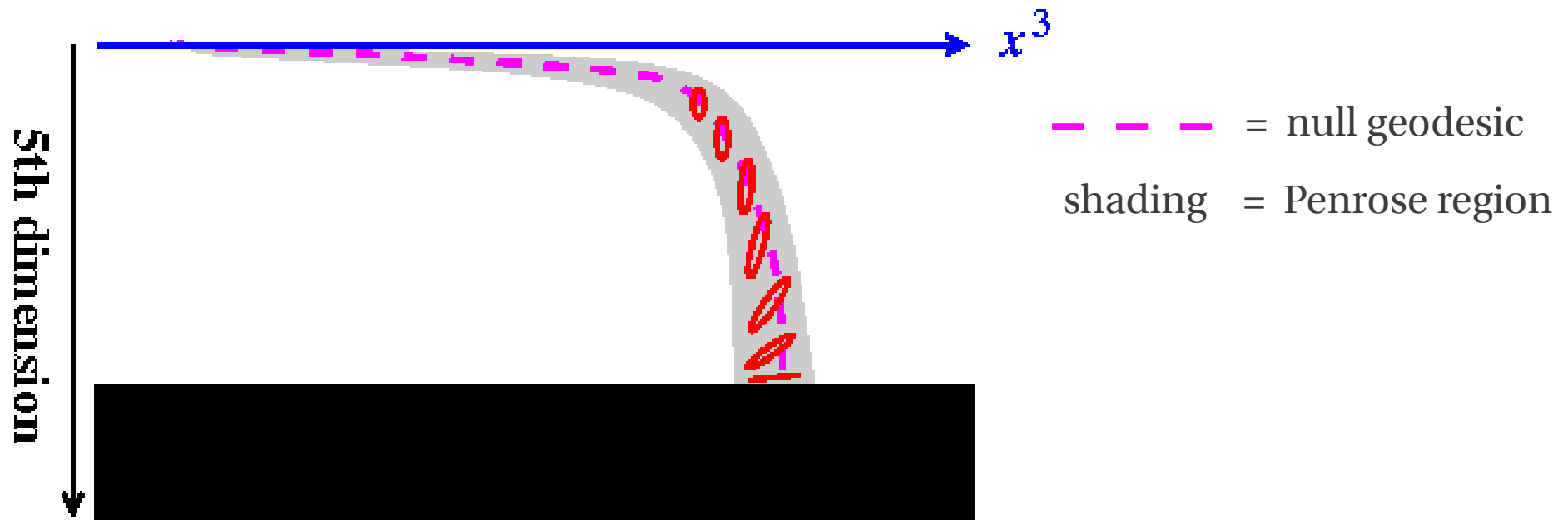


Answer: Replace the high-energy graviton by a closed **string** quantized in the AdS-Schwarzschild background.

Problem: No practical calculation method for quantizing strings in generic backgrounds.

Solution: For our problem, it's good enough to replace AdS-Schwarzschild by a **Penrose limit** (with caveats to be mentioned later).

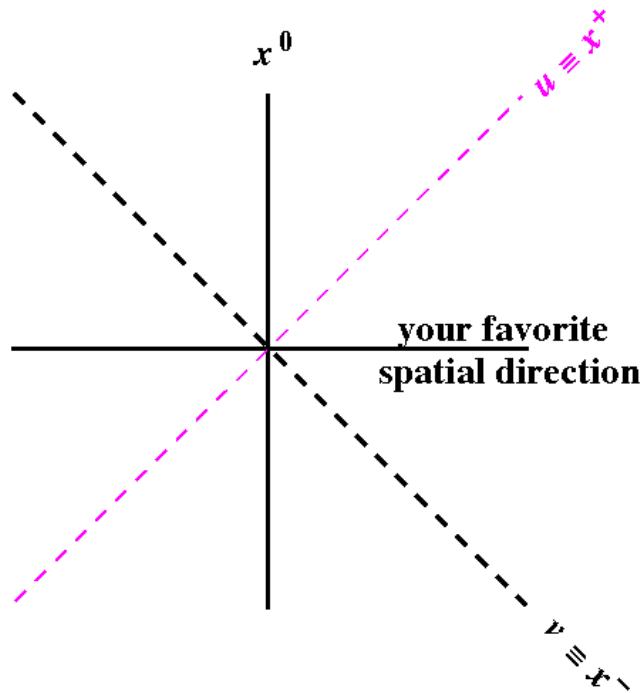
The Penrose Limit



In neighborhood of a null geodesic, can approximate the metric in a form (a pp-wave metric) for which practical calculations involving string quantization are possible!

How do we know if string will stay close enough to reference geodesic?
 Assume it does, calculate answer, and then check.

Reminder of string quantization in flat space



$$ds^2 = -du dv + d\vec{x}_\perp^2$$

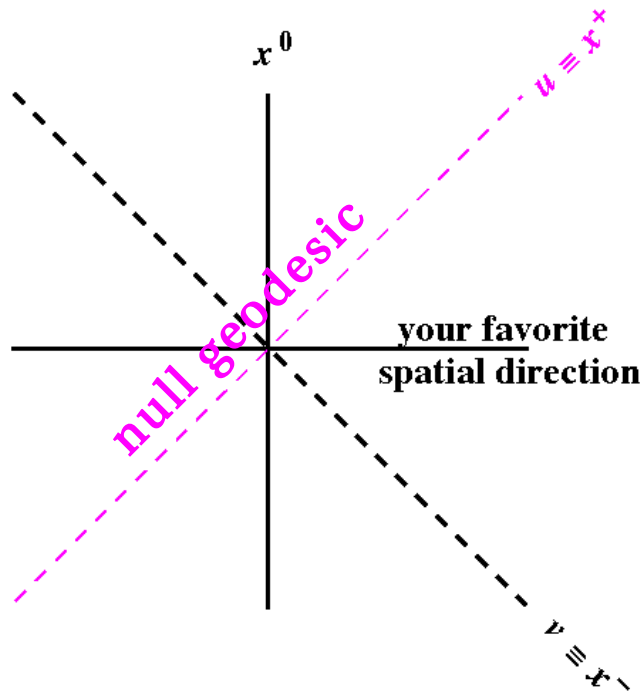
Take world sheet time to be $\tau = u$

→ string decomposes into independent, transverse, harmonic oscillators

$$X_\perp^k(\sigma, \tau) = \sum_n X_{\perp n}^k(\tau) e^{in\sigma}$$

$$X_{\perp n}^k(\tau) = \text{harm. osc. w/ frequency } \omega_{k,n} = \frac{n^2}{(\alpha' p^u)^2}$$

Reminder of string quantization in flat space



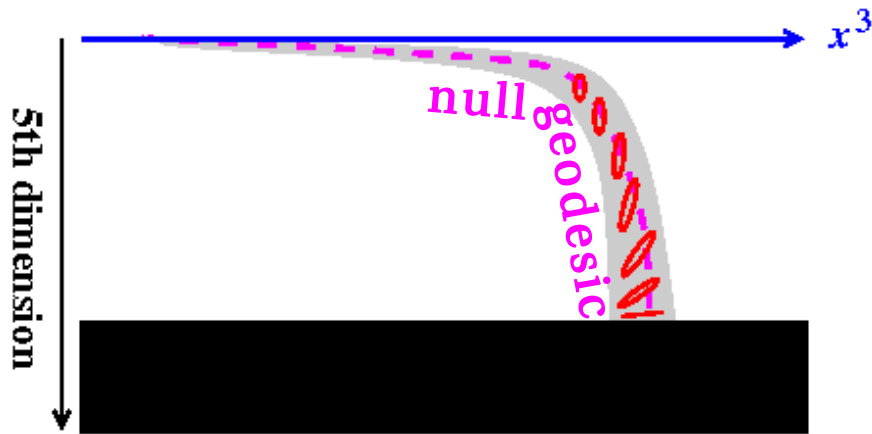
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Take u = affine parameter along reference null geodesic

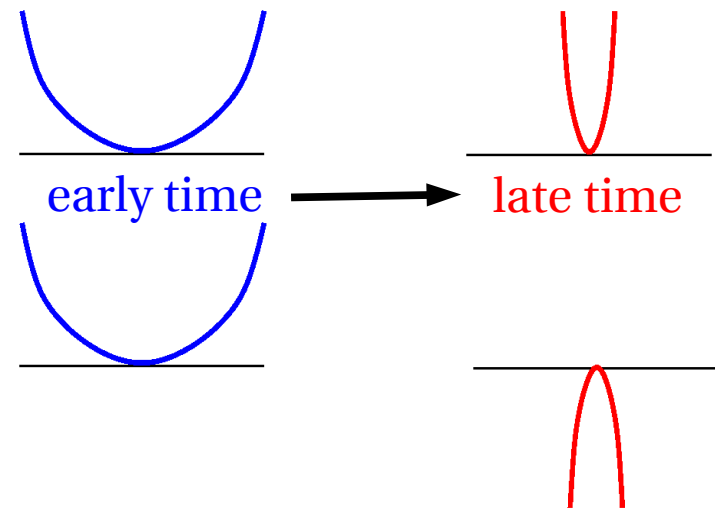
$$ds^2 = -du dv + \dots$$

Take world sheet time to be $\tau = u$

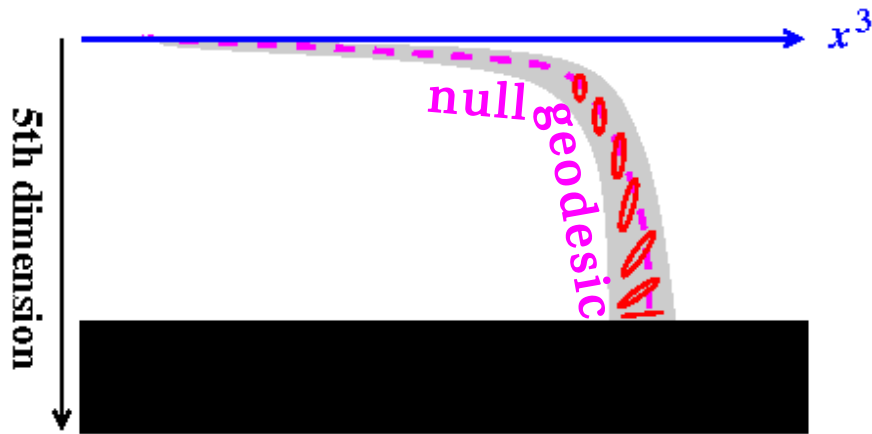
→ as before, but HO's pick up tidal force terms from curvature of space-time:

$$\omega_{1,n}^2 = \omega_{2,n}^2 = \frac{n^2}{(\alpha' p^u)^2} + \frac{1}{2} G(\tau)$$

$$\omega_{3,n}^2 = \frac{n^2}{(\alpha' p^u)^2} - G(\tau)$$



$G(\tau)$ grows as one moves away from the boundary.



Take u = affine parameter along reference null geodesic

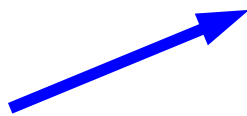
$$ds^2 = -du dv + \dots$$

Take world sheet time to be $\tau = u$

→ as before, but HO's pick up tidal force terms from curvature of space-time:

$$\omega_{1,n}^2 = \omega_{2,n}^2 = \frac{n^2}{(\alpha' p^u)^2} + \frac{1}{2} G(\tau)$$

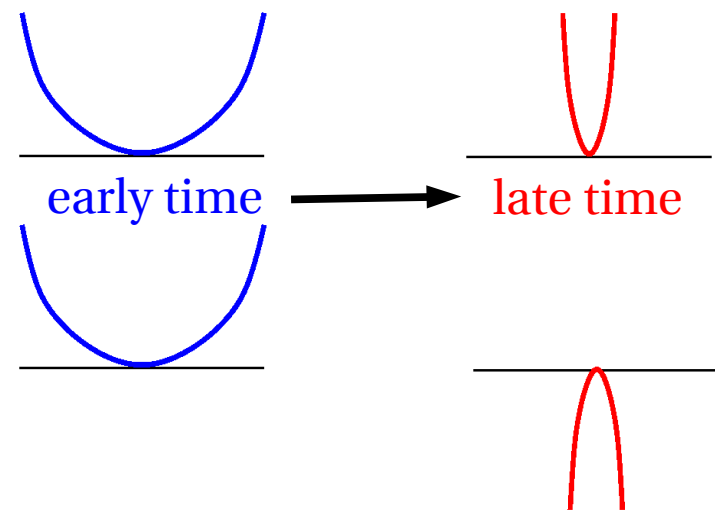
$$\omega_{3,n}^2 = \frac{n^2}{(\alpha' p^u)^2} - G(\tau)$$



STRING TENSION



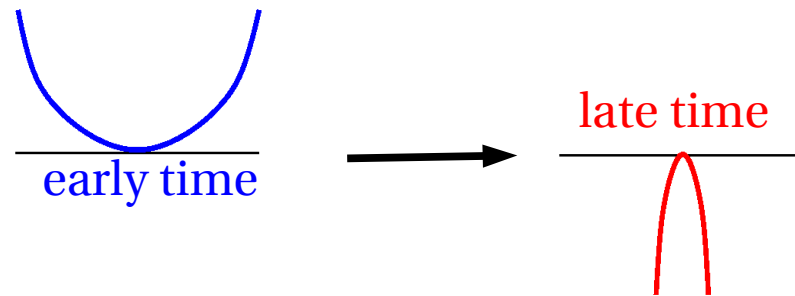
TIDAL FORCES



$G(\tau)$ grows as one moves away from the boundary.

Now just a QM problem:

Need QM solution to a time-dependent harmonic oscillator that starts in its ground state.

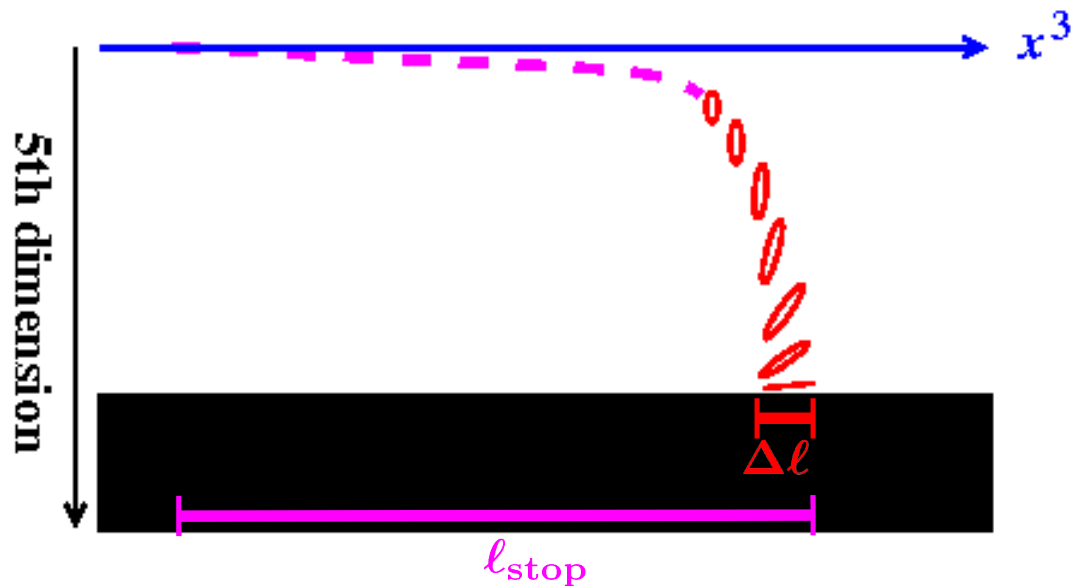


At late times, dynamics become **classical**.

Can calculate late-time probability distribution for each oscillator (i.e. for the amplitude X_n^k of each string harmonic).

→ Can calculate the late-time size of the classical string.

Results



For case $l_{\text{stop}} \lesssim \lambda^{-1/6} l_{\text{max}}$ where string excitation is important,

$$(\Delta l)_{\text{rms}} \simeq 0.8660 \lambda^{-1/4} l_{\text{stop}} \ln^{1/2} n_*$$

where

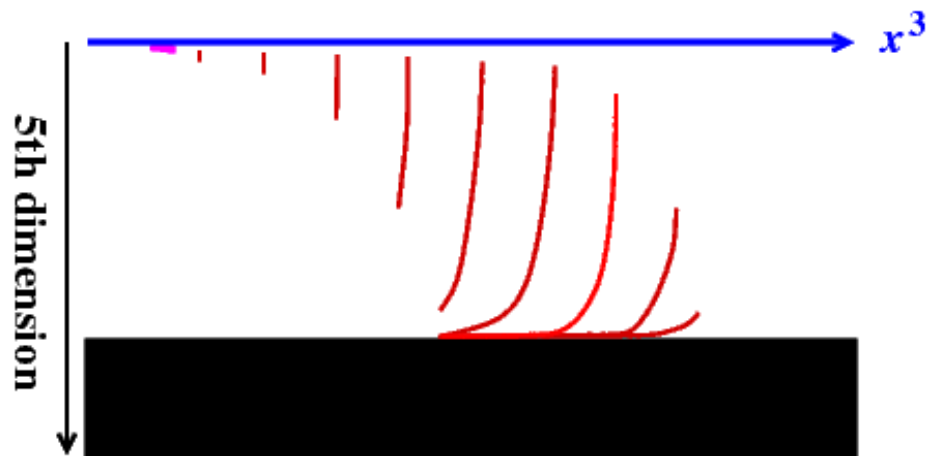
$$n_* \sim \# \text{ harmonics excited} \sim \frac{\lambda^{-1/6} l_{\text{max}}}{l_{\text{stop}}}$$

Moral: Stretching of string has negligible impact on jet stopping unless n_* is exponentially large!

And what if it is? ...

Large $\ln(n_*)$

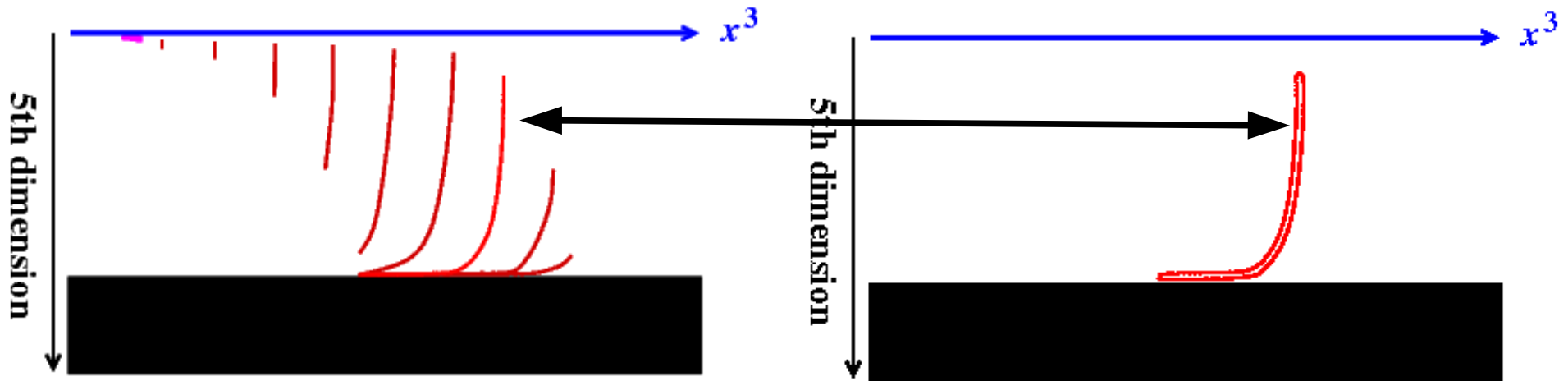
$$\lambda^{-1/4} \ln^{1/2} n_* \sim 1$$



Note: Penrose limit breaks down

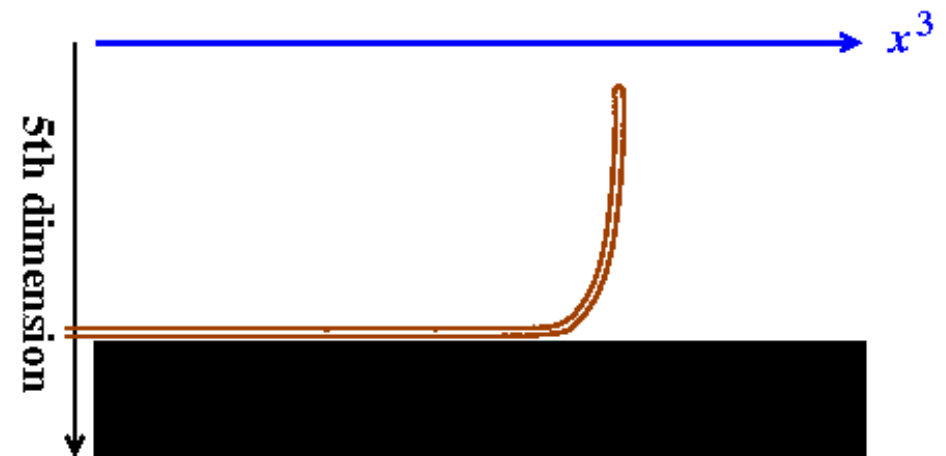
Large $\ln(n_*)$

$$\lambda^{-1/4} \ln^{1/2} n_* \sim 1$$



Note: Penrose limit breaks down

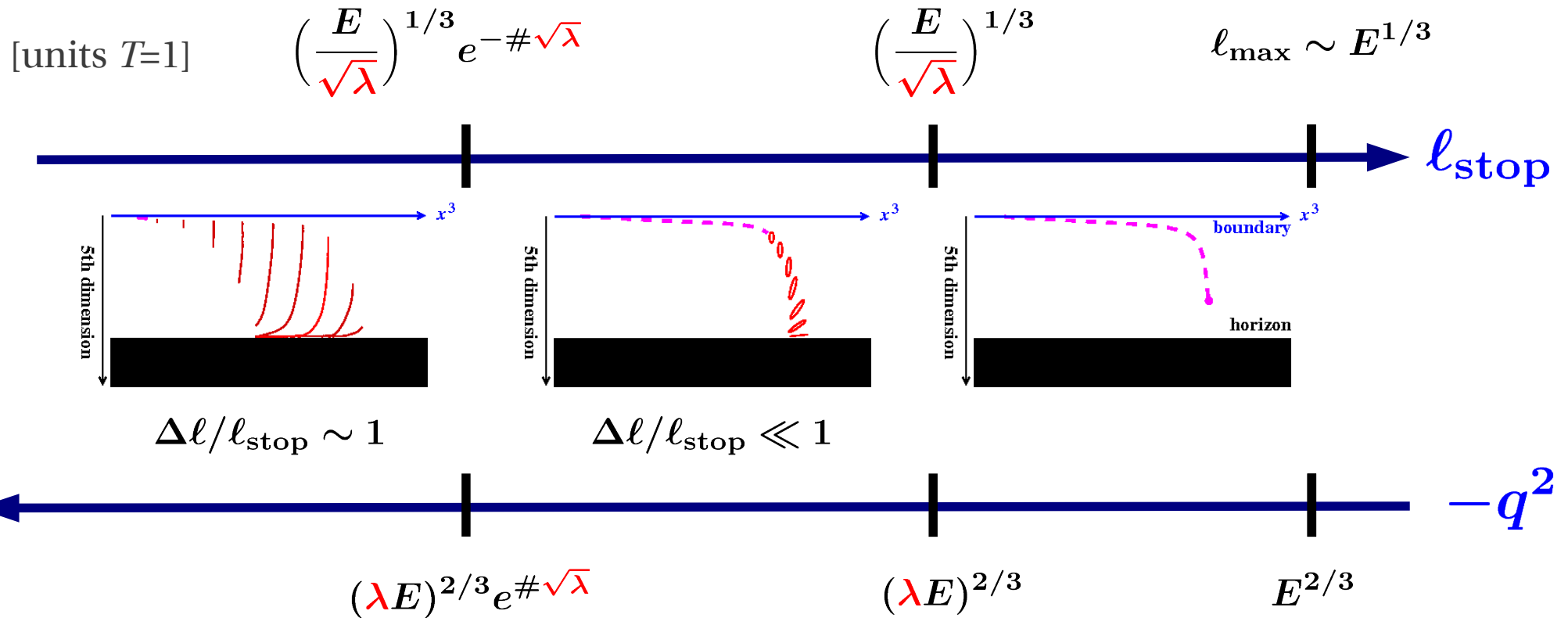
qualitatively similar to



Gubser, Gulotta, Pufu, Rocha (2008)

Summary

At very high energy $E \gg \sqrt{\lambda} e^{\#\sqrt{\lambda}} T$



What happens as λ decreases?

All the above scales coalesce as $\lambda \rightarrow 1$.